

GENERAL ATOMIC
DIVISION OF
GENERAL DYNAMICS

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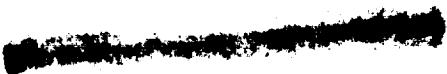
NUCLEAR PULSE SPACE VEHICLE STUDY

**Vol. IV--MISSION VELOCITY REQUIREMENTS
AND SYSTEM COMPARISONS
(SUPPLEMENT)**

George C. Marshall Space Flight Center
Future Projects Office
National Aeronautics and Space Administration
Huntsville, Alabama

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FOREWORD

This report is a supplement to Vol. IV, Mission Velocity Requirements and System Comparisons (Unclassified) of the report on the Nuclear Pulse Vehicle Study performed under National Aeronautics and Space Administration Contract NAS8-11053. This supplement, referred to as Vol. IVA contains mission velocity and mass ratio charts which are background material for some of the missions studied under the above mentioned contract and which in part also represent relevant results of in-house studies.

The report is divided into six sections, one for each planet from Mercury to Saturn, including Earth. For Mercury, Venus and Jupiter, both, hyperbolic departure and arrival data, as well as planetocentric data are presented. For Saturn, only hyperbolic data, and for Earth and Mars only planetocentric data are presented. The hyperbolic departure and arrival data refer to mono-elliptic heliocentric transfer paths between Earth and target planet, and vice versa. In the case of Venus they are presented as impulsive velocity changes Δv (10^3 ft/sec) at 1.1 planet radii distance. By conversion to EMOS (97,700 ft/sec = 1 EMOS), the associated hyperbolic excess velocity v_∞^* (both in EMOS) can be read from $\Delta v^*/v_\infty^*$ (both in EMOS) versus v_∞^* charts for the respective capture conditions (1.1 planet radii, circular capture orbit). For Mercury, Jupiter and Saturn, the hyperbolic excess velocities are shown (in EMOS). The associated impulsive capture or departure maneuver is readily found from $\Delta v^*/v_\infty^*$ charts for a large range of planetocentric distances and capture orbit ellipticities.

The planetocentric data for Mercury, Venus and Jupiter refer to the (impulsive) maneuvers required for transition into circular orbit at the apapsis of elliptic capture orbits. Those for Earth and Mars refer to powered flight time and mass ratio versus hyperbolic excess velocity ranging from $F/W_0 = 0.001$ to $0.1 g$ at various specific impulses ranging from 1000 to 10,000 seconds. For Earth, four out of the five charts presented refer to departure operations, one ($0.05g$) to arrival at Earth. In comparing the chart for departure at $0.05 g$ with that for arrival at $0.05 g$, it is seen that, for equal values of v_∞^* and I_{sp} , the mass ratio for the arrival maneuver is slightly higher than for the departure maneuver. The reason for this is that, for departure maneuvers, the thrust acceleration (thrust/weight ratio F/W_0) refers to the initial value, whereas, for the arrival maneuver, it designates the terminal value. The same definition applies to the Mars charts all of which pertain to arrival maneuvers. Therefore, the designation on the arrival maneuver charts should say F/W_1 , rather than F/W_0 . Initial and terminal conditions are circular orbit at 1.1 Earth radii and 1.3 Mars radii. Since, at Mars the gravitational losses are lower, the differences between departure and arrival maneuvers

of finite burning time (i. e. non-impulsive) is even smaller than for Earth. Therefore, the arrival maneuver charts for Mars are representative (though slightly conservative) also for Mars departure. For Venus, the values are sufficiently close to those for Earth to permit use of the Earth charts for the accuracy level required in this study.

It should be emphasized, that the thrust accelerations specified on top of each chart are in Earth surface g-values, not in local g-values. In other words, the thrust acceleration is given in units of 32.17 ft/sec² or 9.81 m/sec². The thrust direction is taken as tangential to the instantaneous velocity vector throughout the powered flight period. This is not the optimum, but is sufficiently close to it, to be representative for the actual powered flight path, if the latter would be optimized by varying the thrust vector between azimuthal and local tangential direction.

Only a few hyperbolic data are given for Venus and none for Mars, because values for missions to these planets are abundantly available¹⁾. In the chart for Earth departure to Venus in 1980, the designation $r^* = 1.3$ on the upper left hand side should be changed to $r^* = 1.1$.

1) cf., for instance, National Aeronautics and Space Administration, Space Flight Handbooks, Vol. 3, Planetary Flight Handbook, Part 1, 2 & 3, NASA, SP-35, Washington, D. C., 1963

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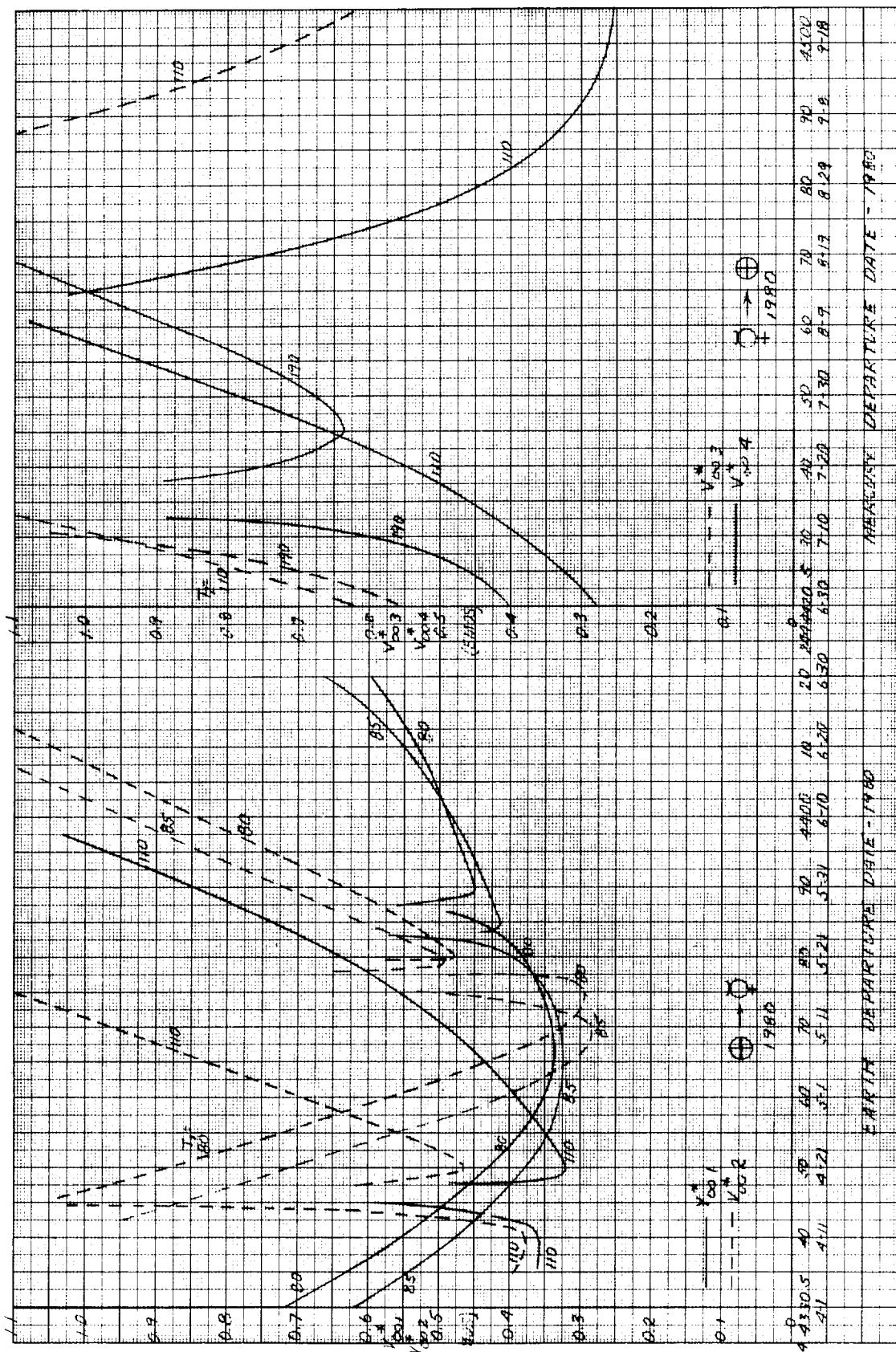
NOMENCLATURE AND DEFINITIONS

| | |
|--------------|--|
| I_{sp} | = Specific impulse (sec) |
| n | = Ratio of apoapsis to periapsis distance of elliptic capture orbit |
| r^* | = Distance (planet radii) |
| r_P^* | = Periapsis distance (planet radii) |
| T_1 | = Transfer period Earth to target planet (days) |
| T_2 | = Transfer period target planet to Earth (days) |
| t_1 | = Time of powered flight |
| v_E^* | = Earth atmospheric entry velocity following unretarded hyperbolic approach (nominally taken as the vacuum velocity along the approach hyperbola at 100 km apogee altitude) (EMOS) |
| W | = Weight (Earth pounds) |
| y | = Altitude (km) |
| Δv | = Impulsive velocity change (10^3 ft/sec) |
| Δv^* | = Impulsive velocity change (EMOS) |
| Δv_A | = Apoapsis impulse for change from elliptic to circular orbit (ft/sec) |
| | = Mass ratio (ignition to engine cut-off mass) |
| v_∞^* | = Hyperbolic excess velocity (EMOS) |

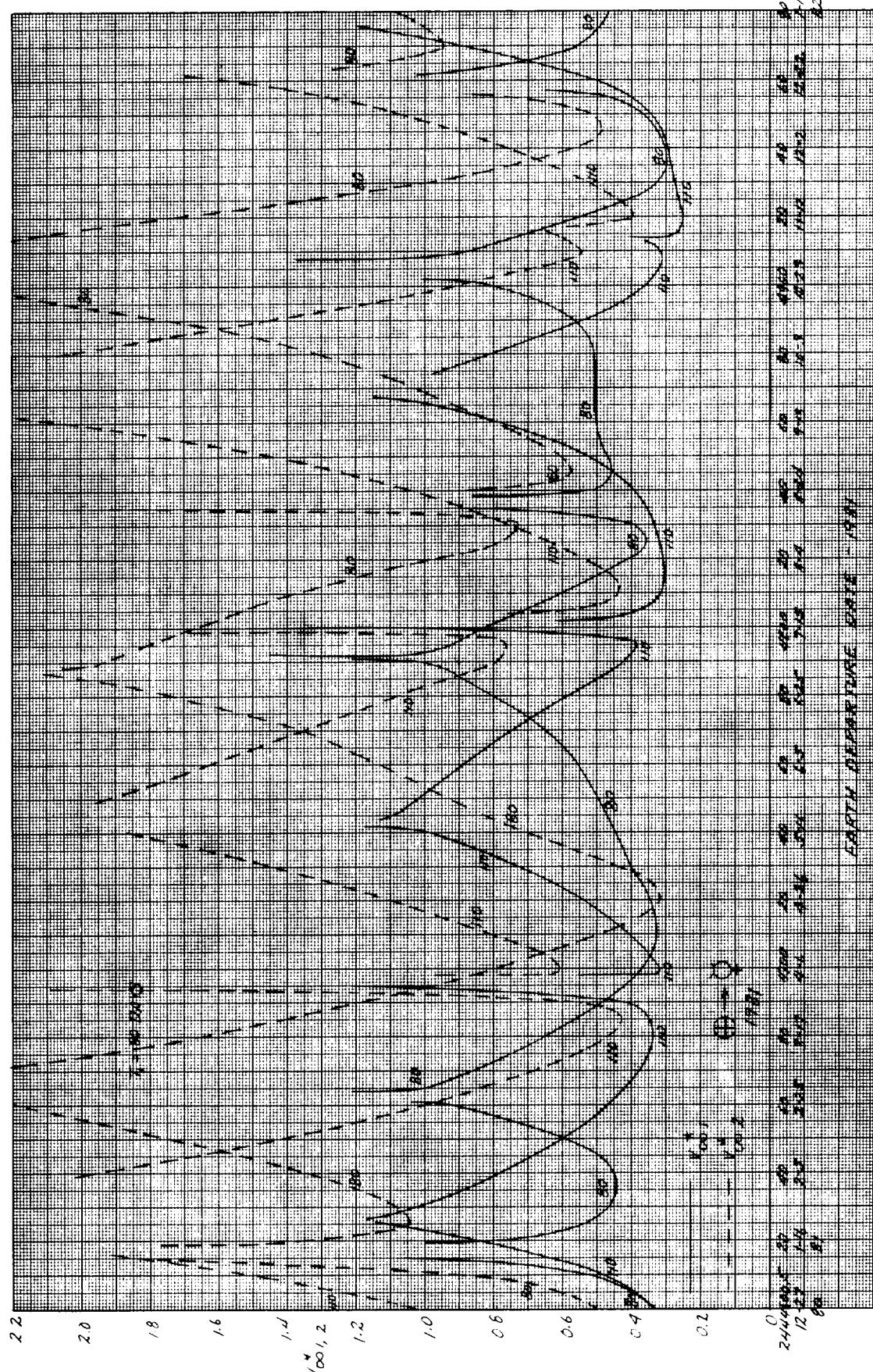
Subscripts:

- 1 Earth departure
- 2 Target planet arrival
- 3 Target planet departure
- 4 Earth arrival

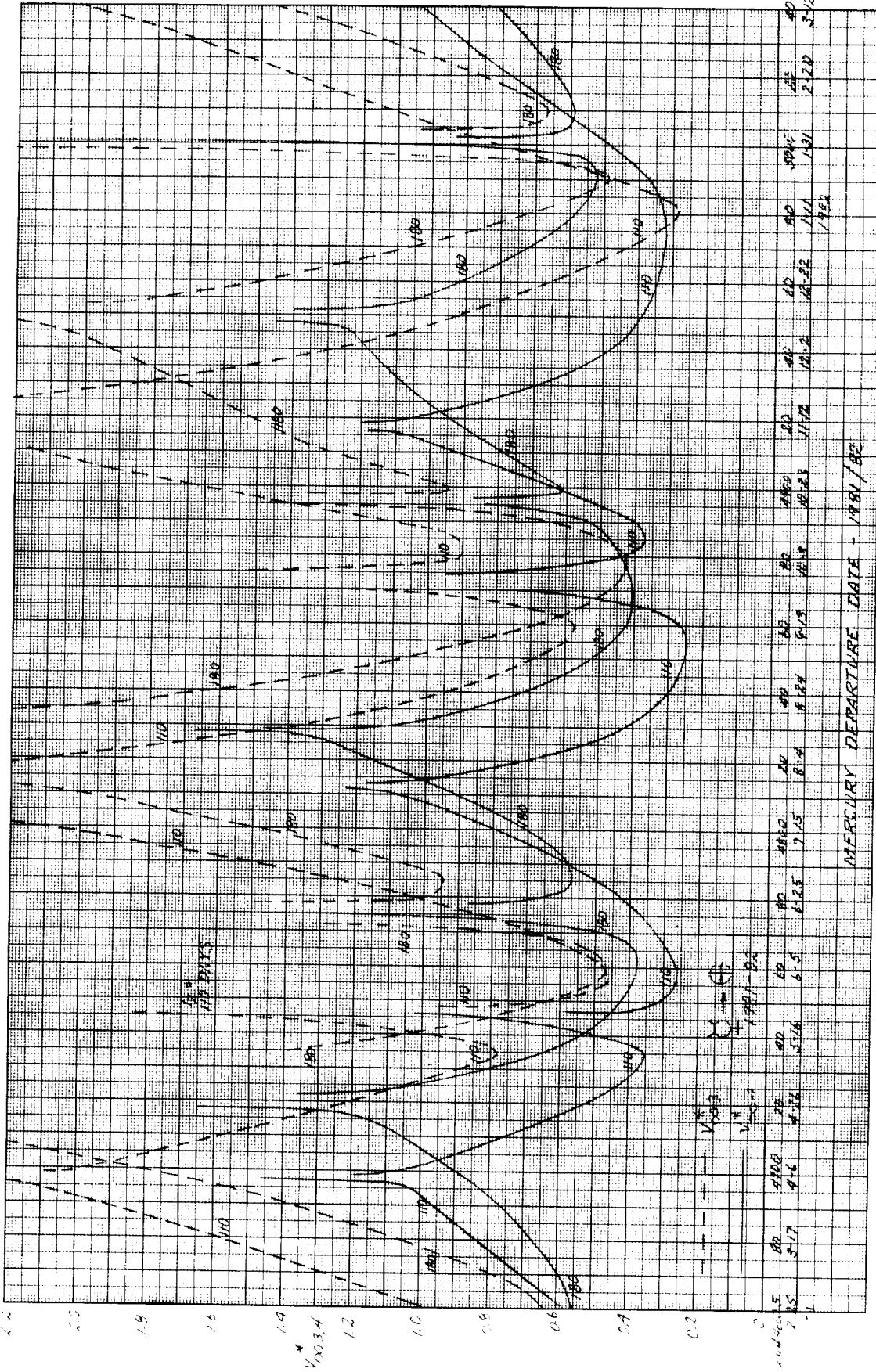
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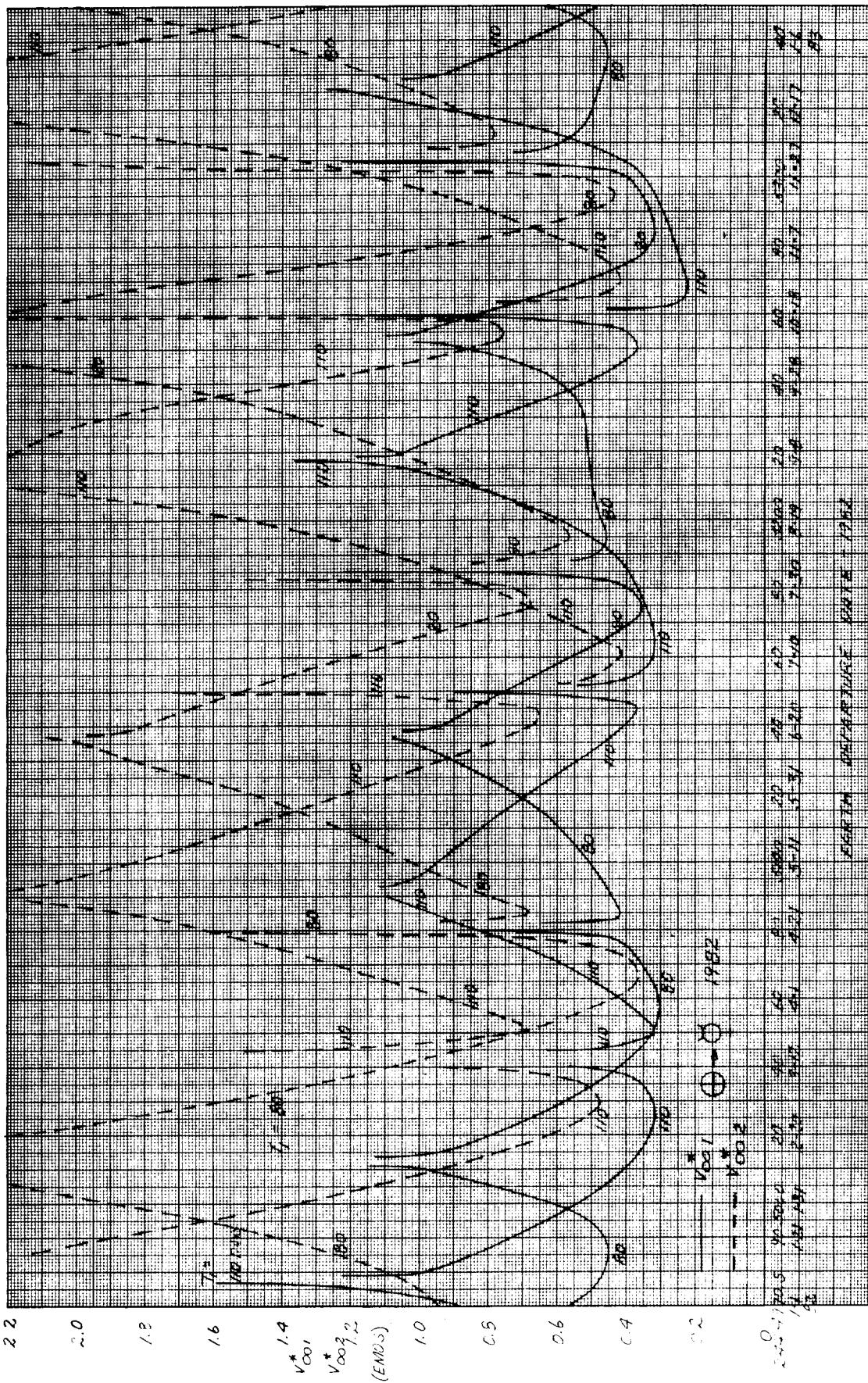
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64 - 57



64 - 150



22

0

10

16

1.4

 $v_{\infty 1}^*$

2

0.6

0.4

0.2

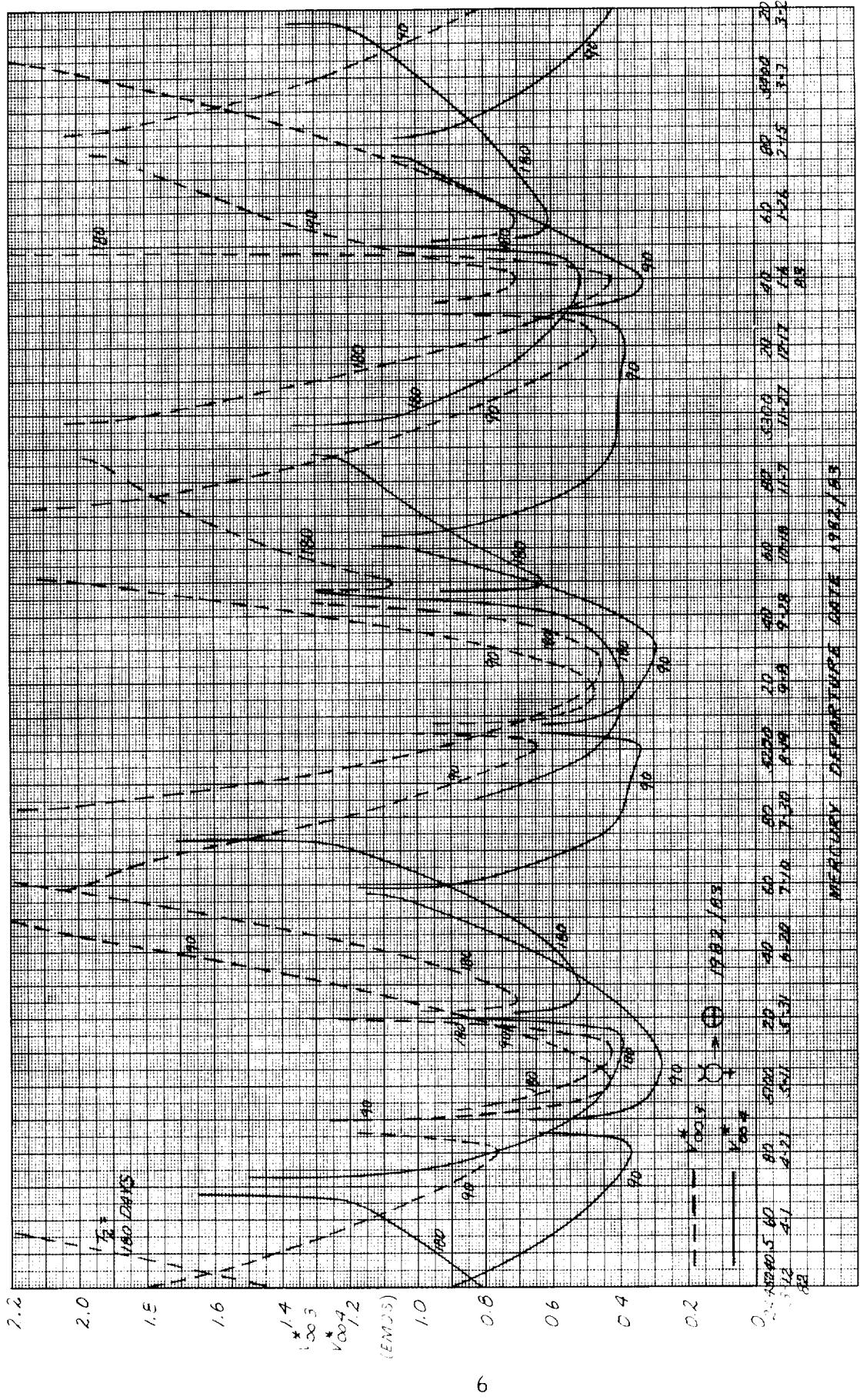
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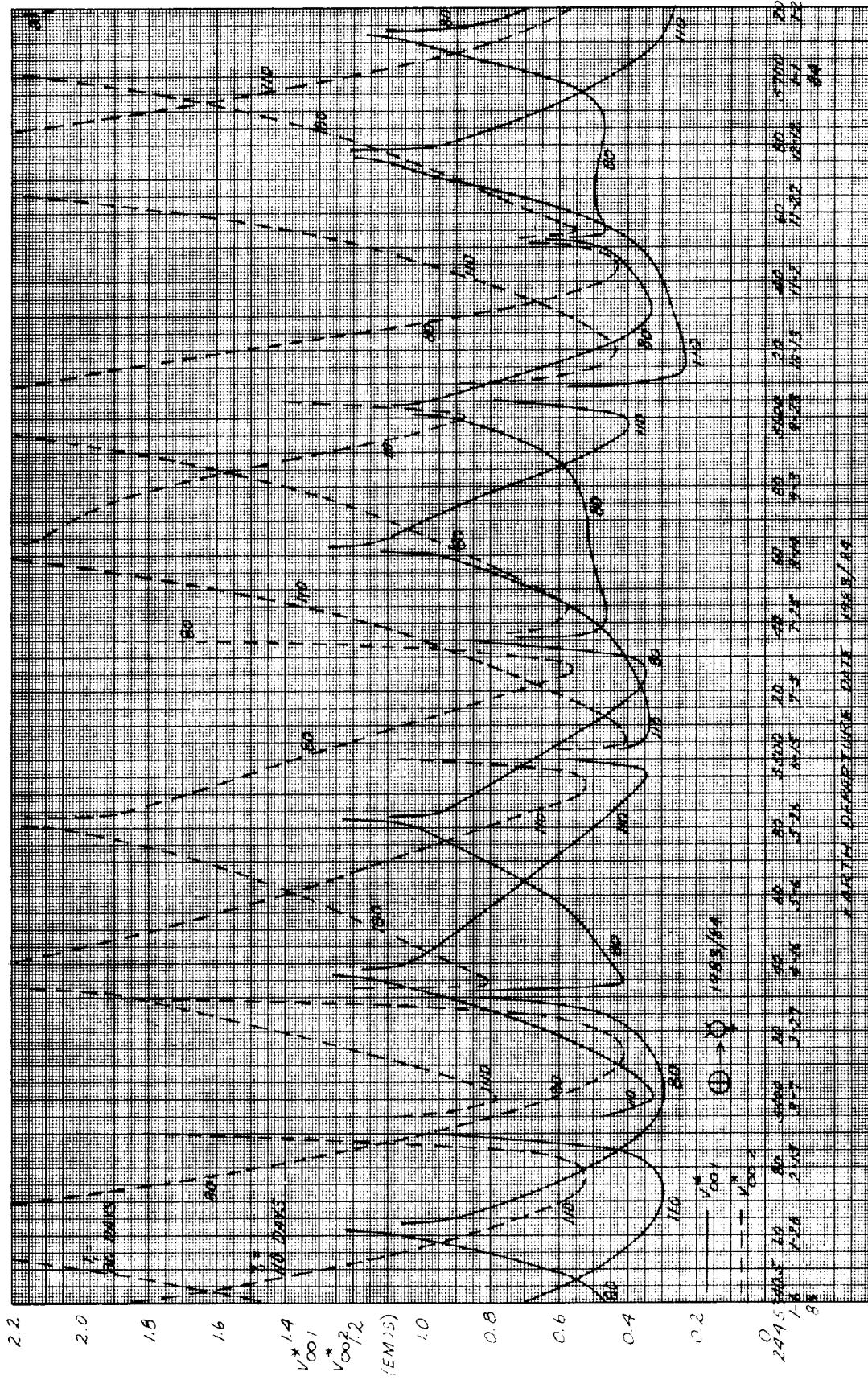
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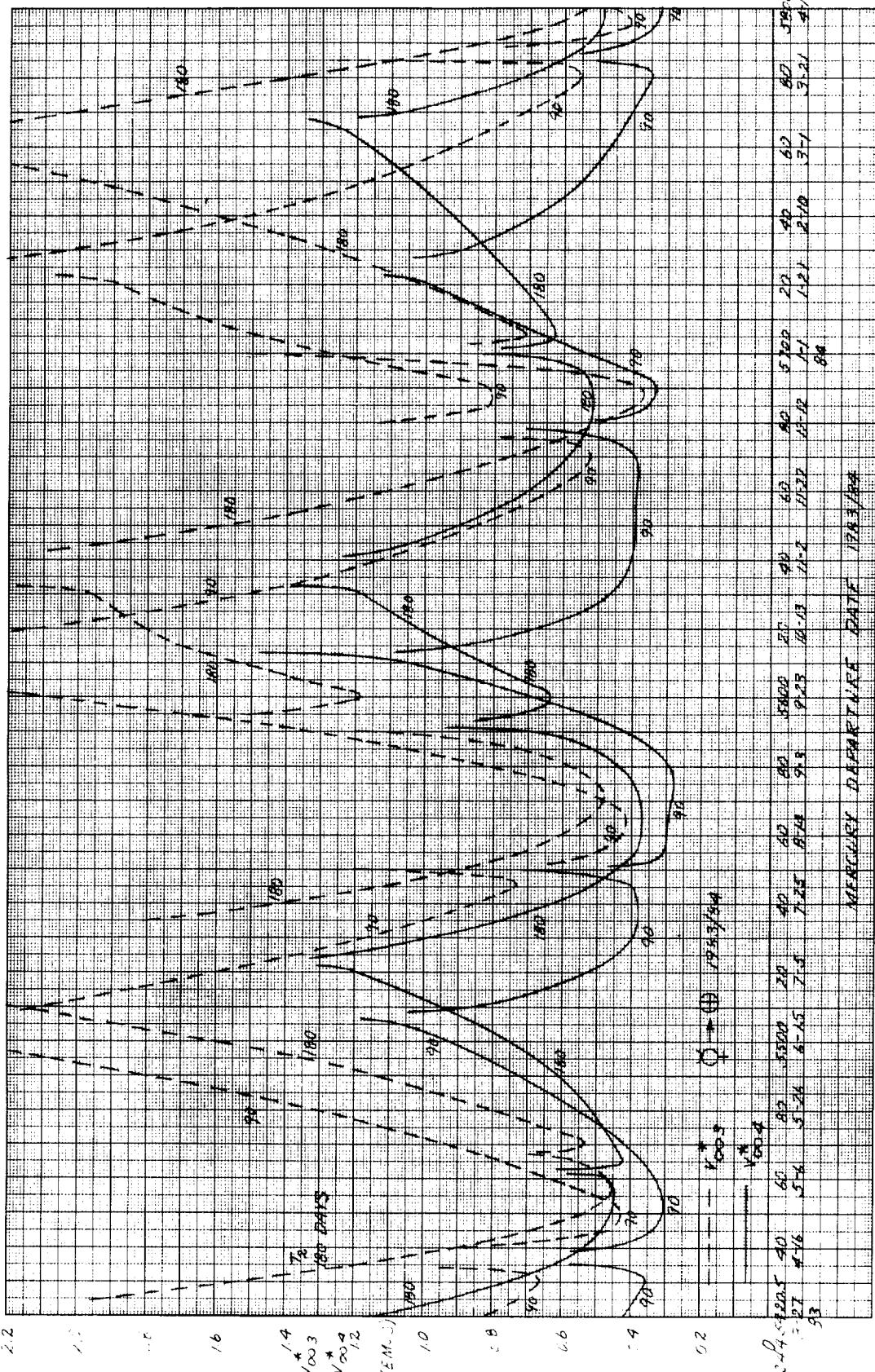
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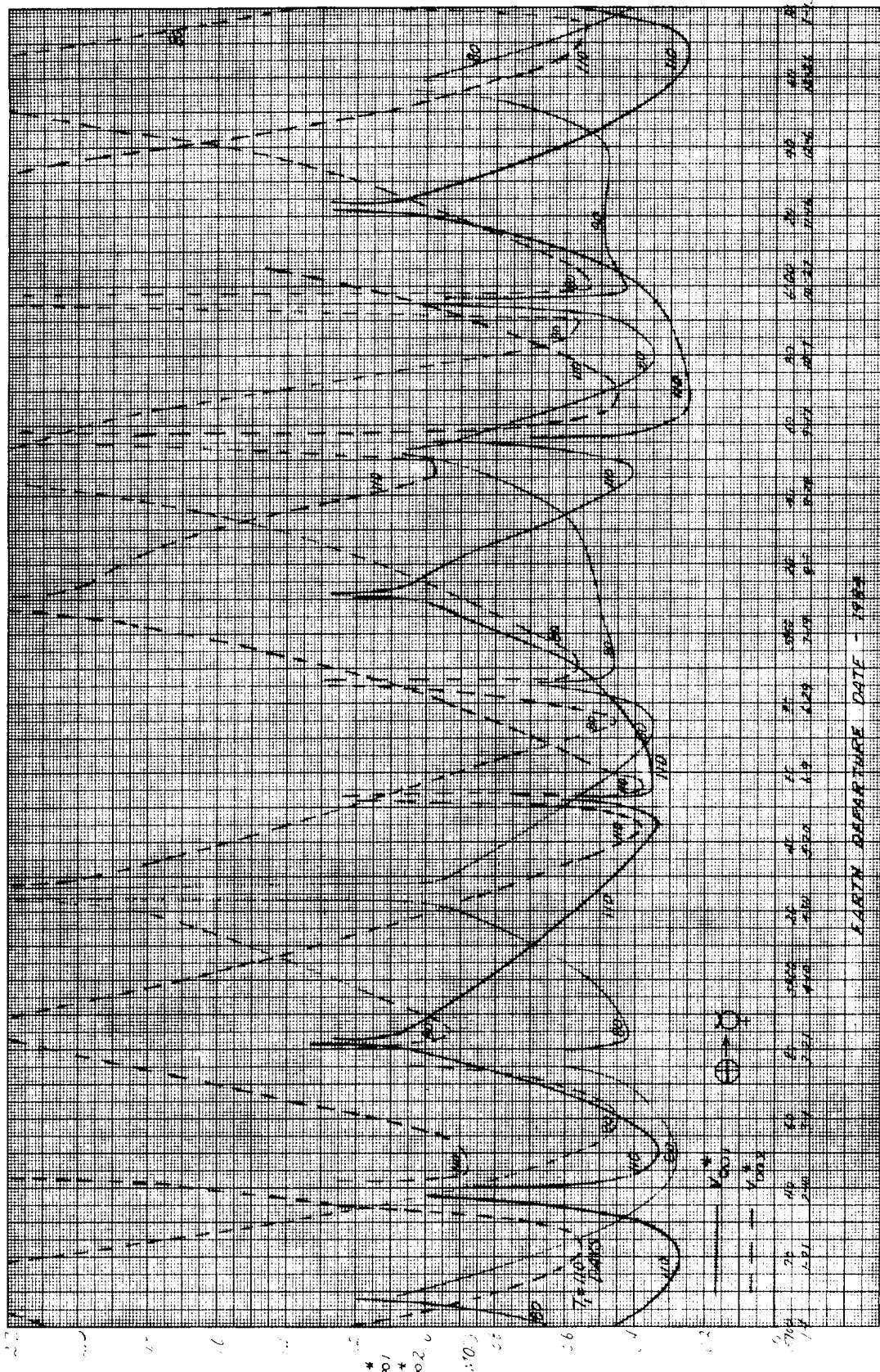
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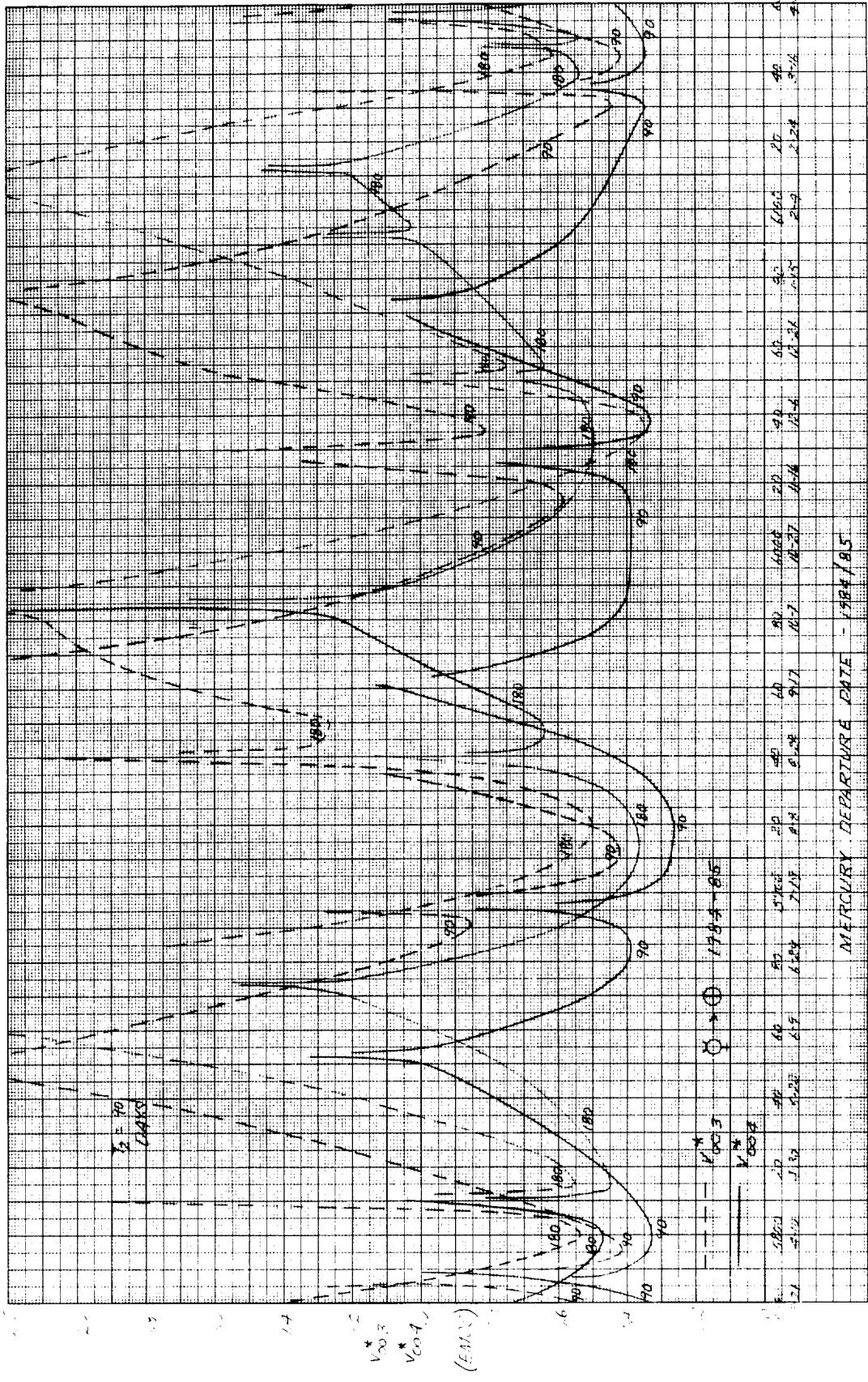


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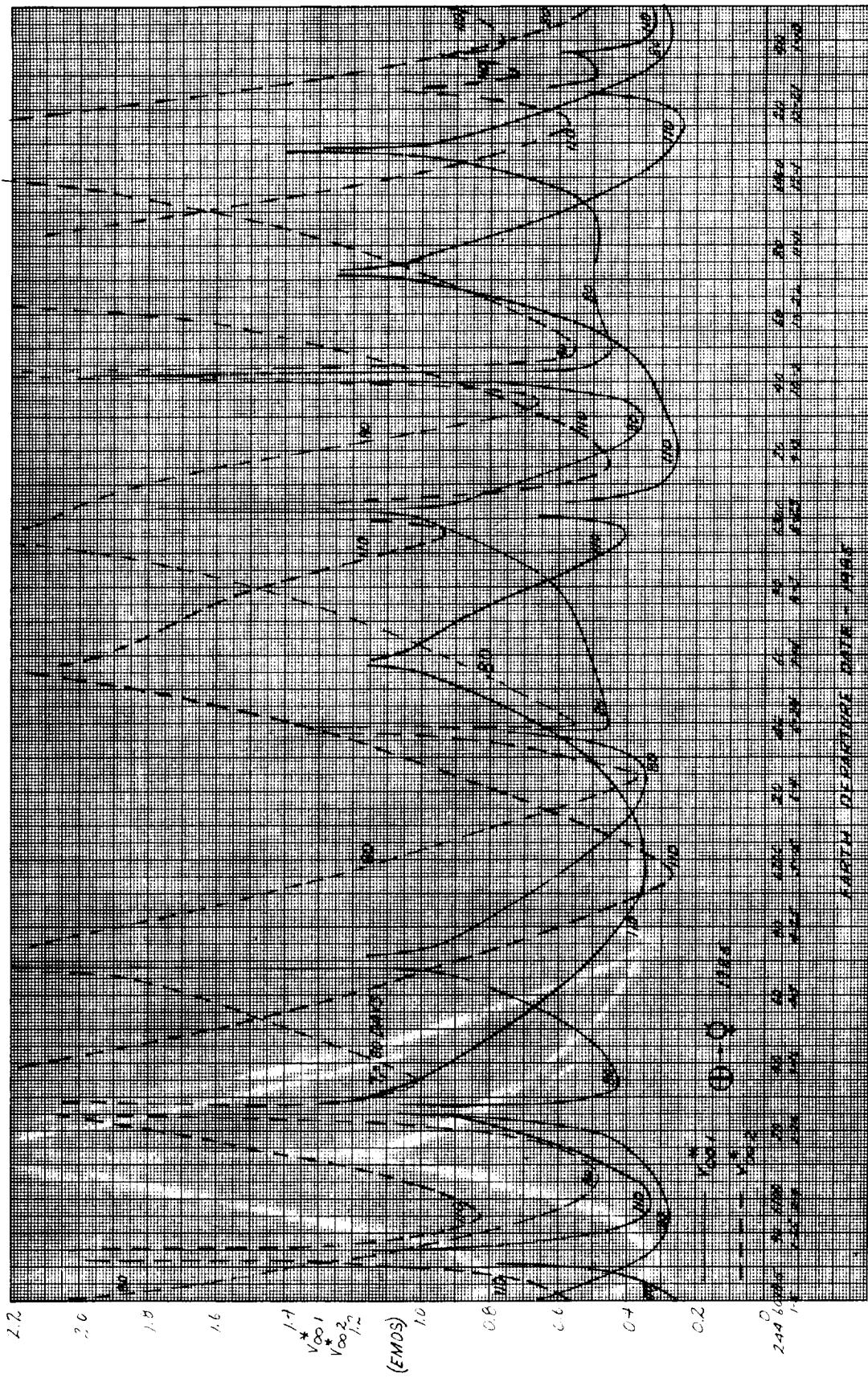
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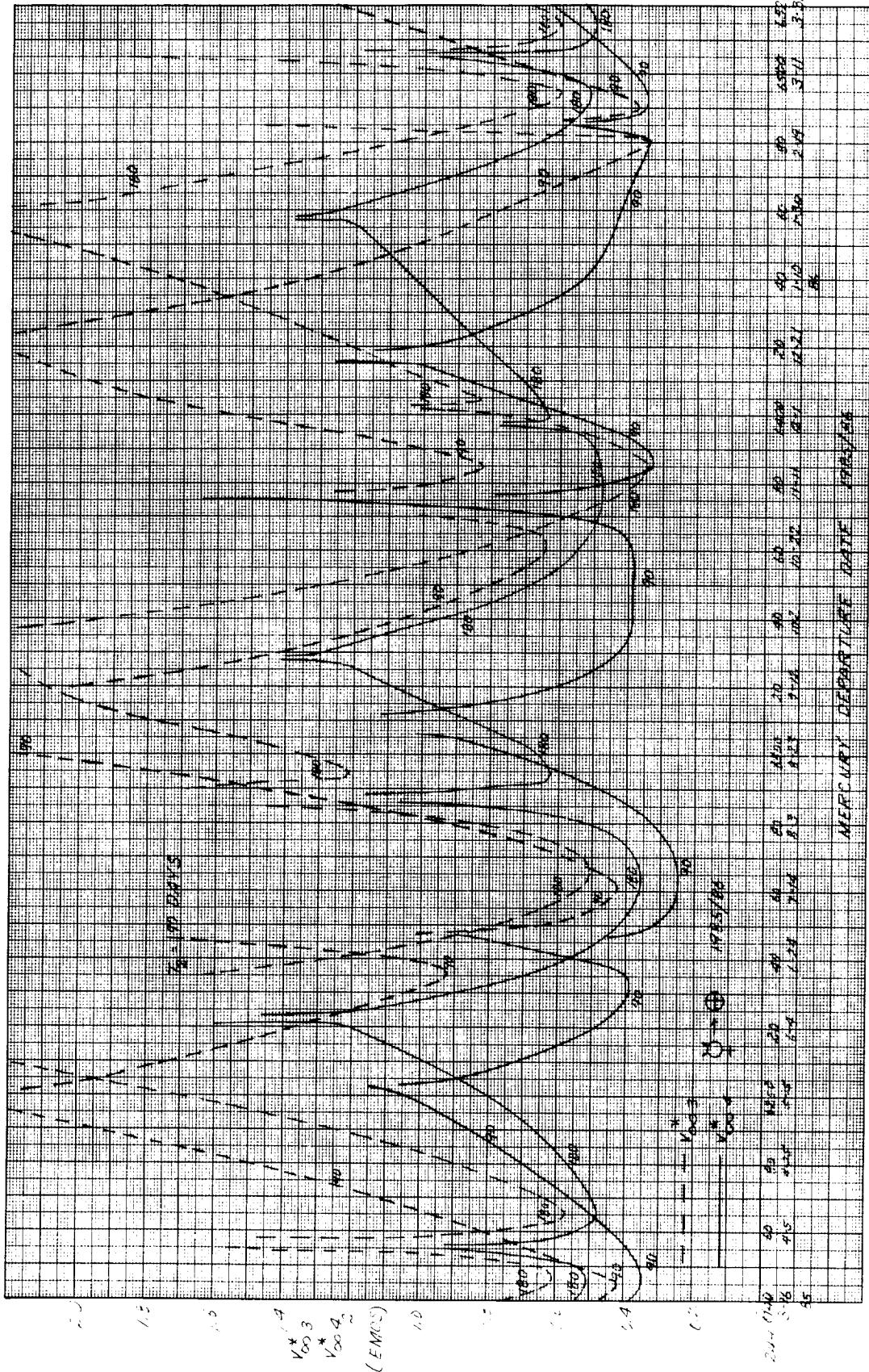


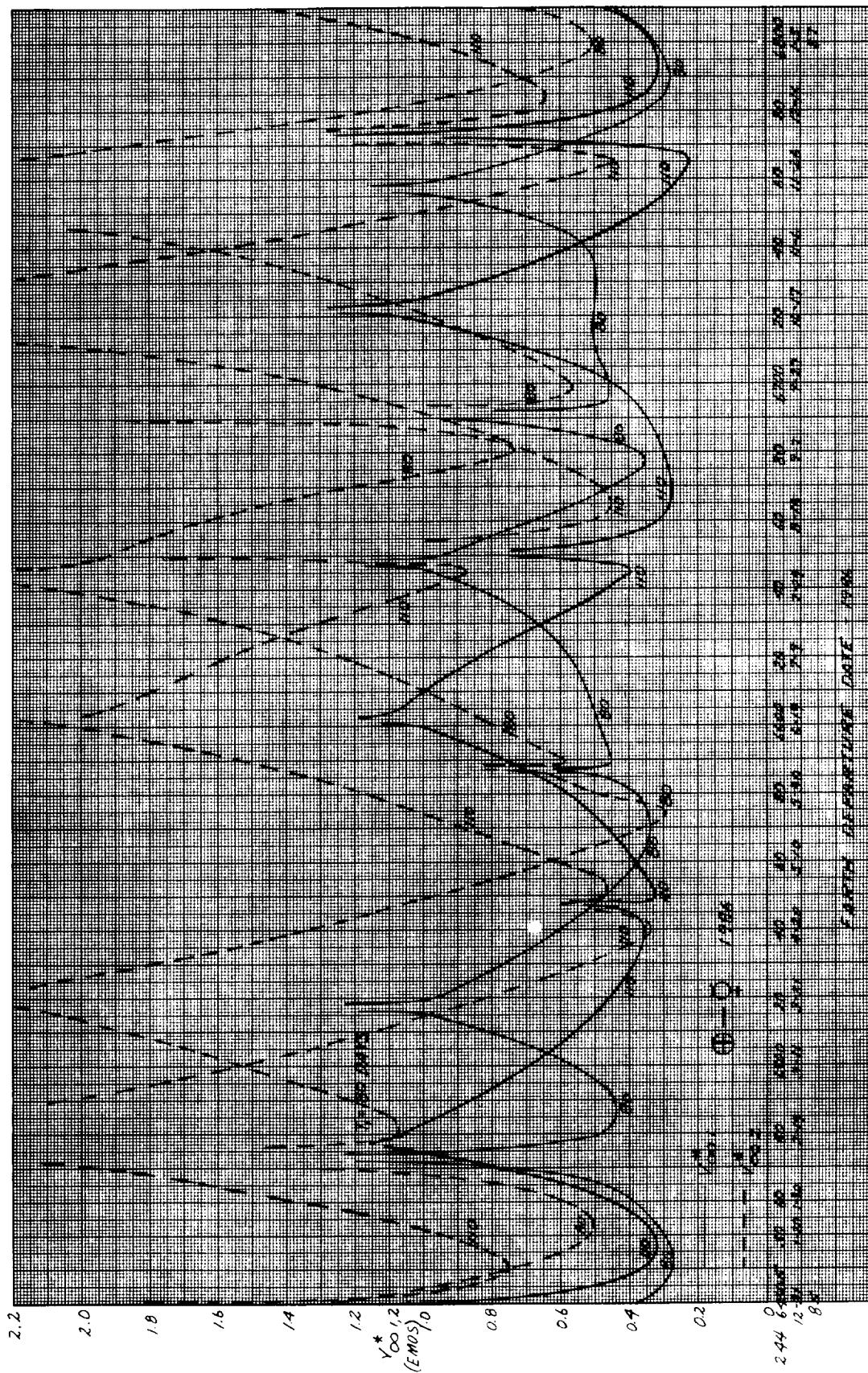
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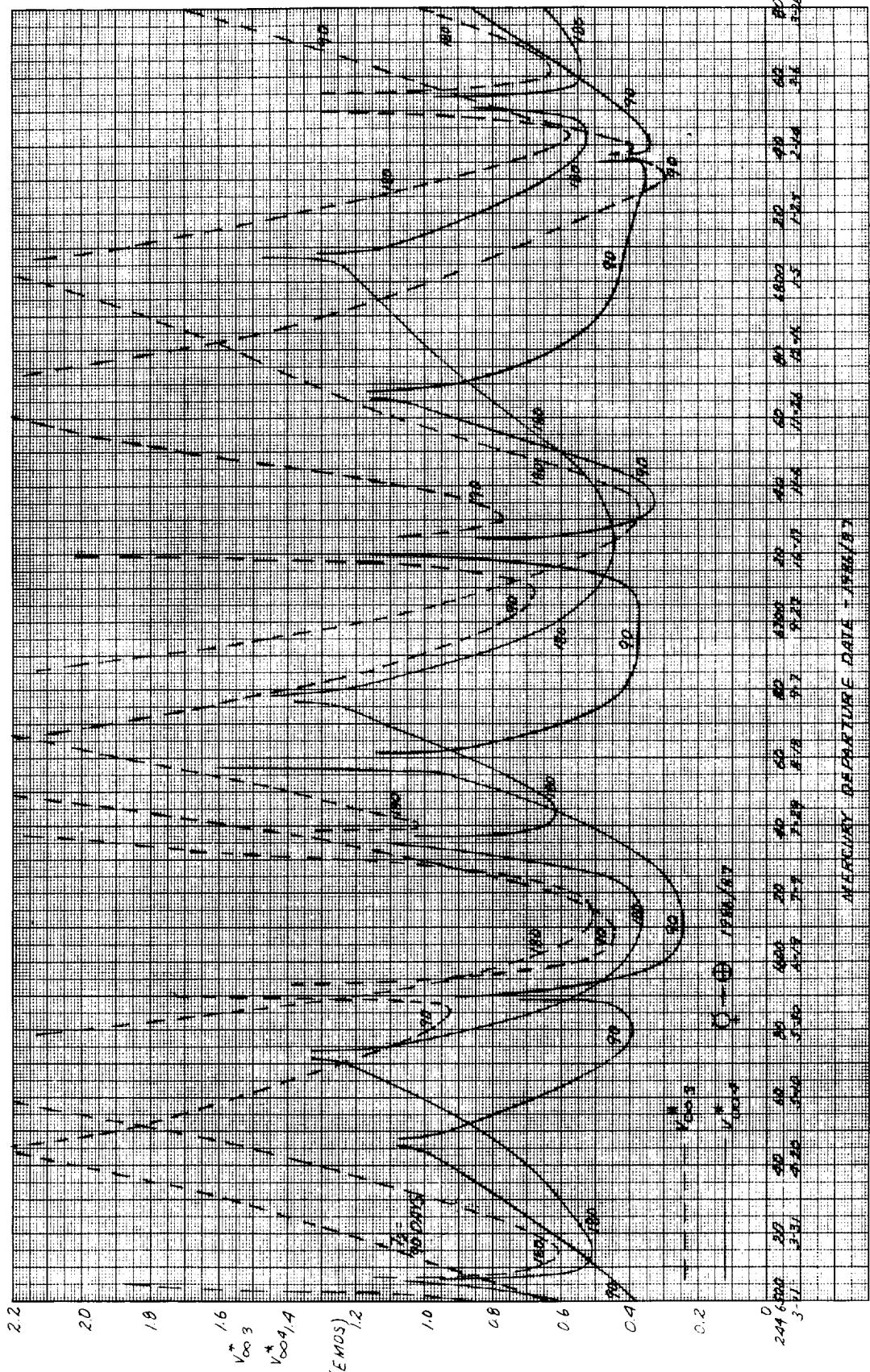


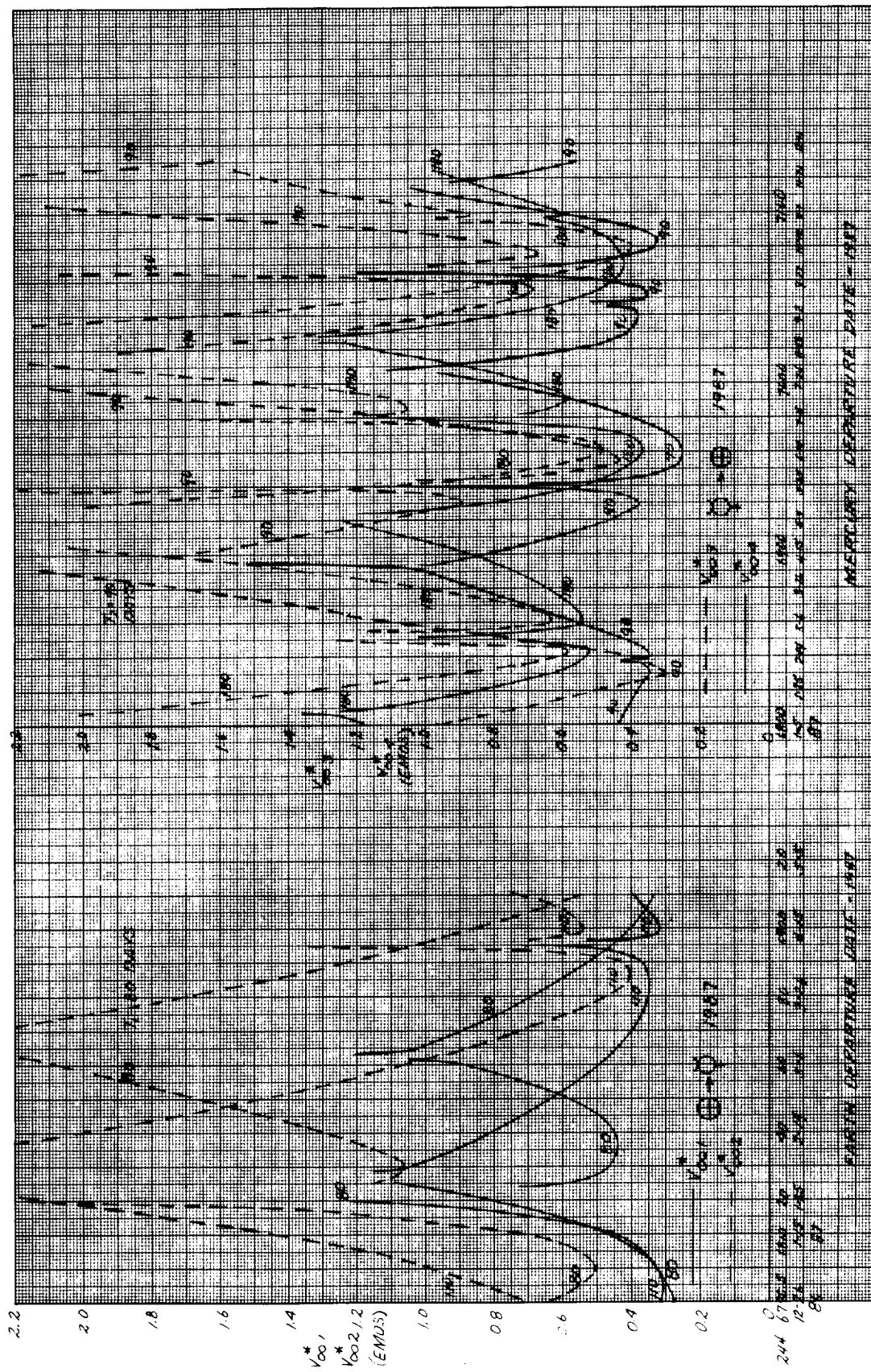
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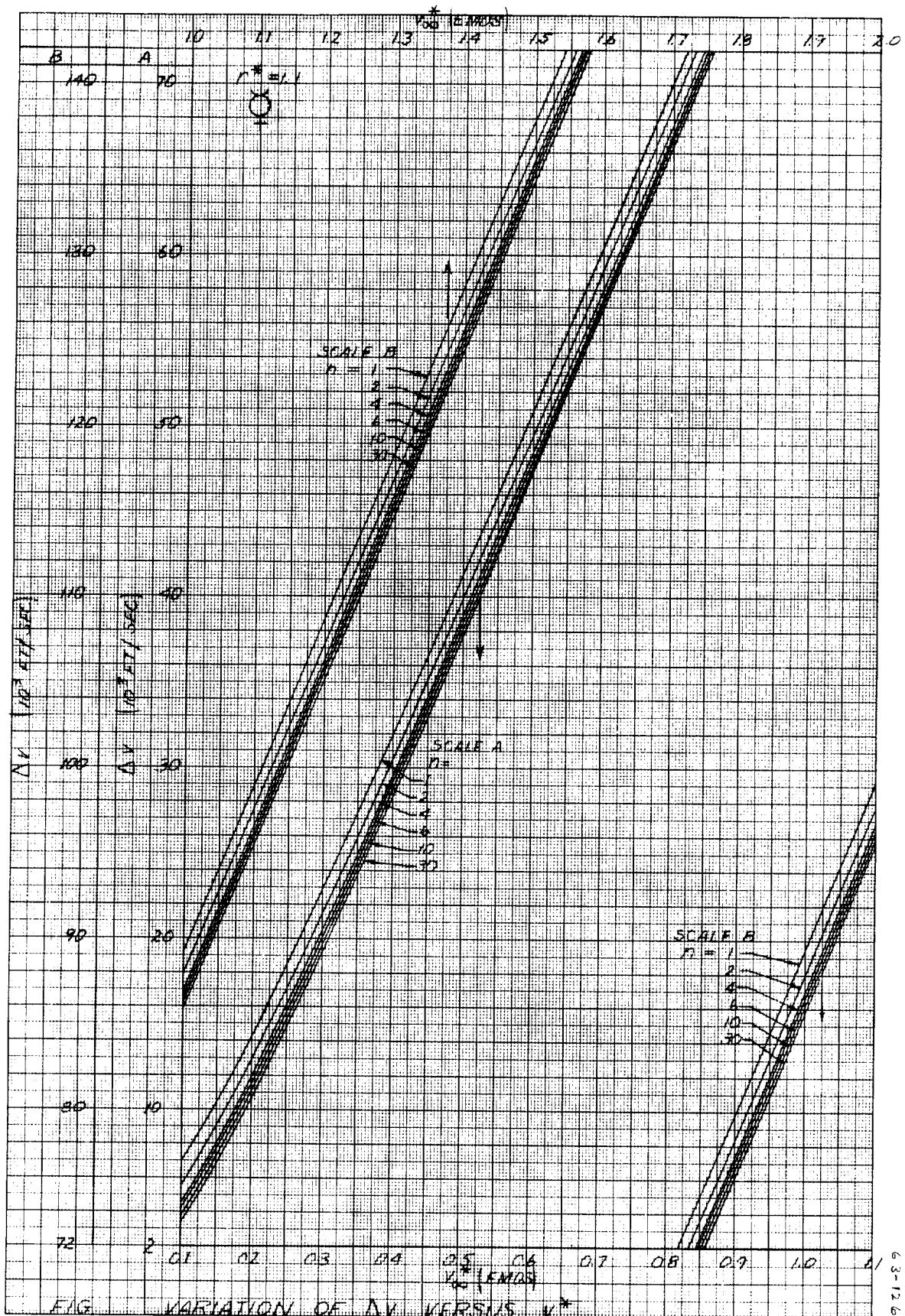


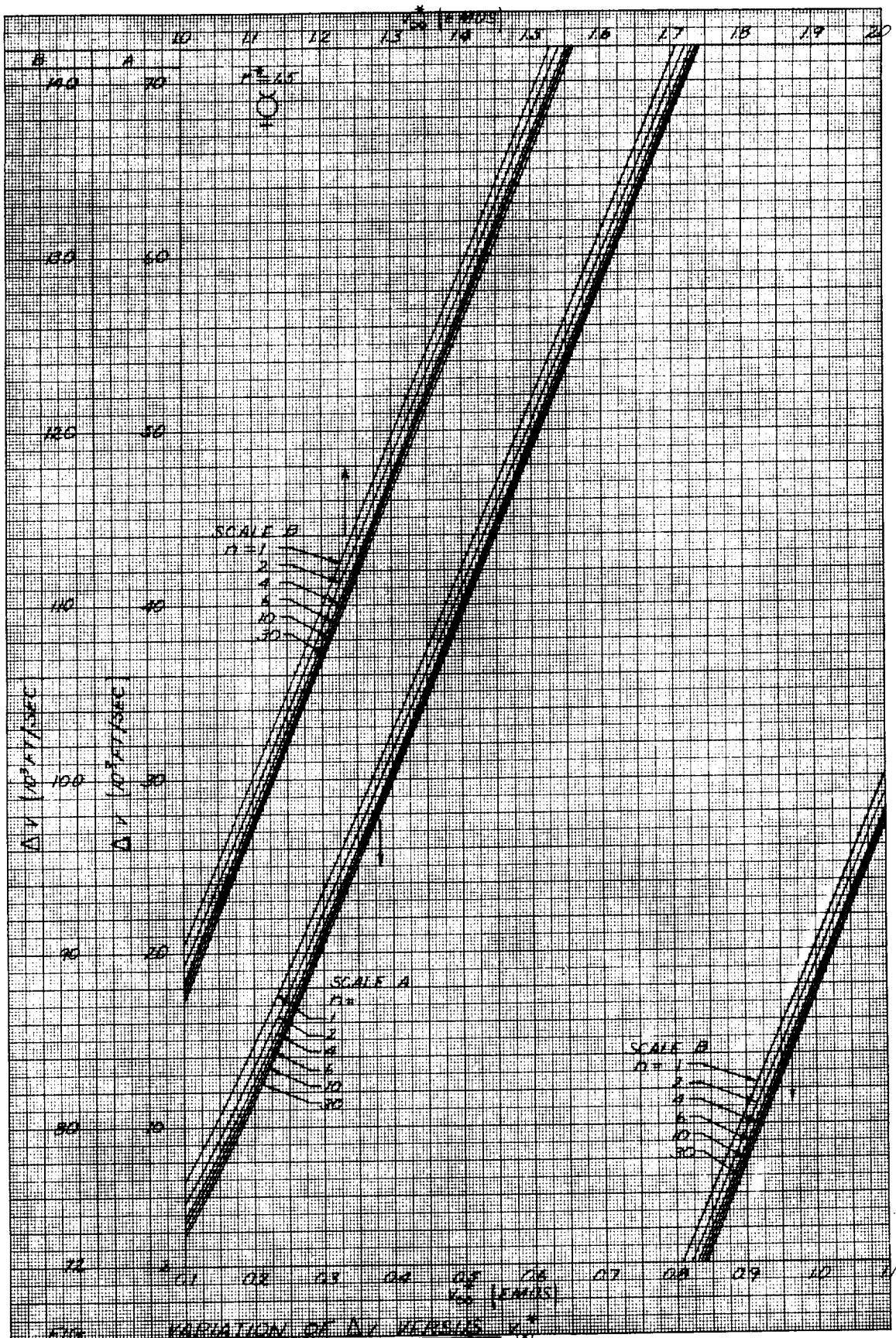




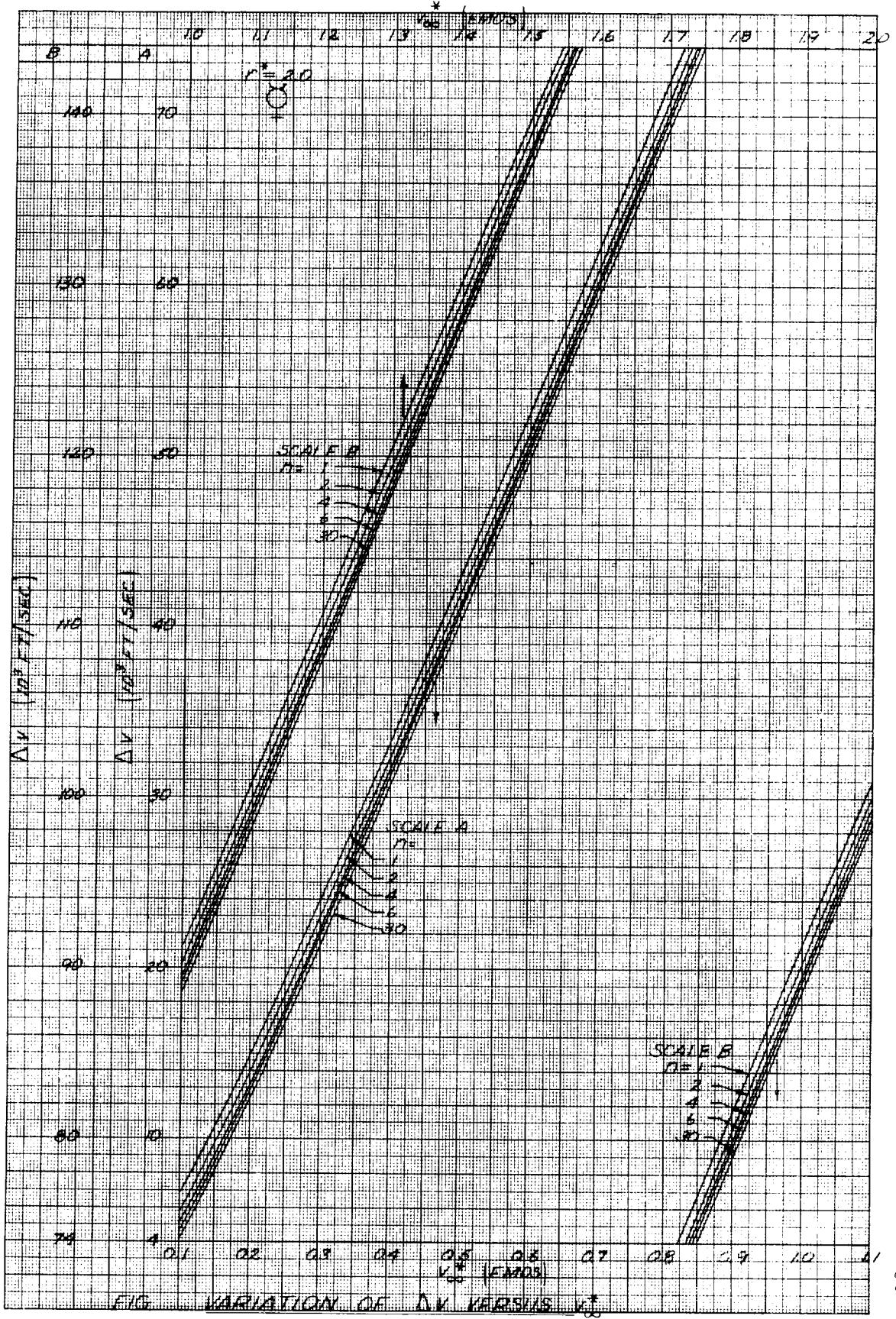






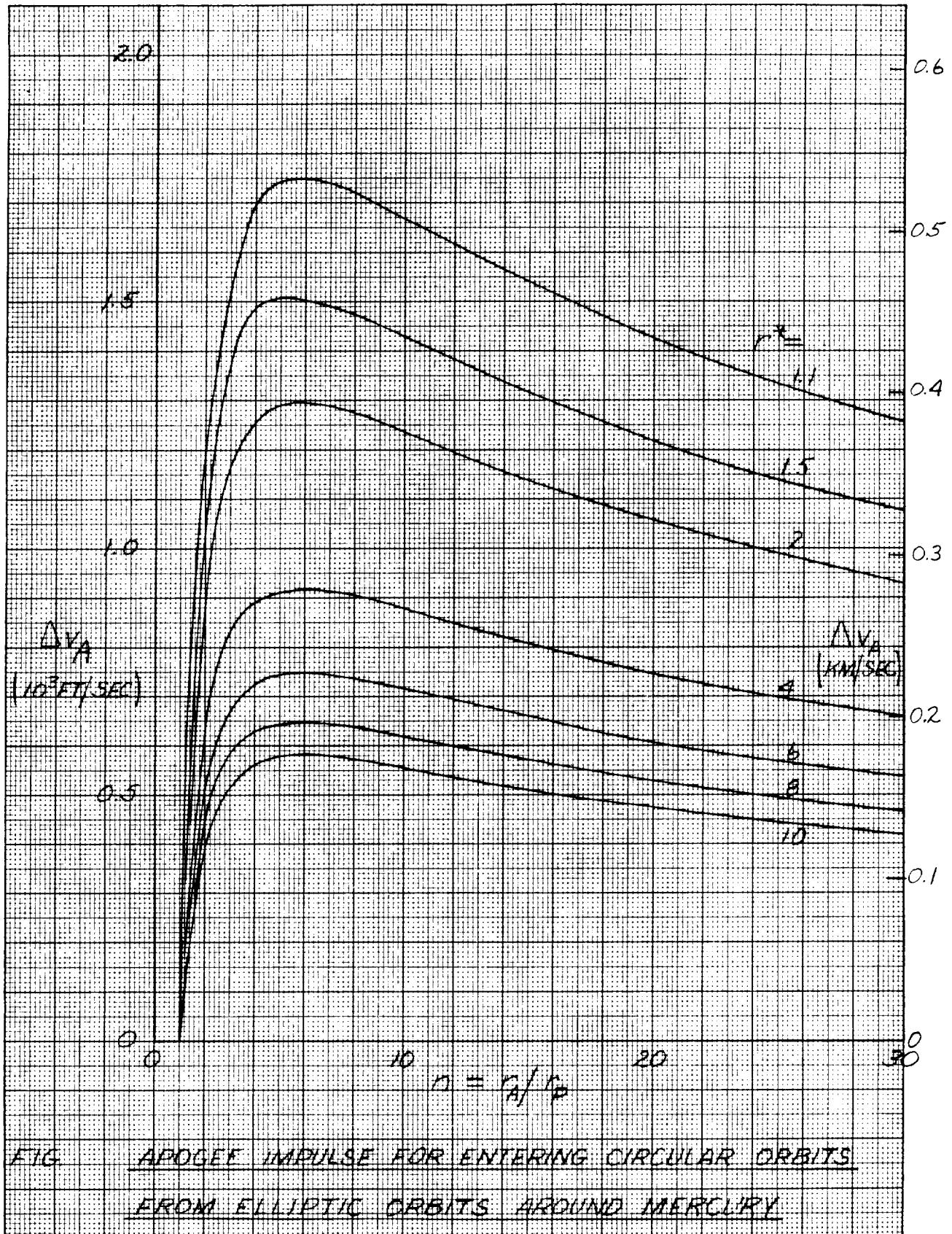


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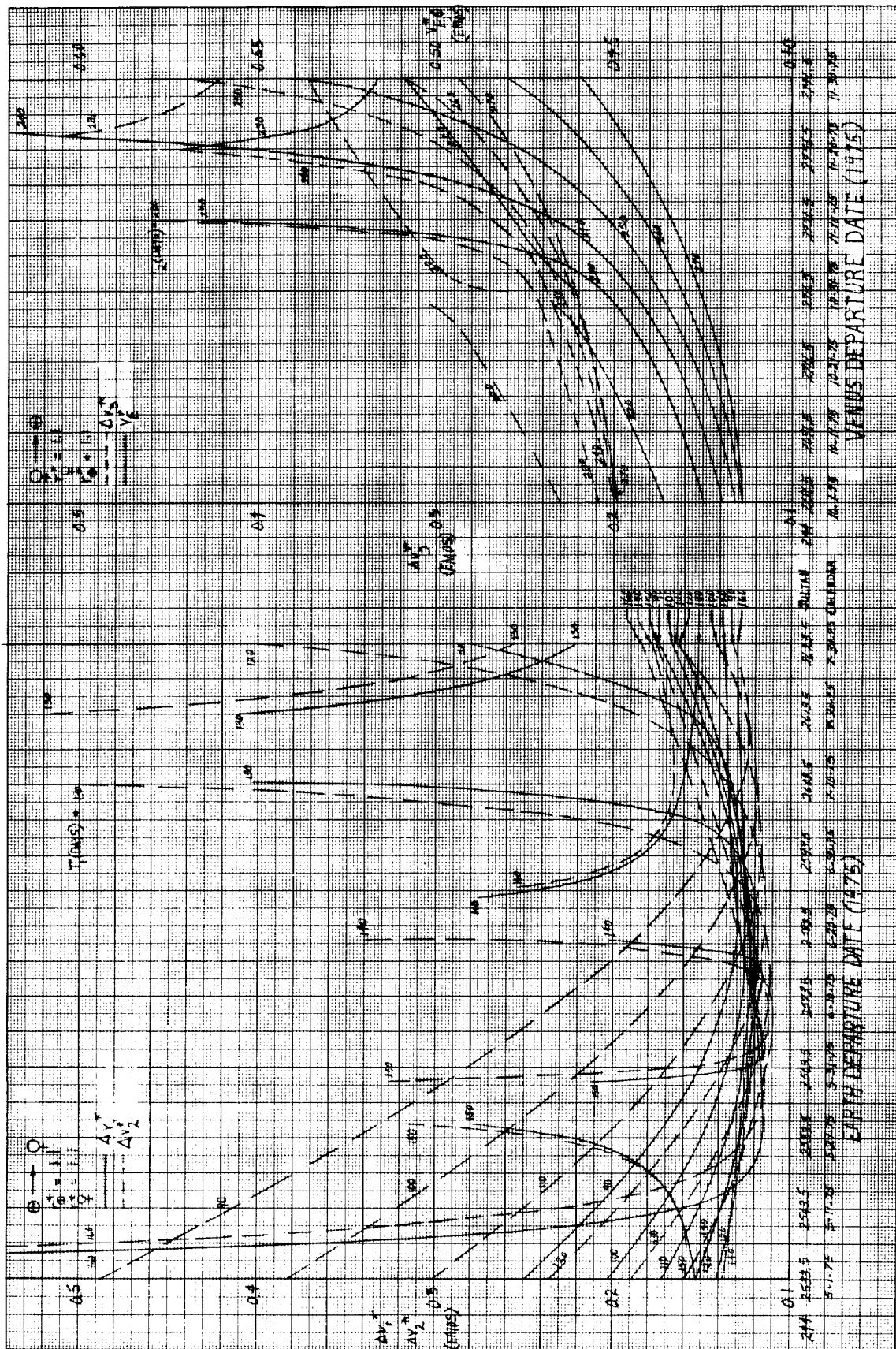
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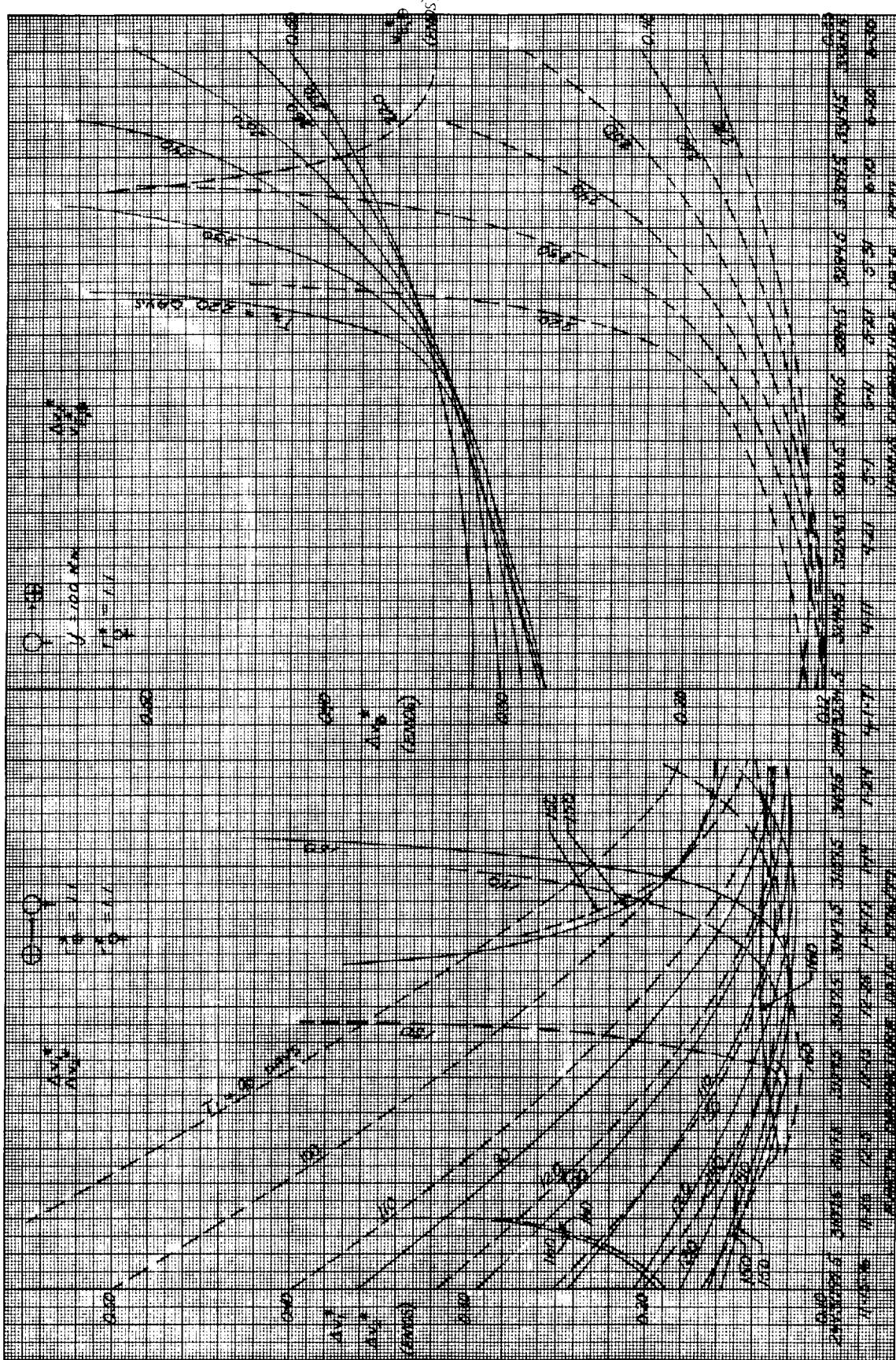


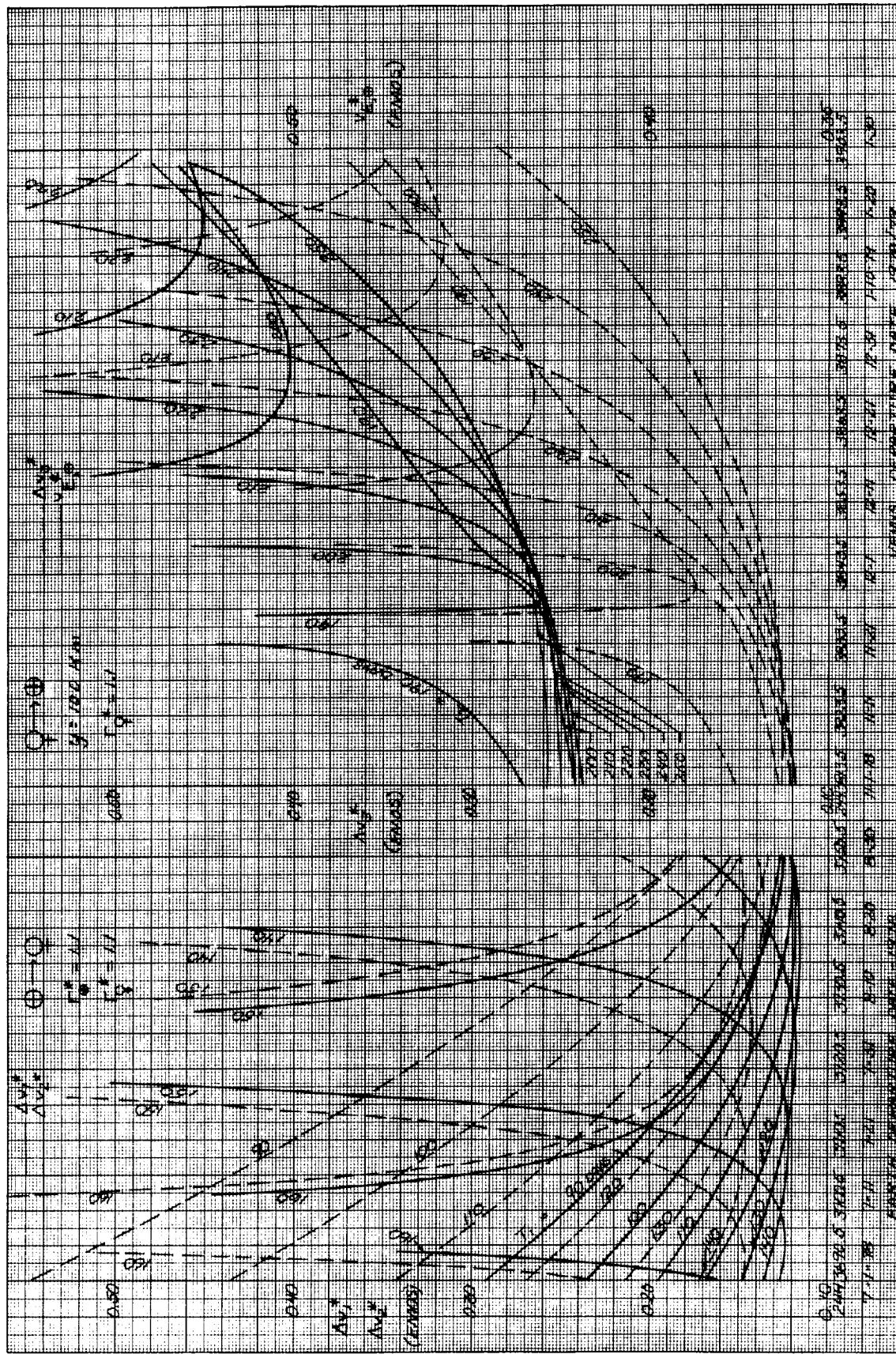
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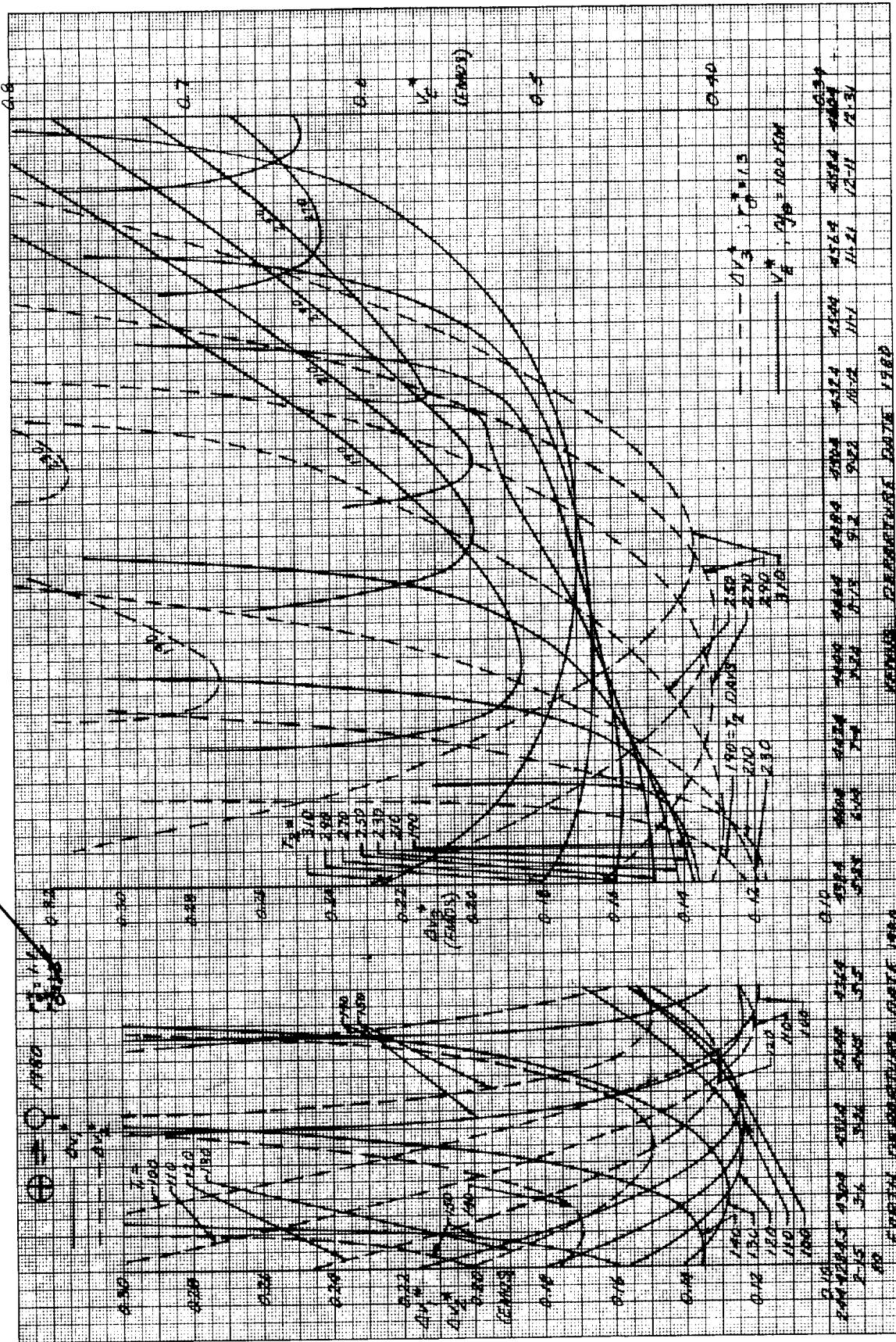
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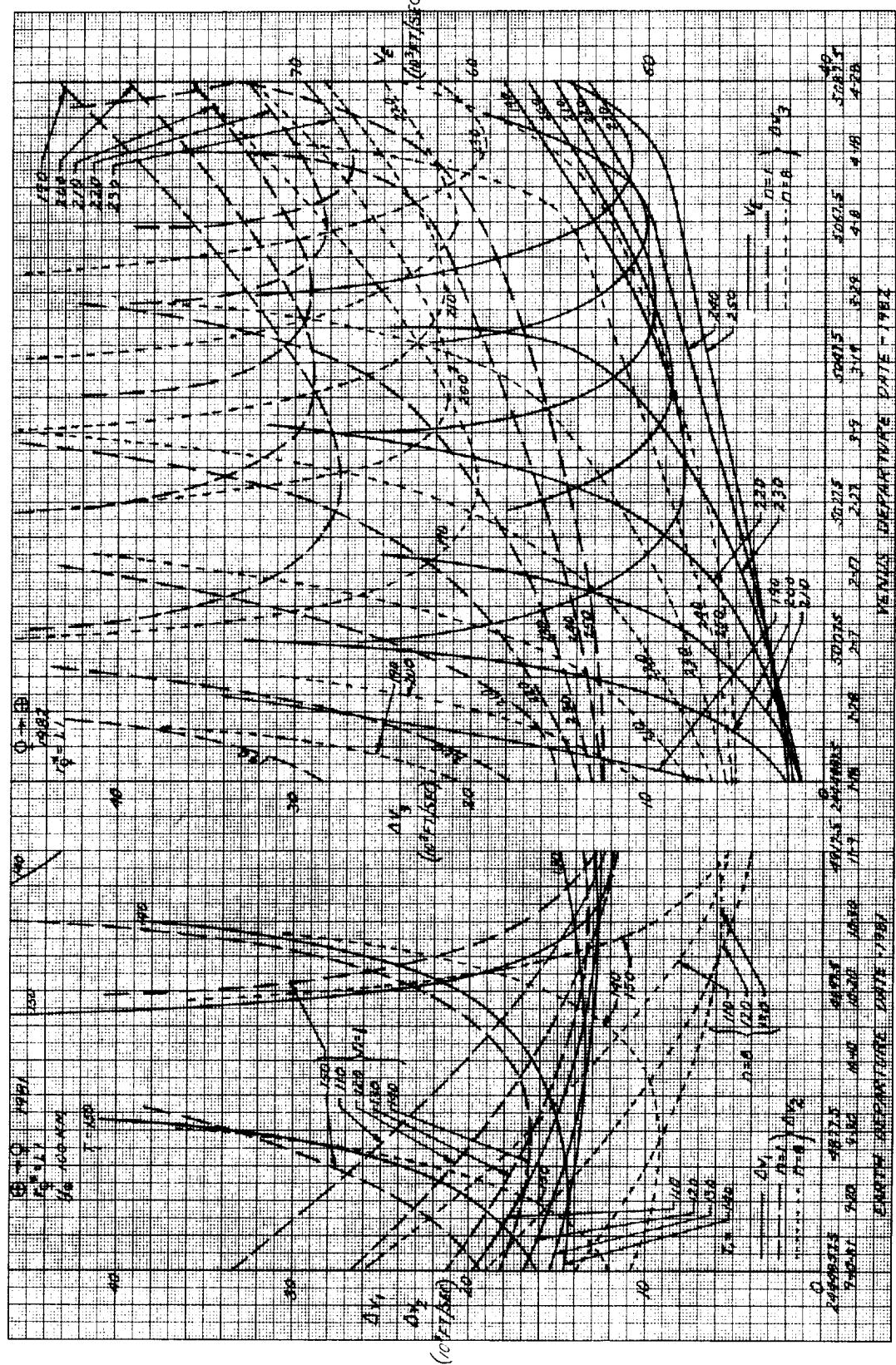




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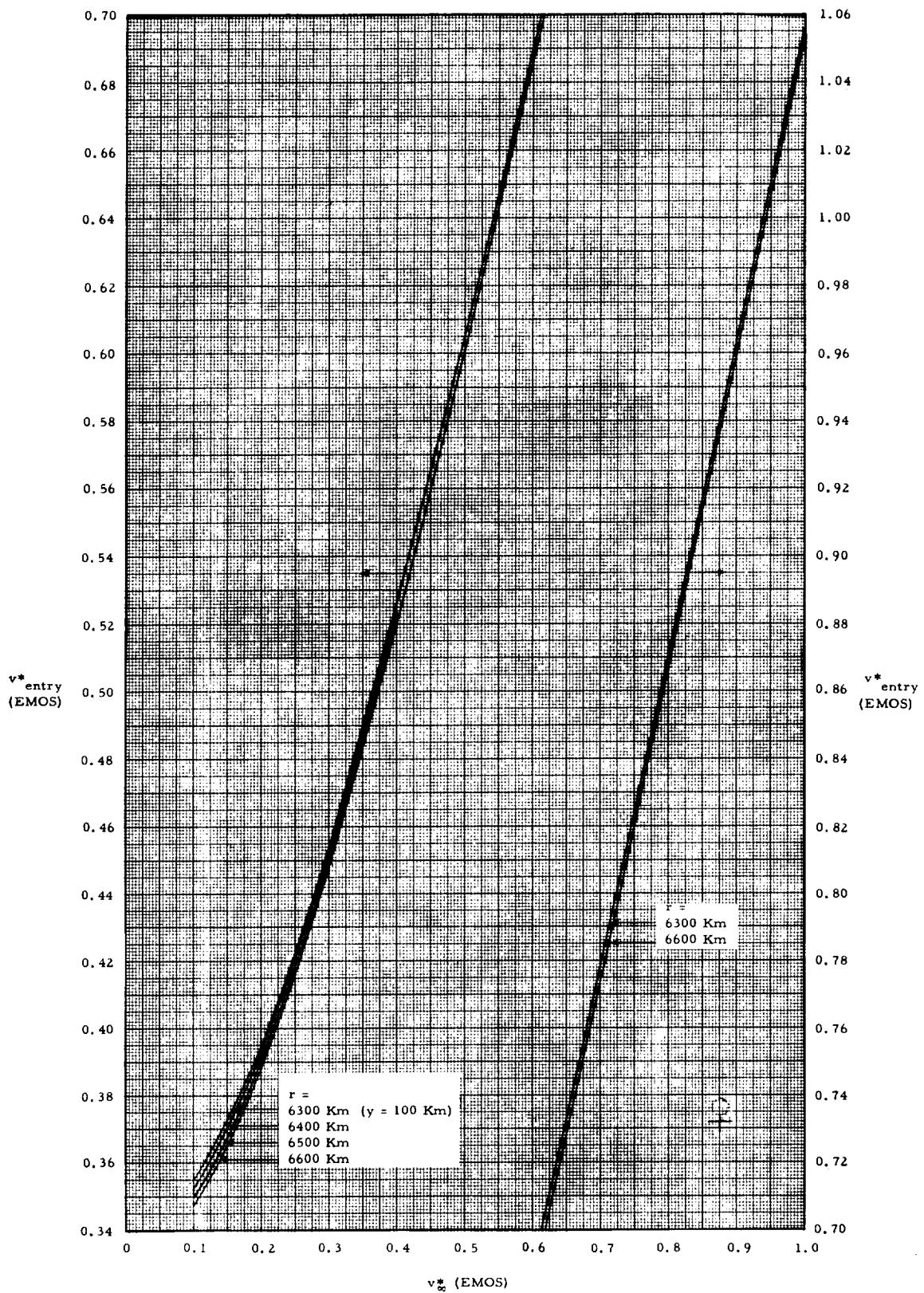


Fig. B-3 CORRELATION BETWEEN HYPERBOLIC EXCESS VELOCITY (v_{∞}^*) TO ATMOSPHERIC ENTRY VELOCITY (v^*_{entry}) AT DIFFERENT ALTITUDES FOR PLANET VENUS

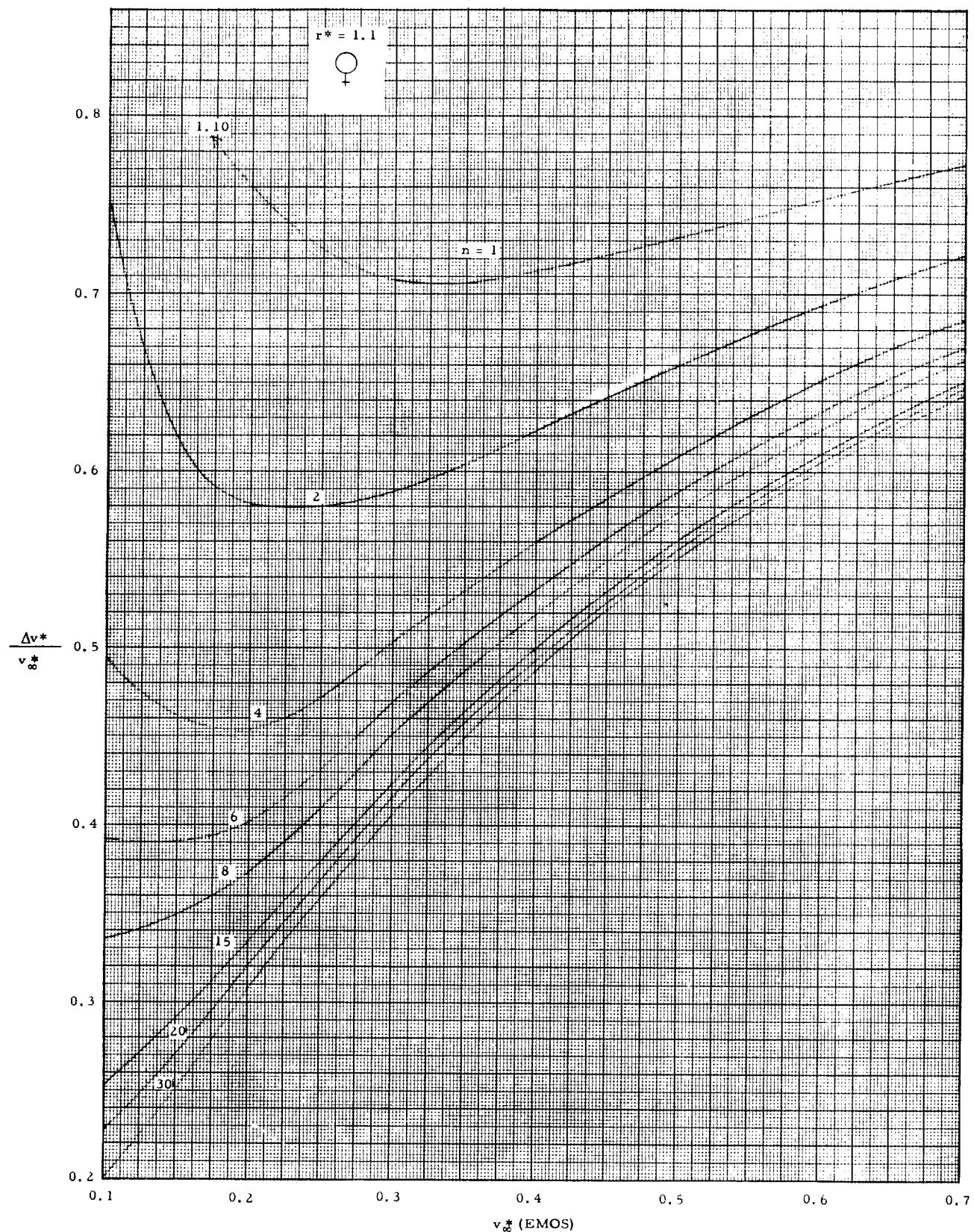


Fig. B-4 VARIATION OF $\frac{\Delta v^*}{v_\infty^*}$ VERSUS v_∞^* FOR CAPTURE IN CIRCULAR
AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^* = 1.1$)

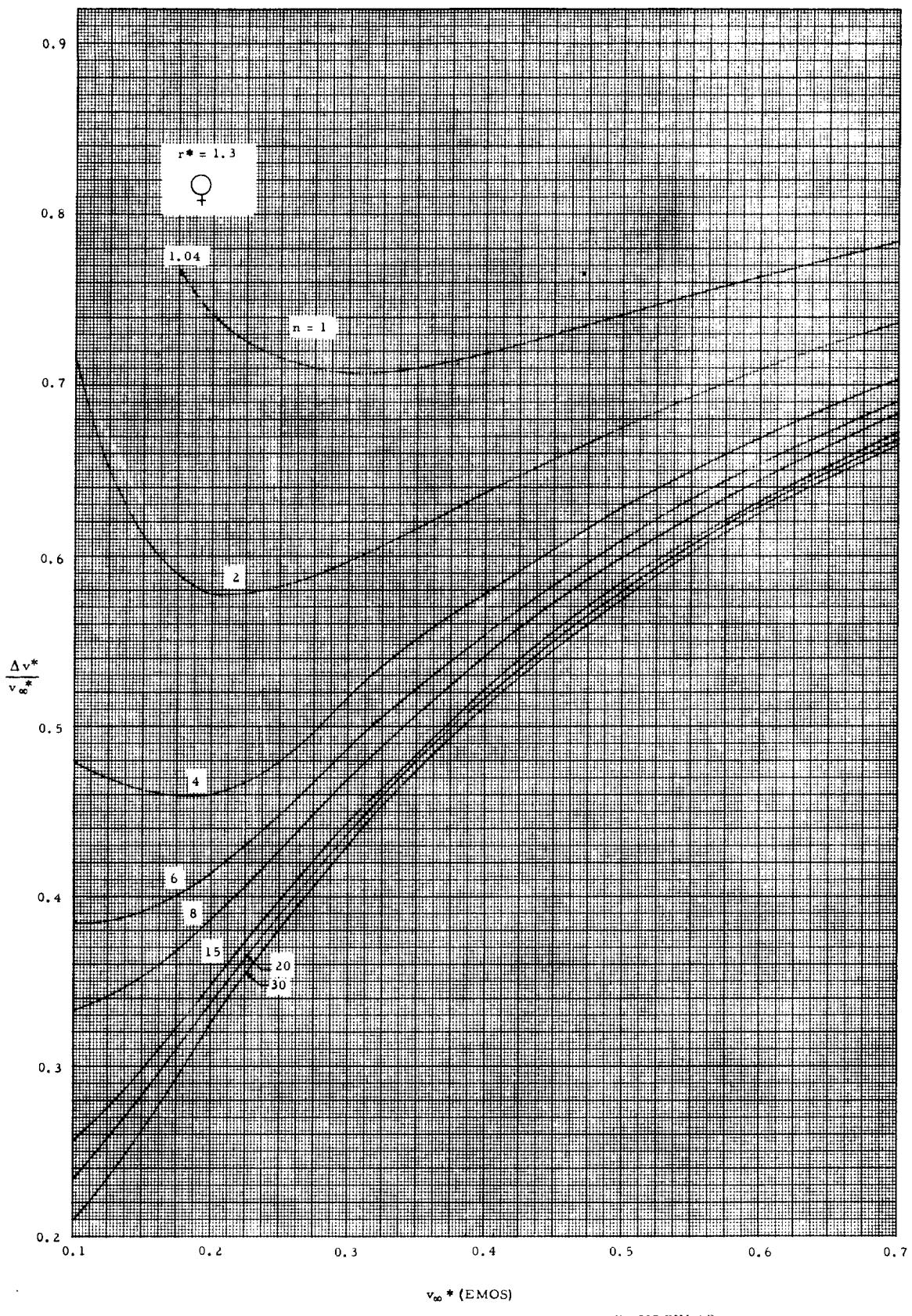


Fig. B-5 VARIATION OF $\Delta v^*/v_{\infty}^*$ VERSUS v_{∞}^* FOR CAPTURE IN CIRCULAR AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^* = 1.3$)

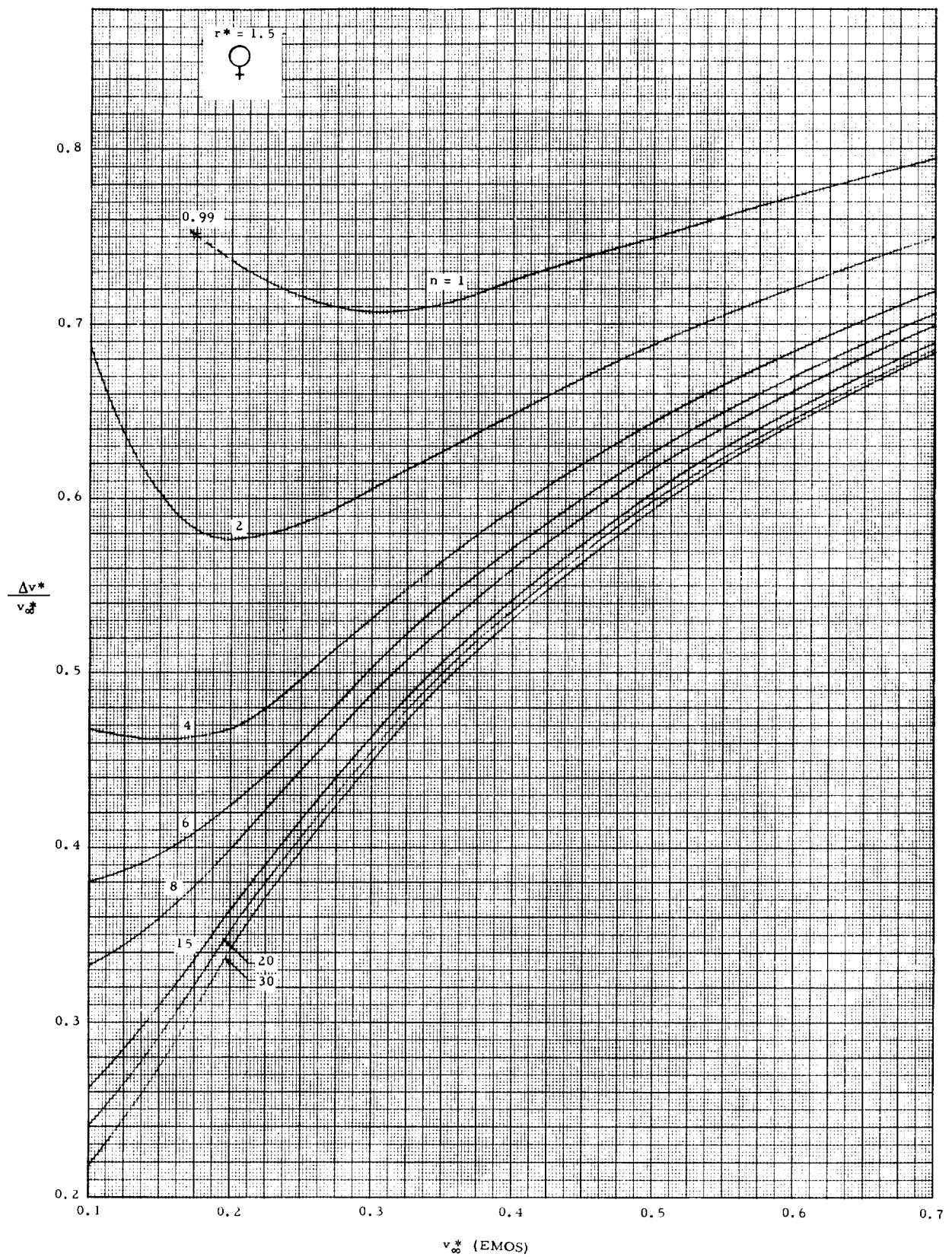


Fig. B-6 VARIATION OF $\frac{\Delta v^*}{v_\infty^*}$ VERSUS v_∞^* FOR CAPTURE IN CIRCULAR AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^* = 1.5$)

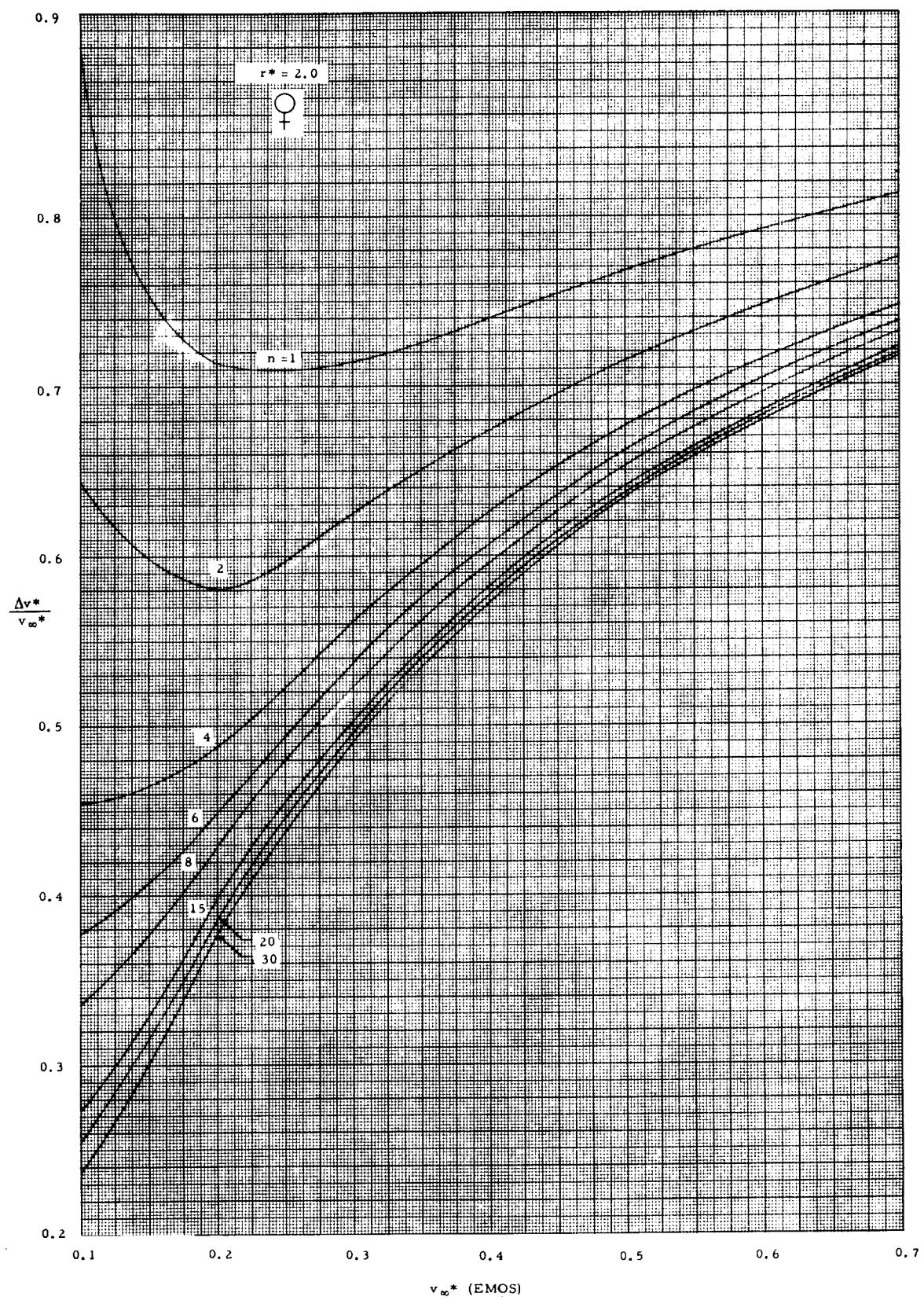


Fig. B-7 VARIATION OF $\frac{\Delta v^*}{v_{\infty}^*}$ VERSUS v_{∞}^* FOR CAPTURE IN CIRCULAR AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^* = 2.0$)

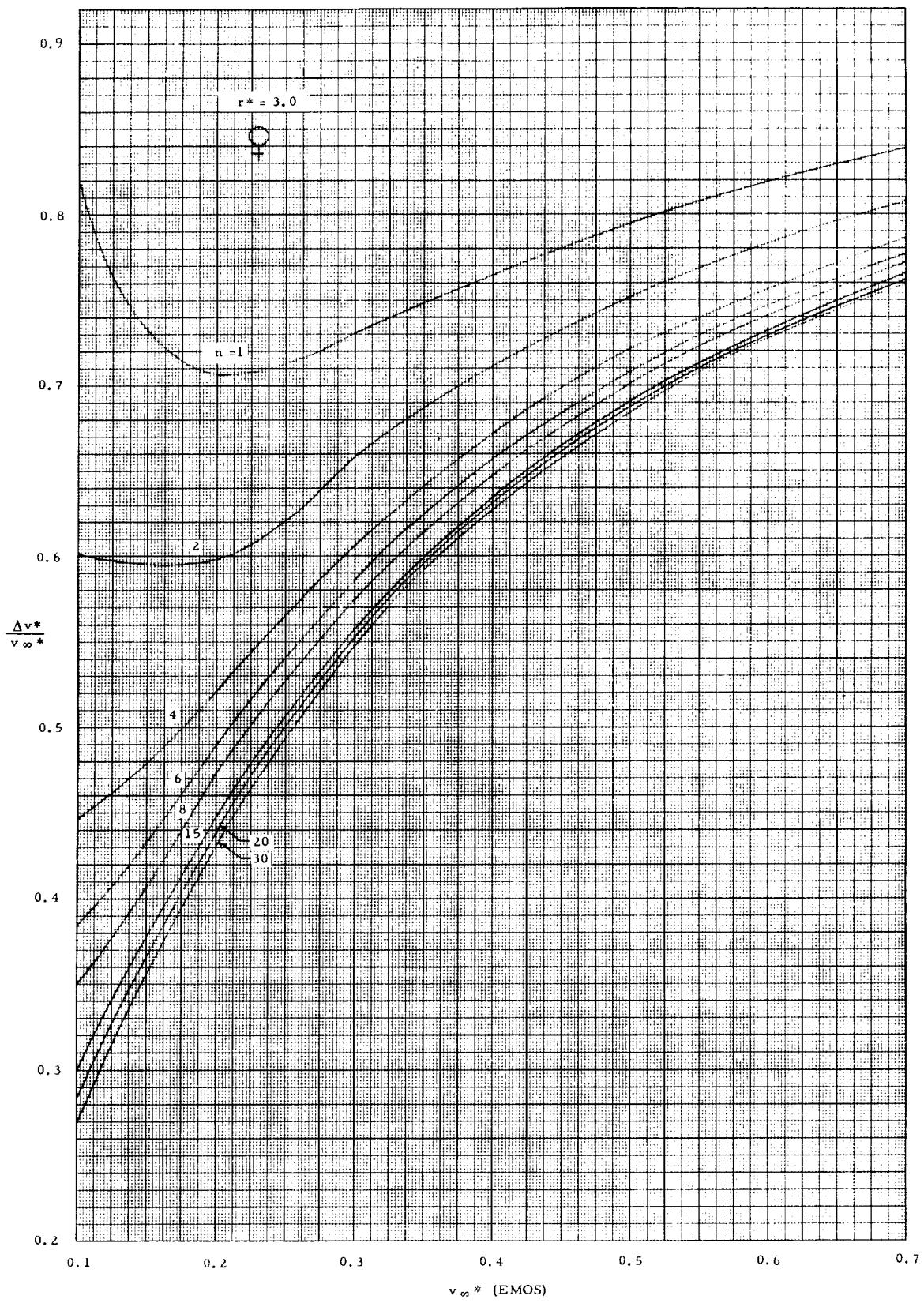


Fig. B-8 VARIATION OF $\frac{\Delta v^*}{v_{\infty}^*}$ VERSUS v_{∞}^* FOR CAPTURE IN CIRCULAR AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^*=3.0$)

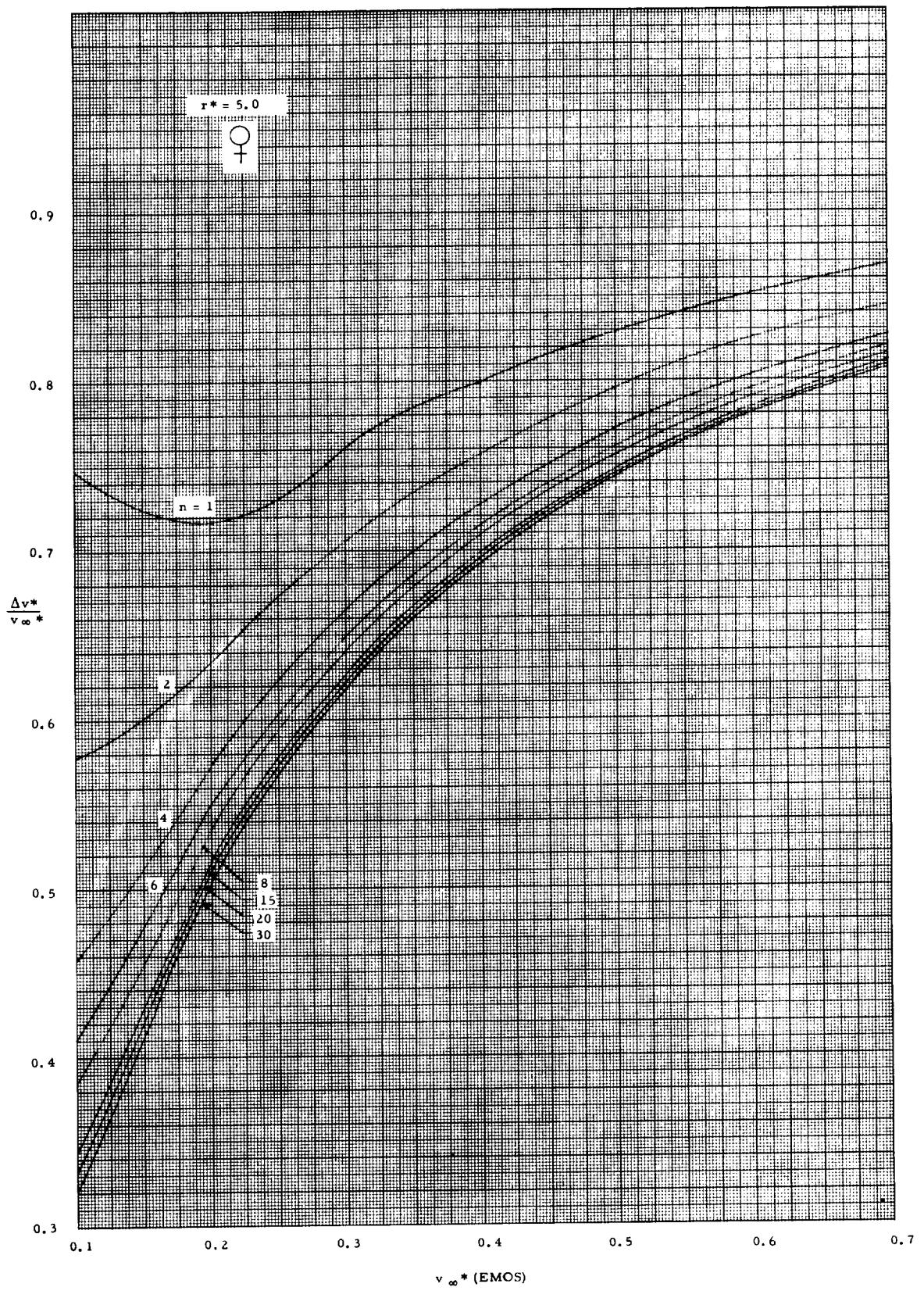
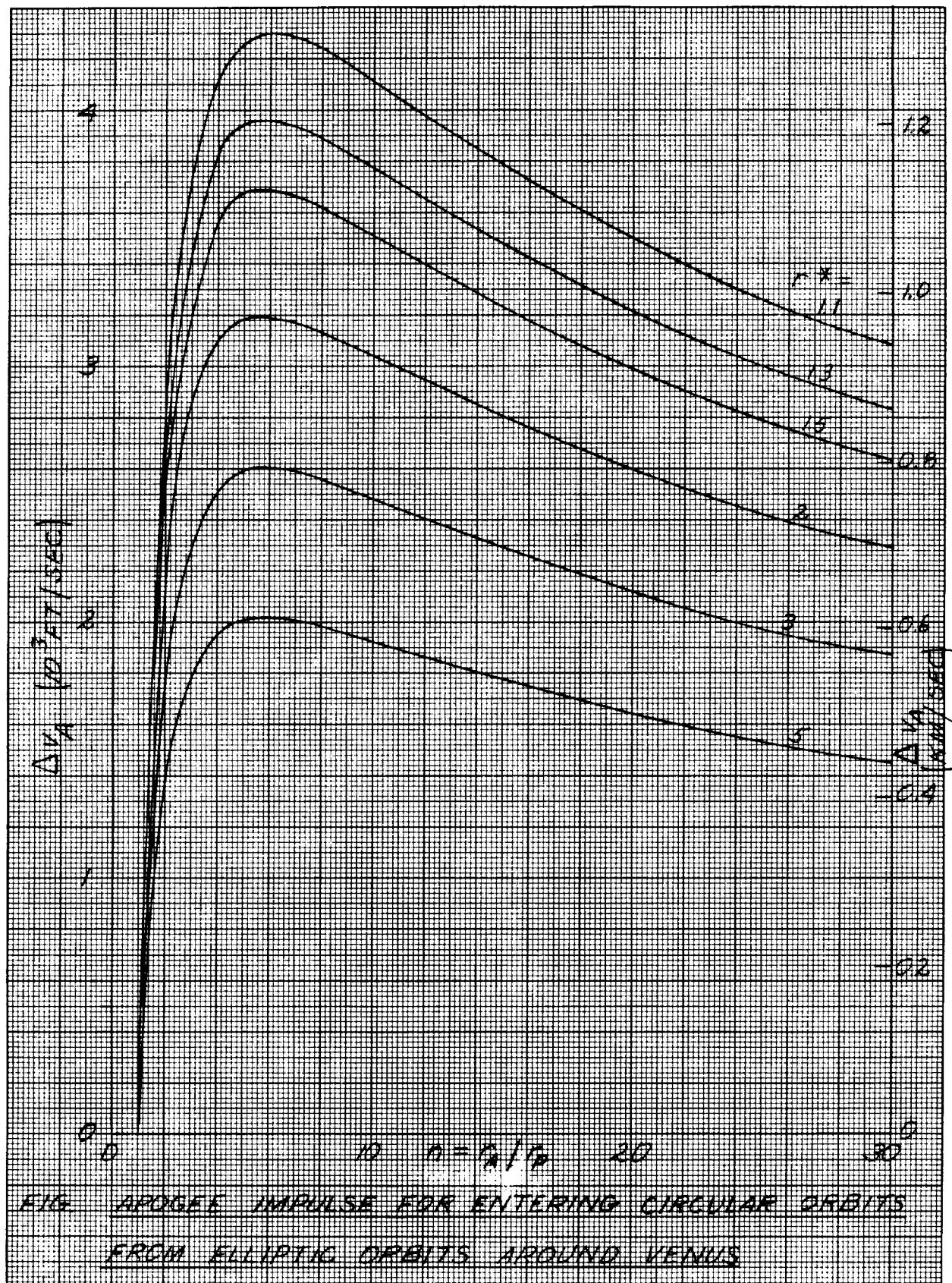
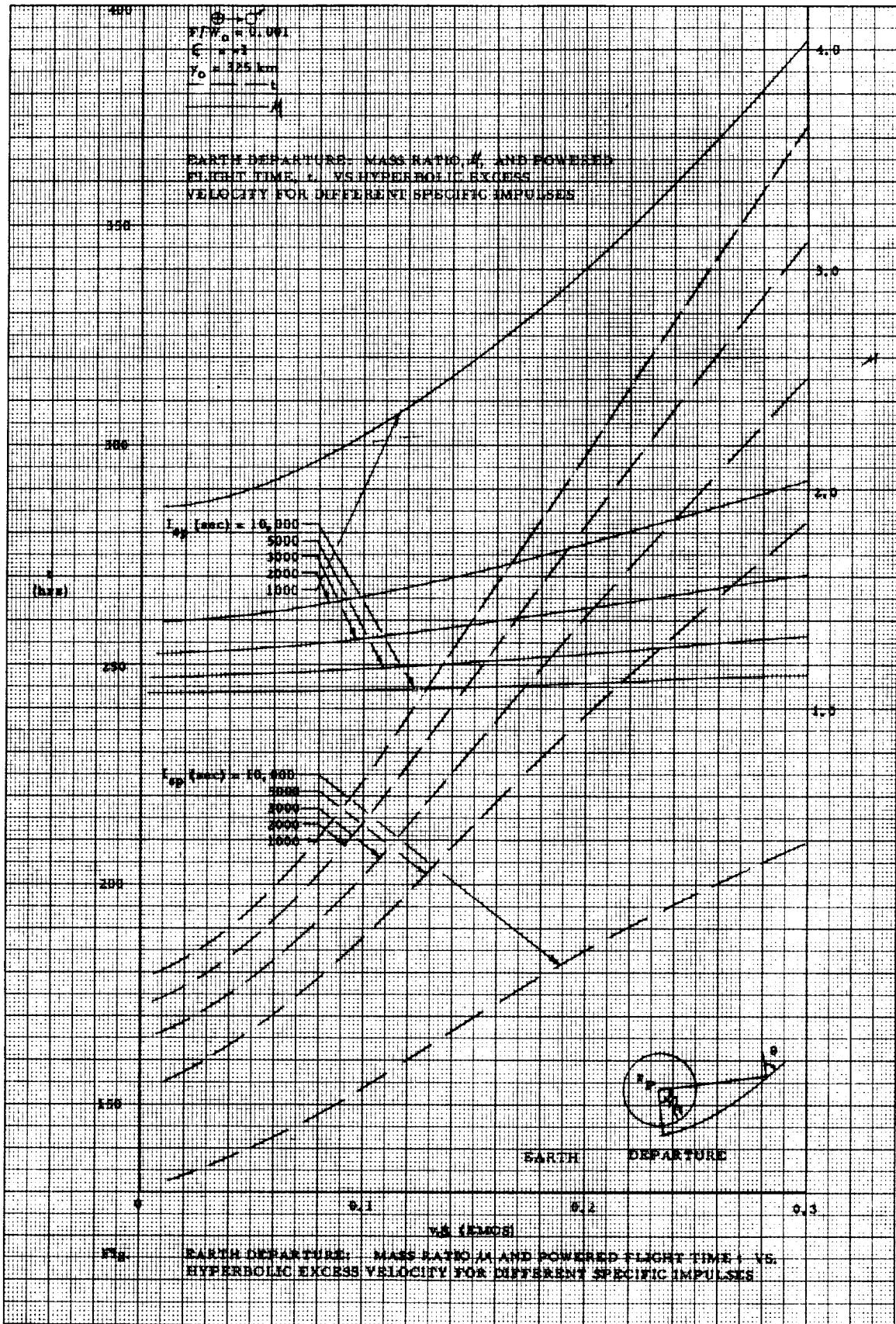
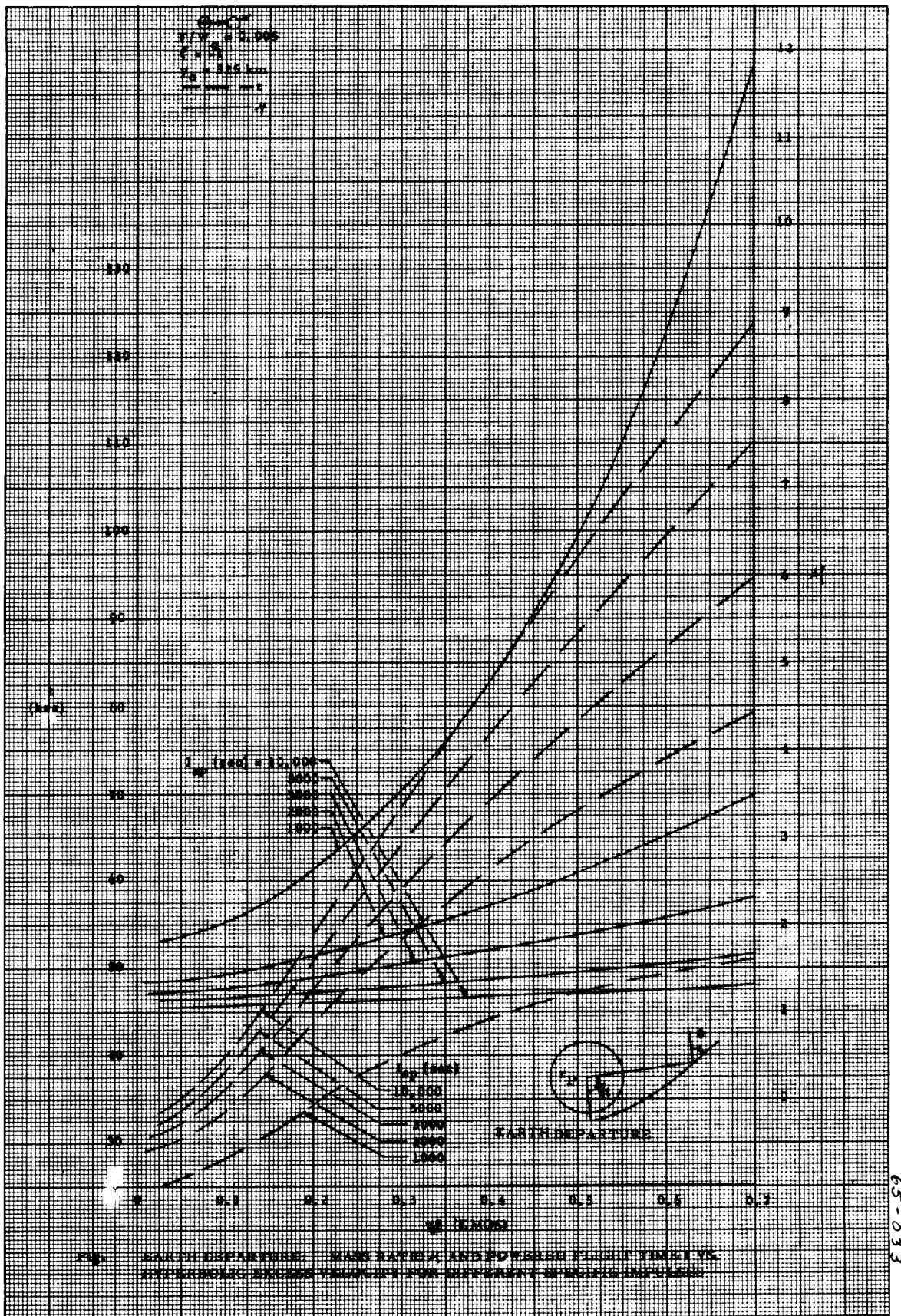


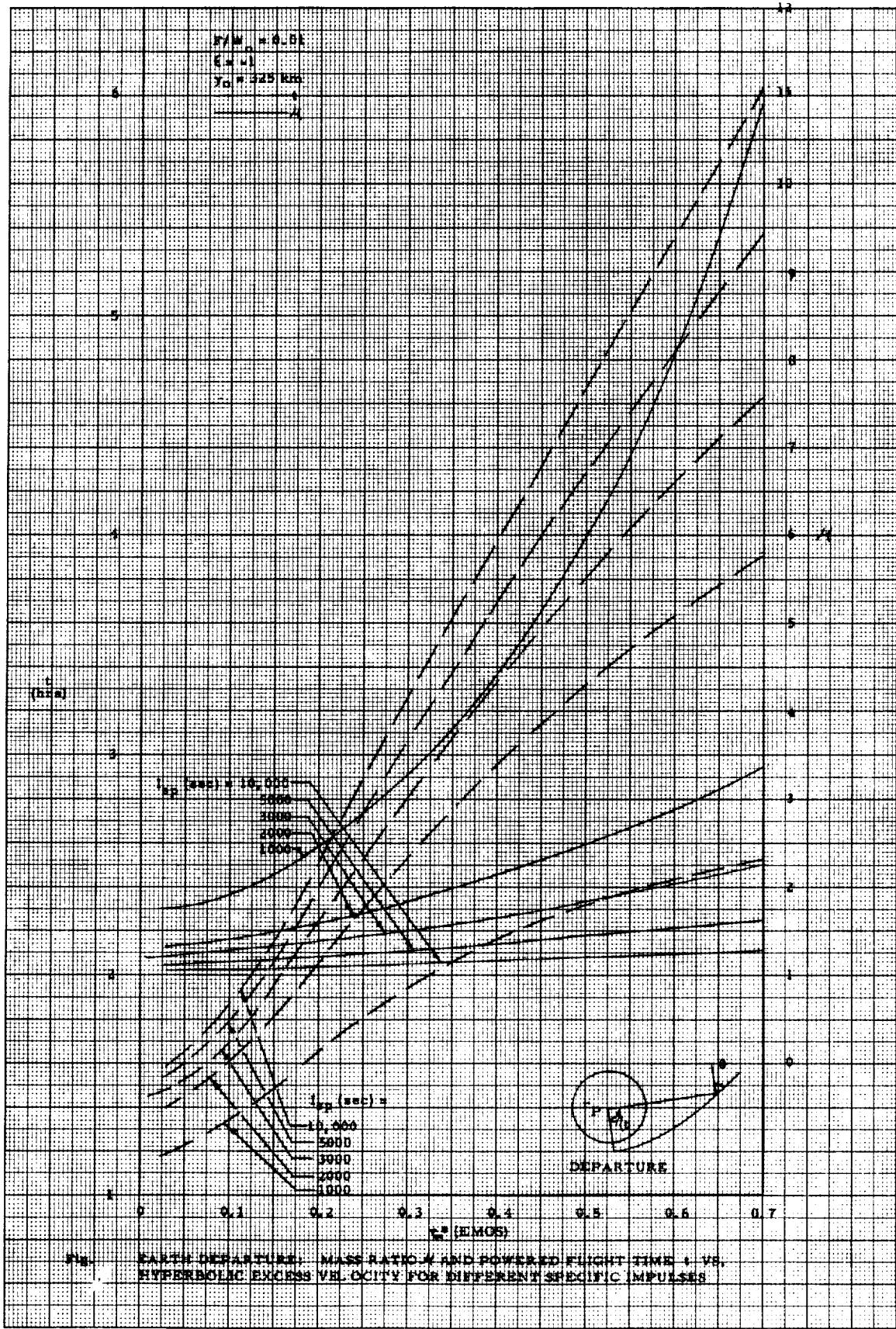
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 AND ELLIPTIC ORBITS AROUND VENUS (PERIAPSIS DISTANCE: $r^*=5.0$)



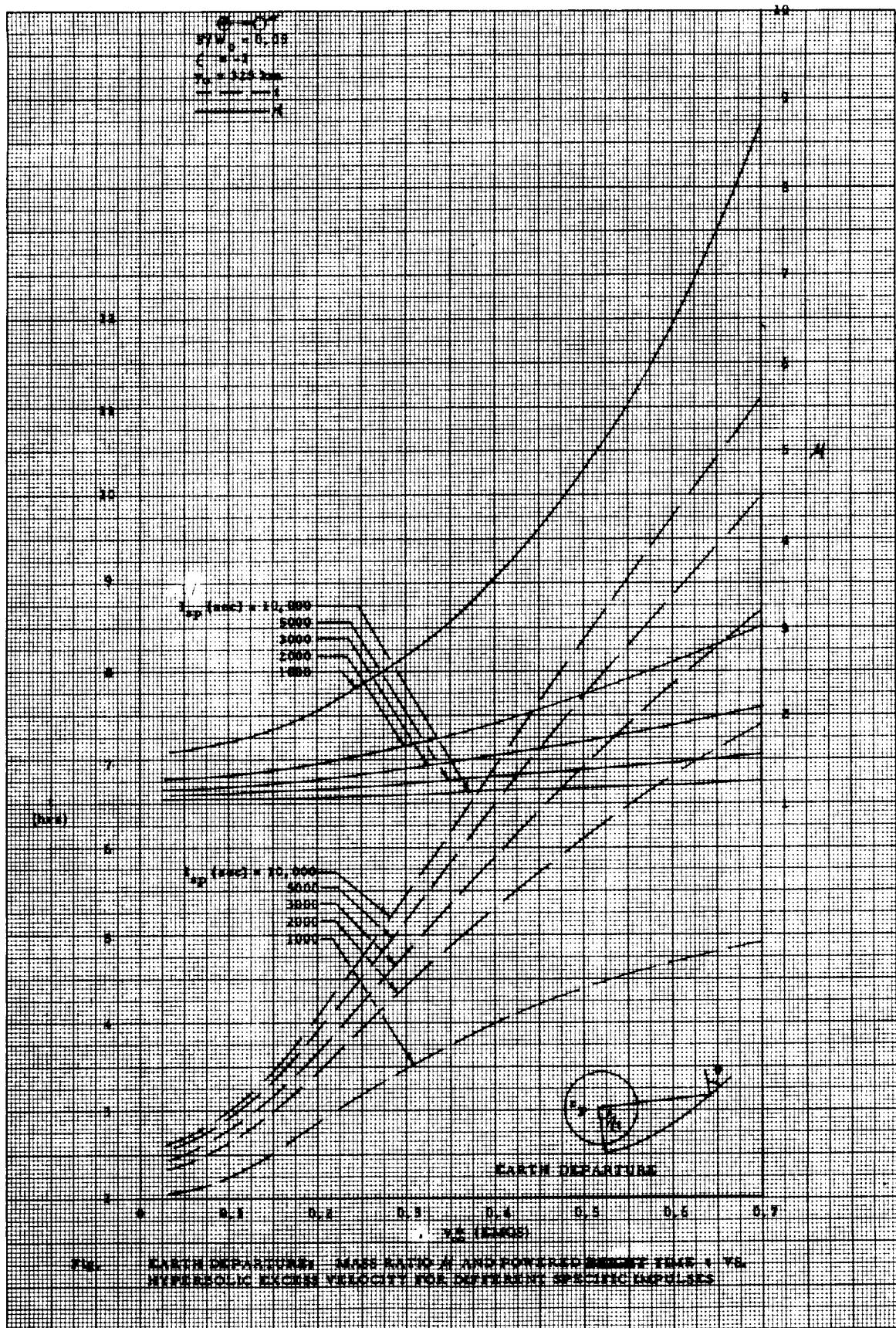
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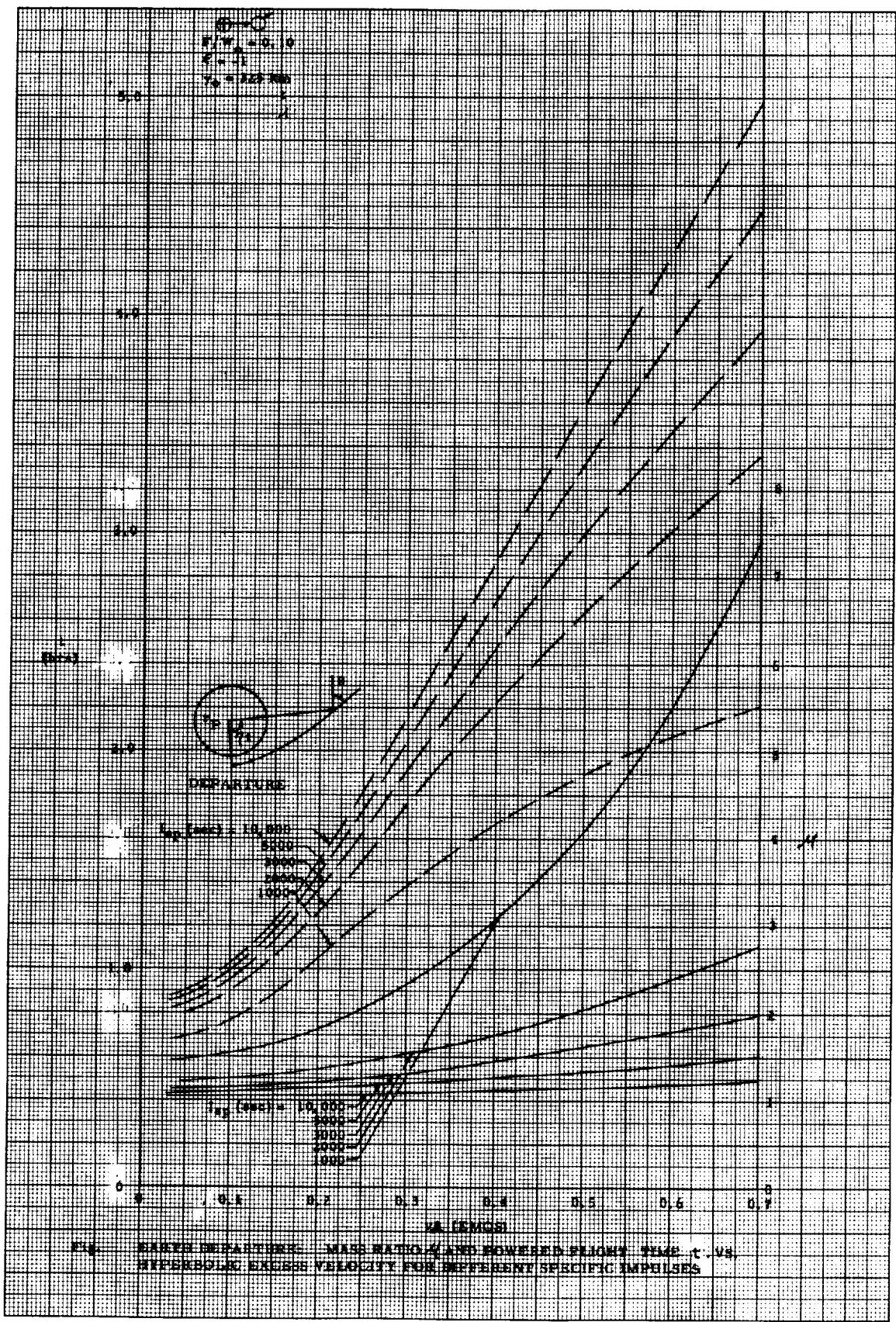


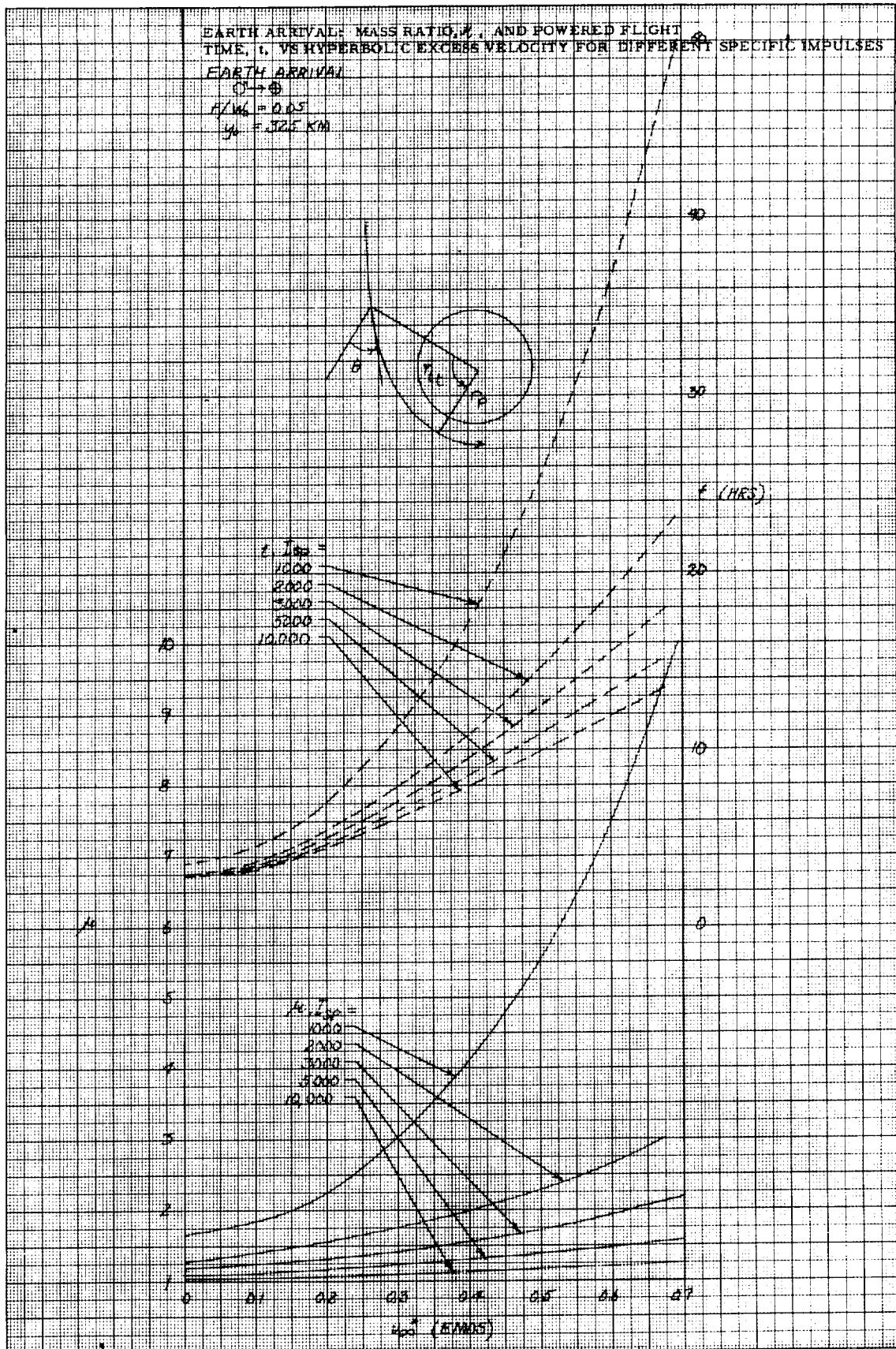




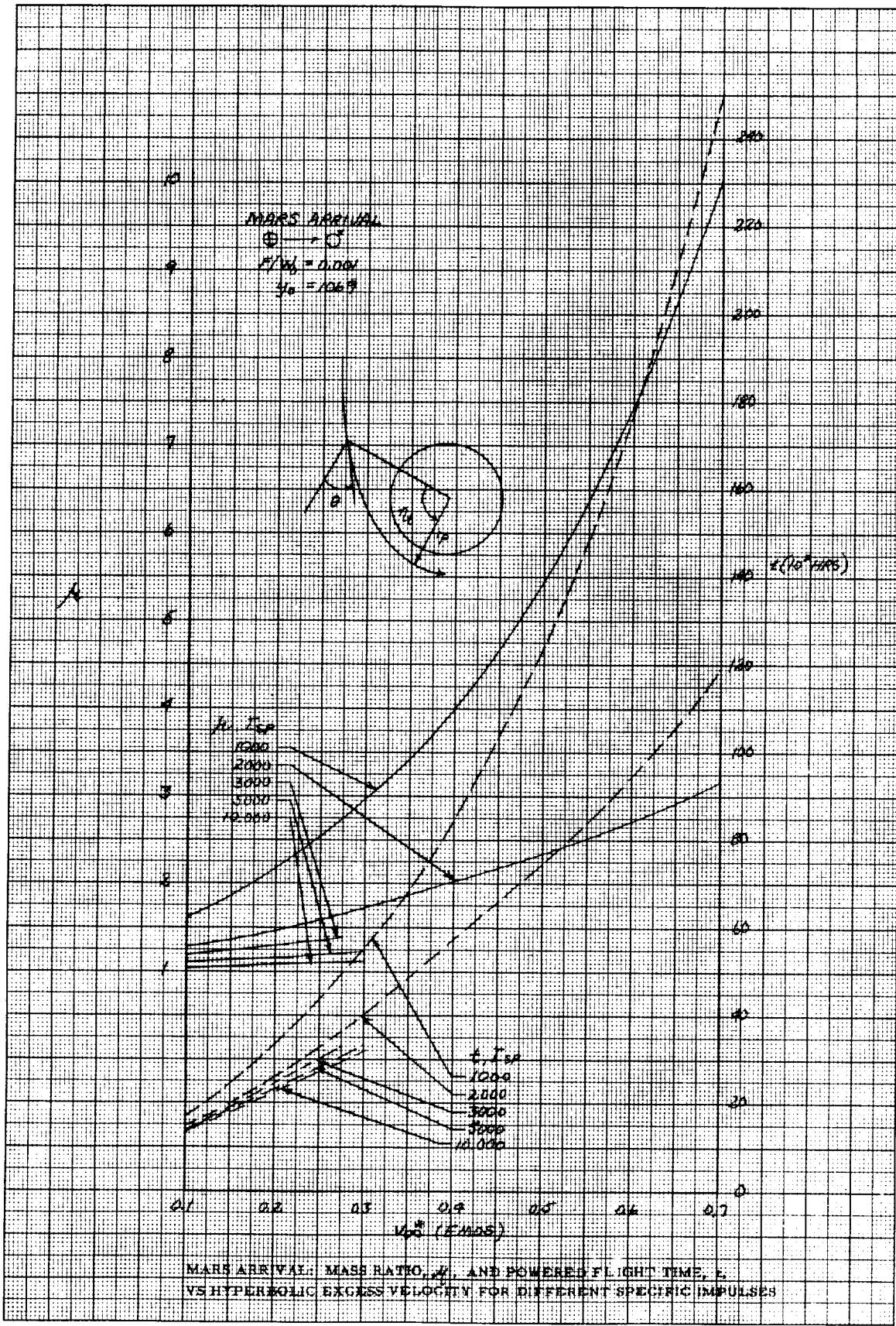
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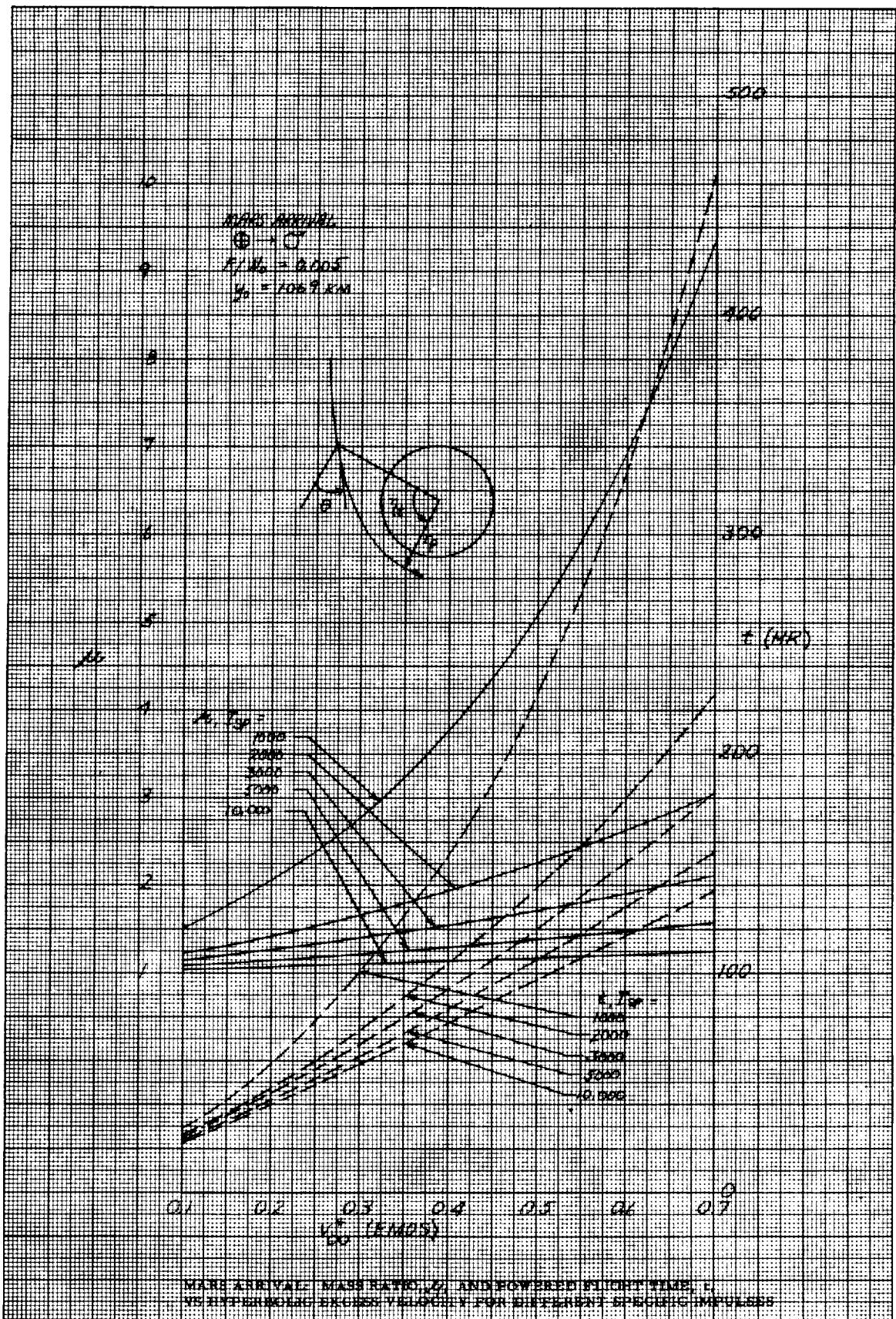


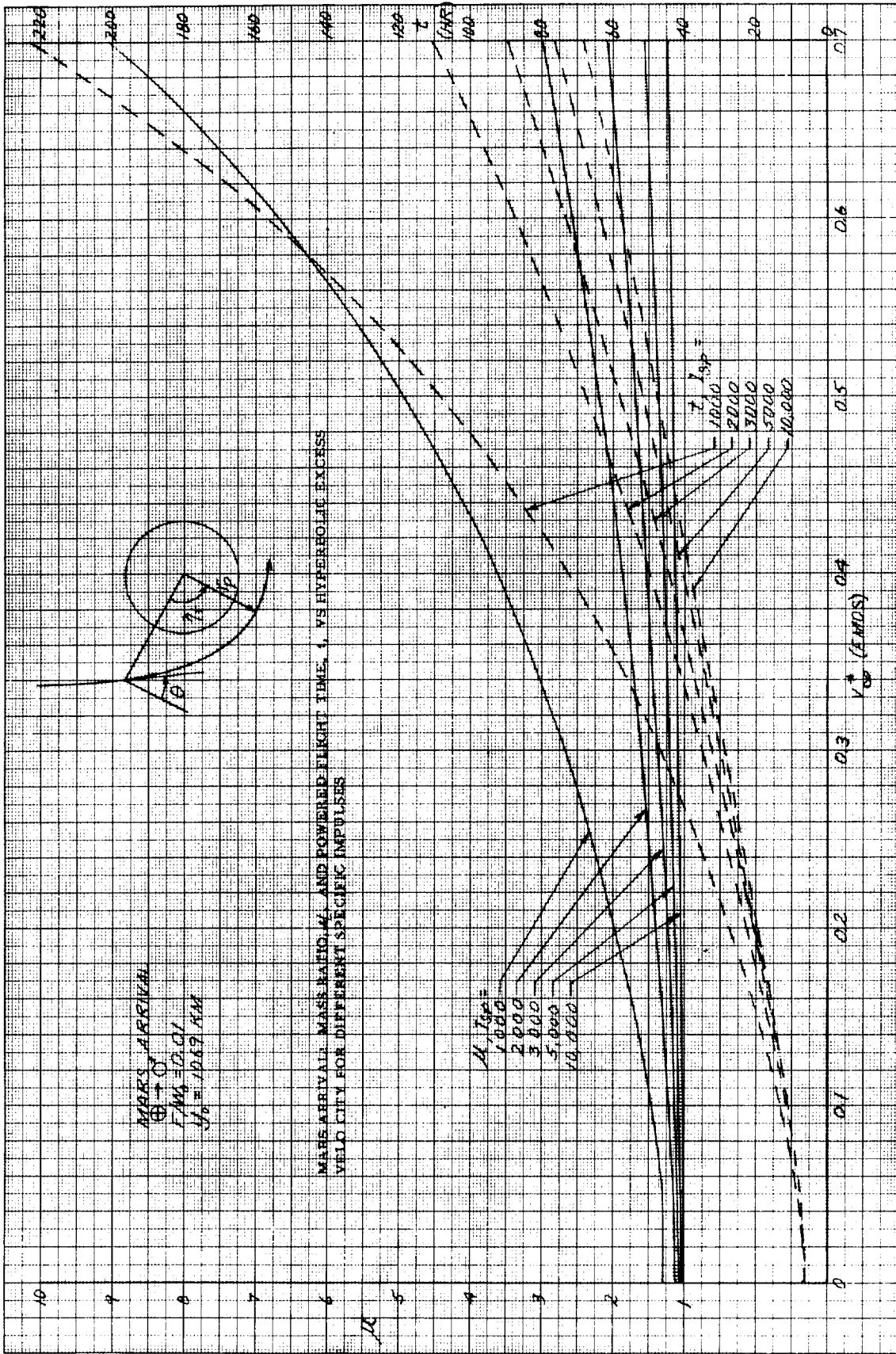


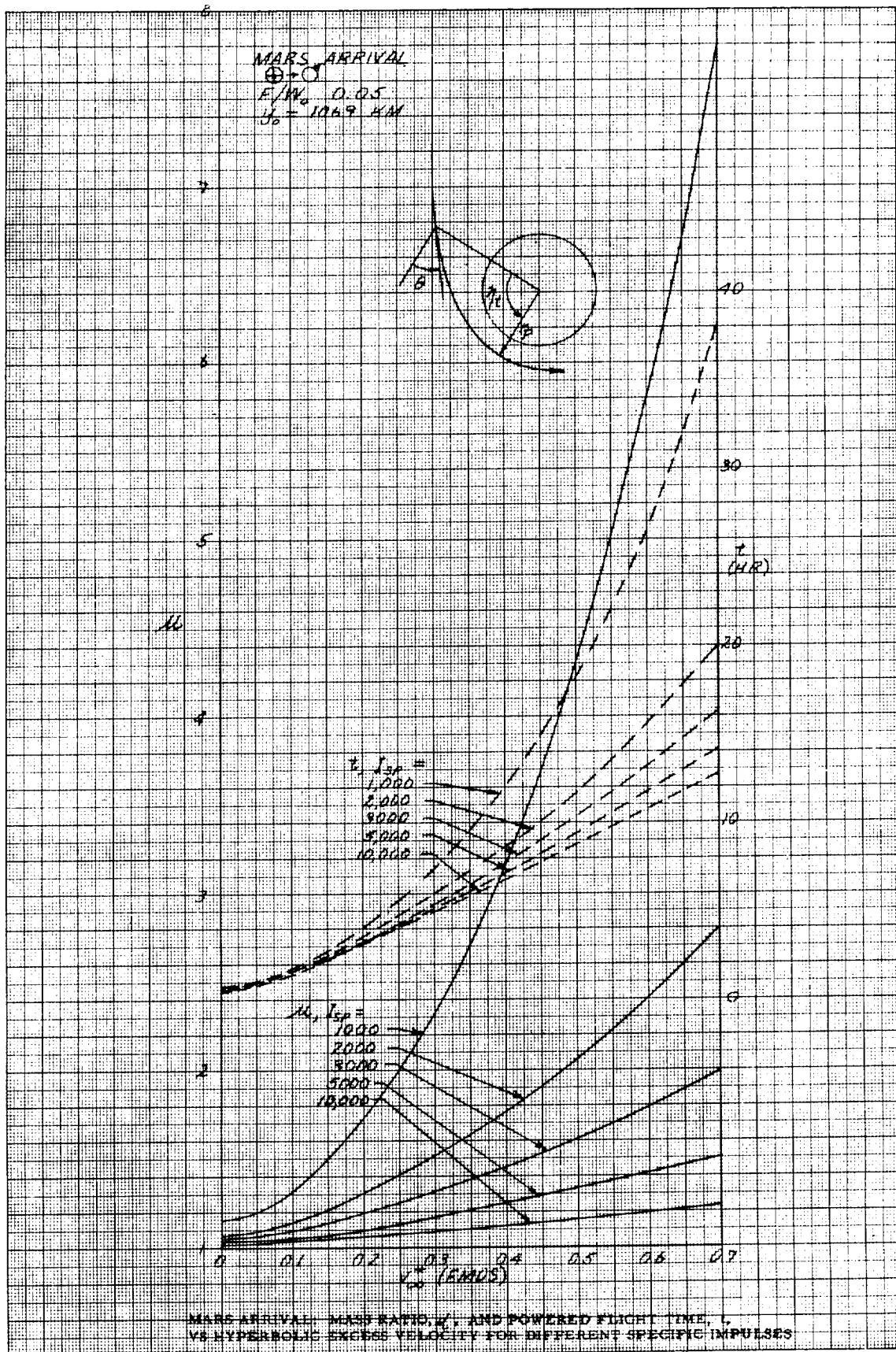
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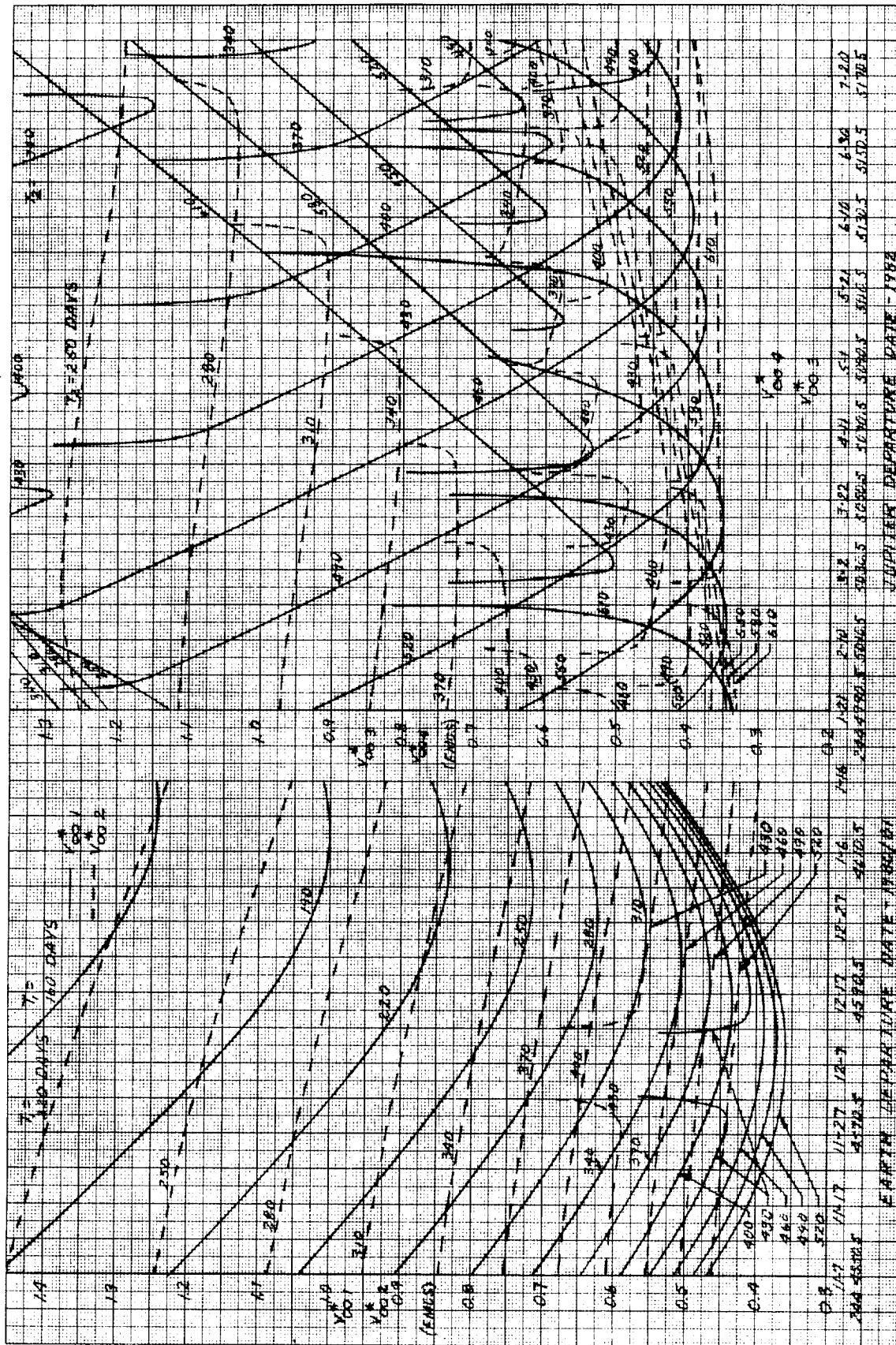


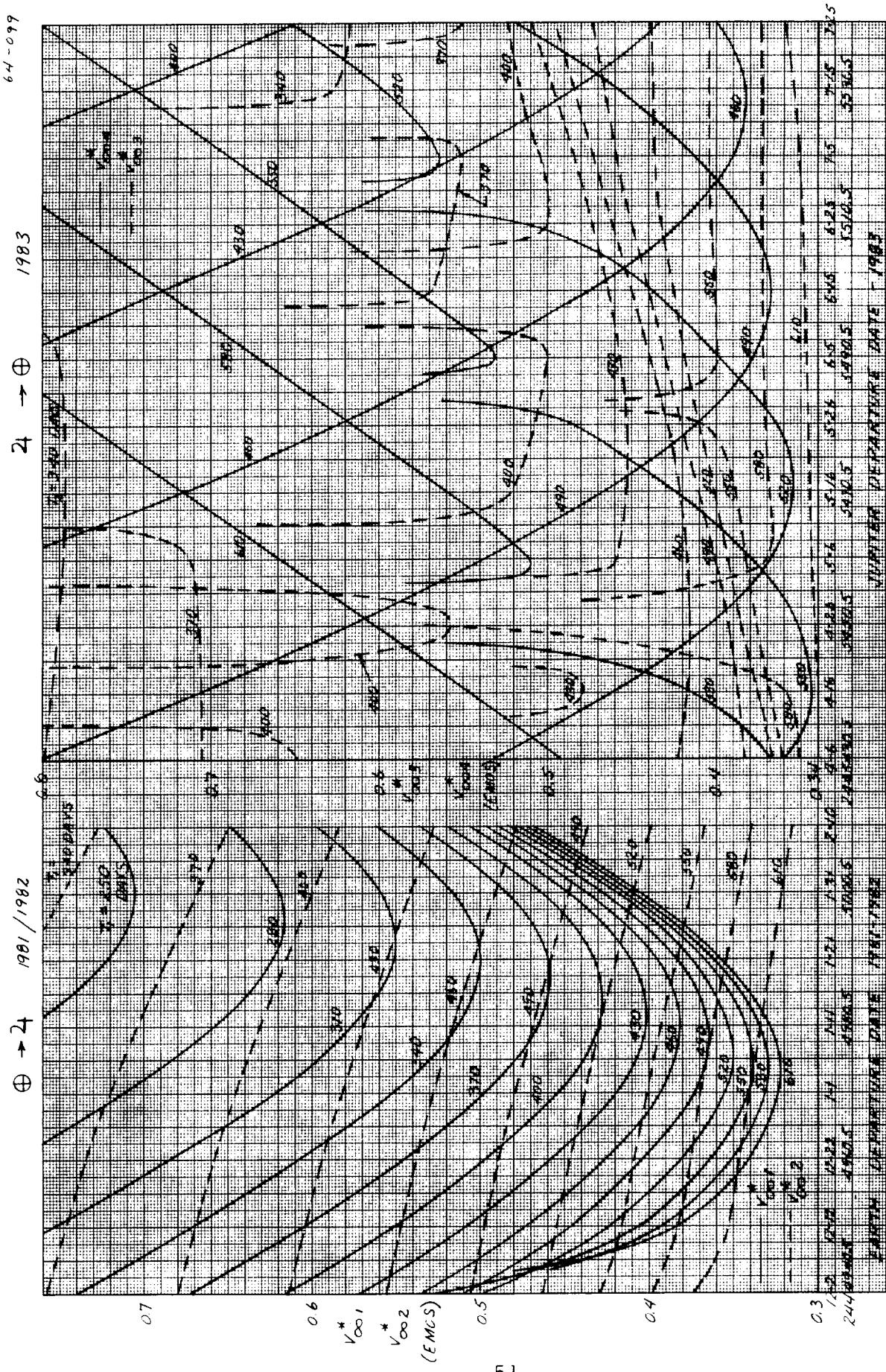


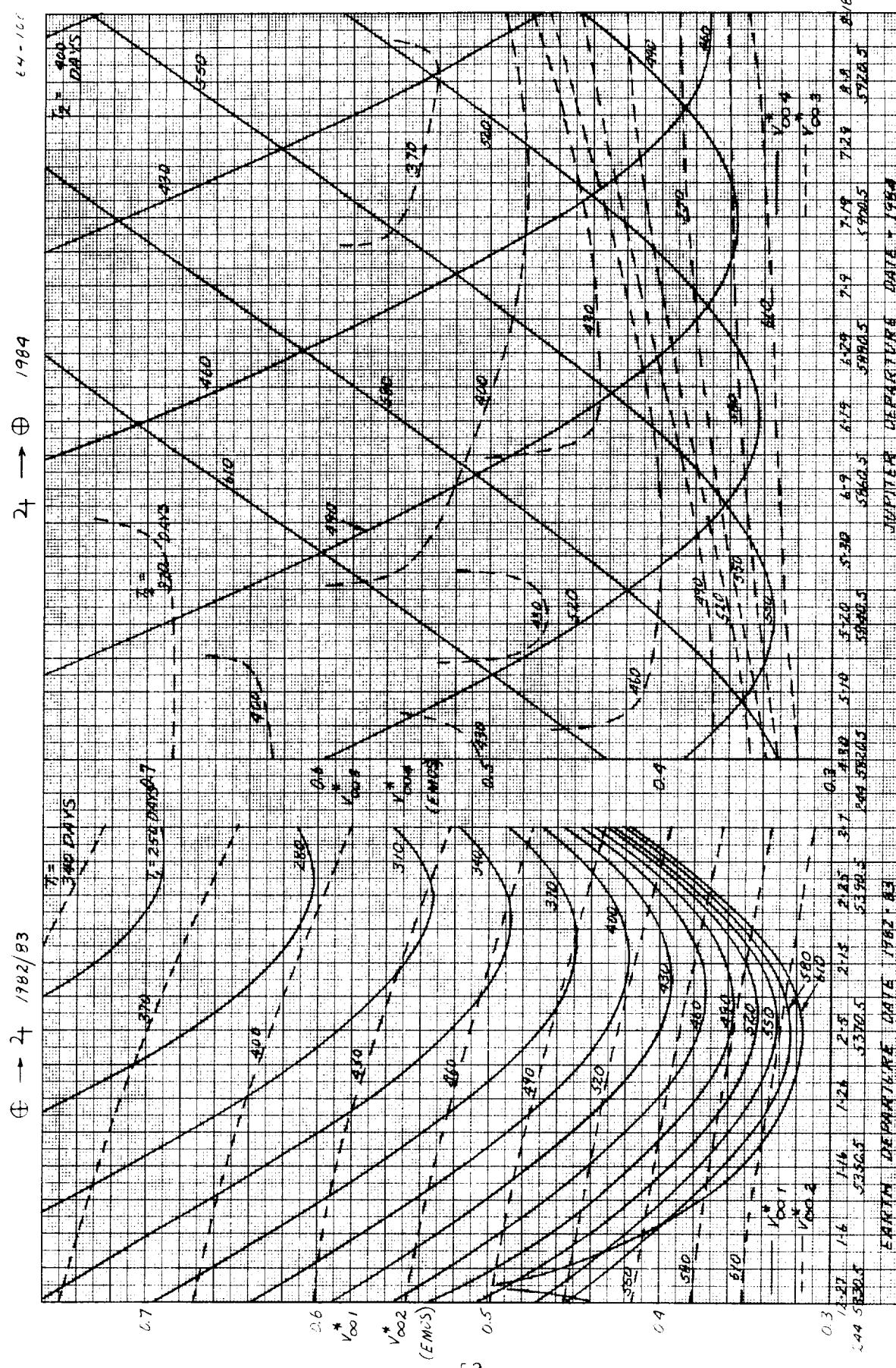
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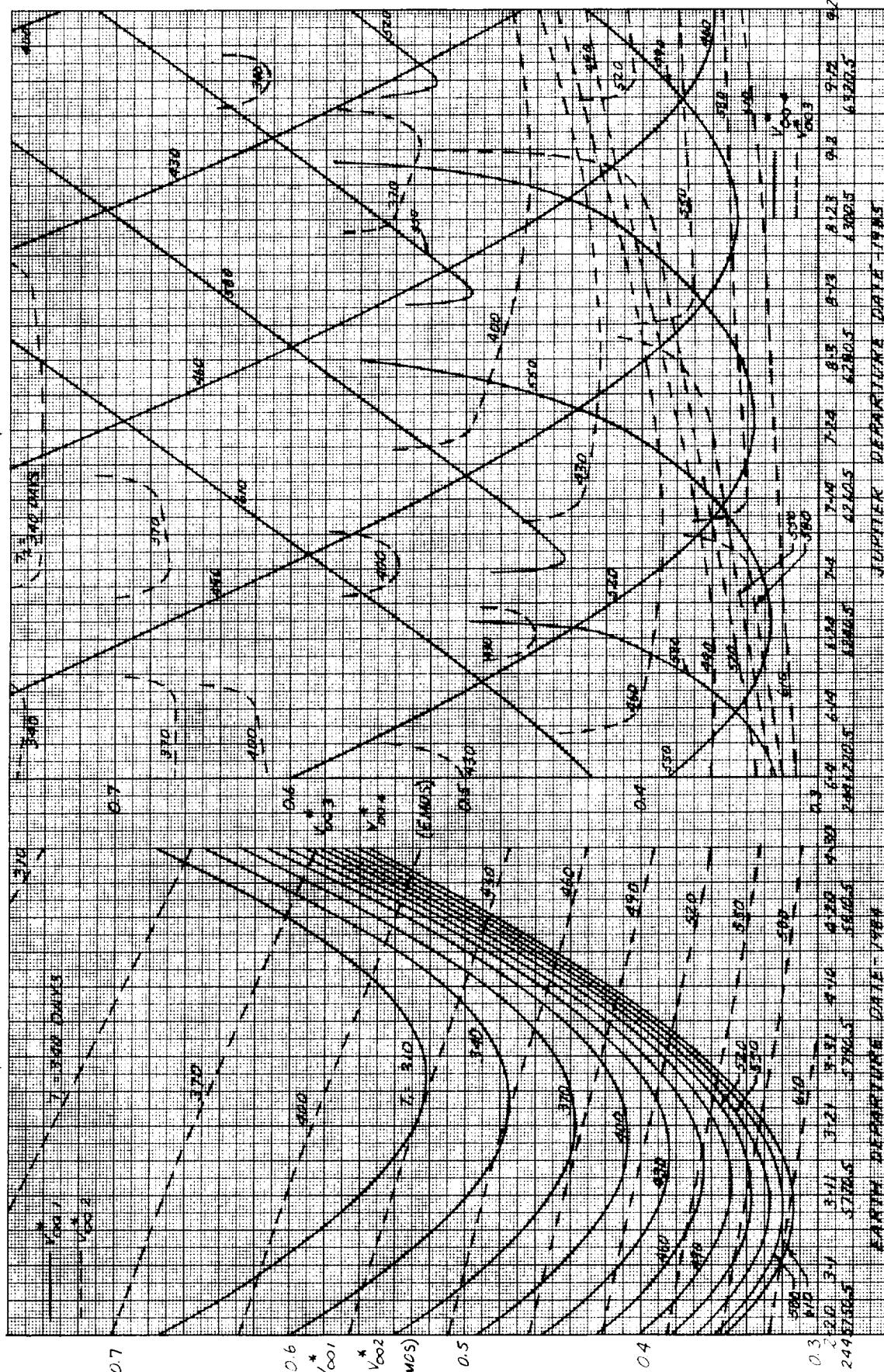


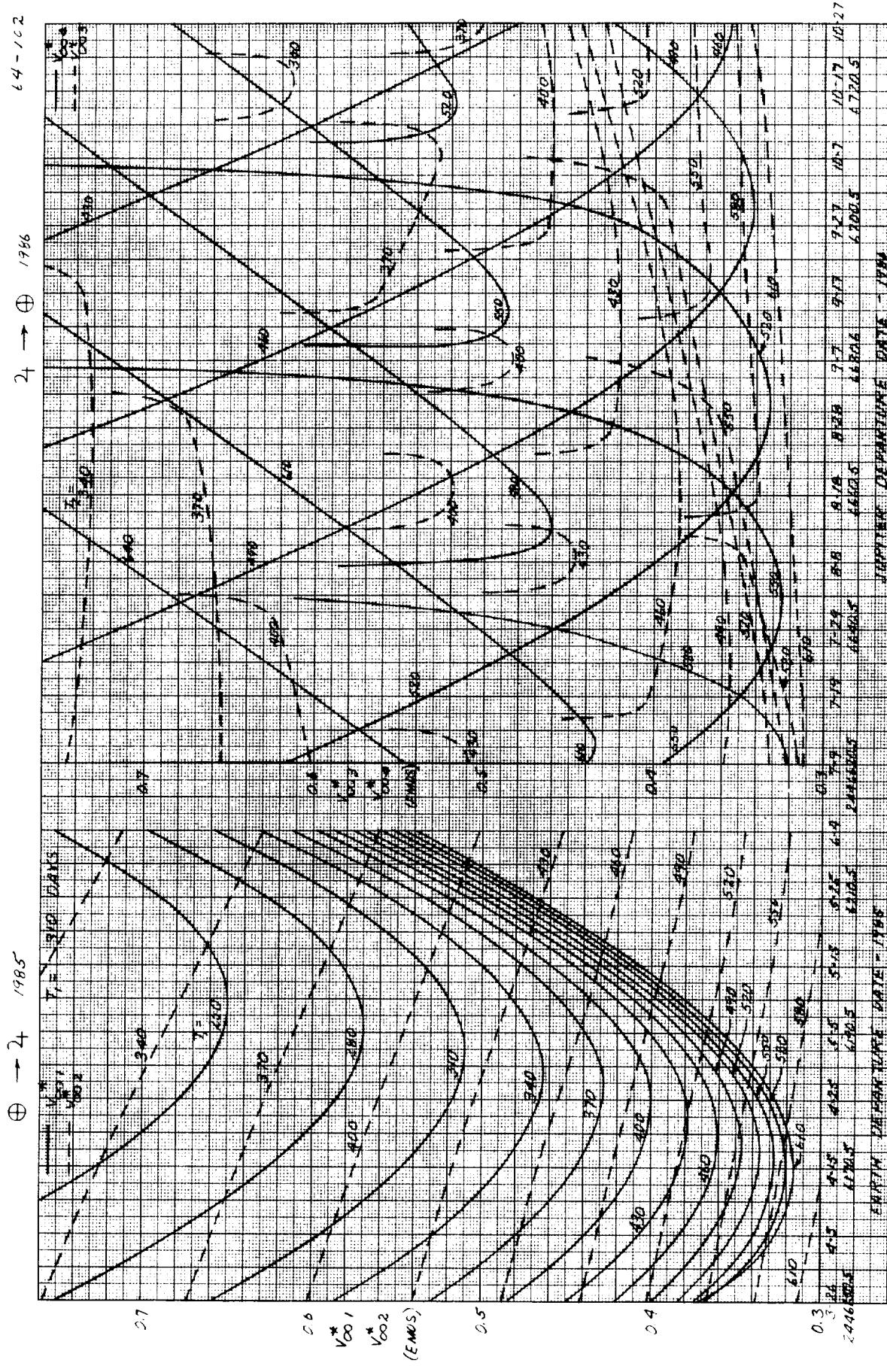


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$2 \rightarrow \oplus$ 1985

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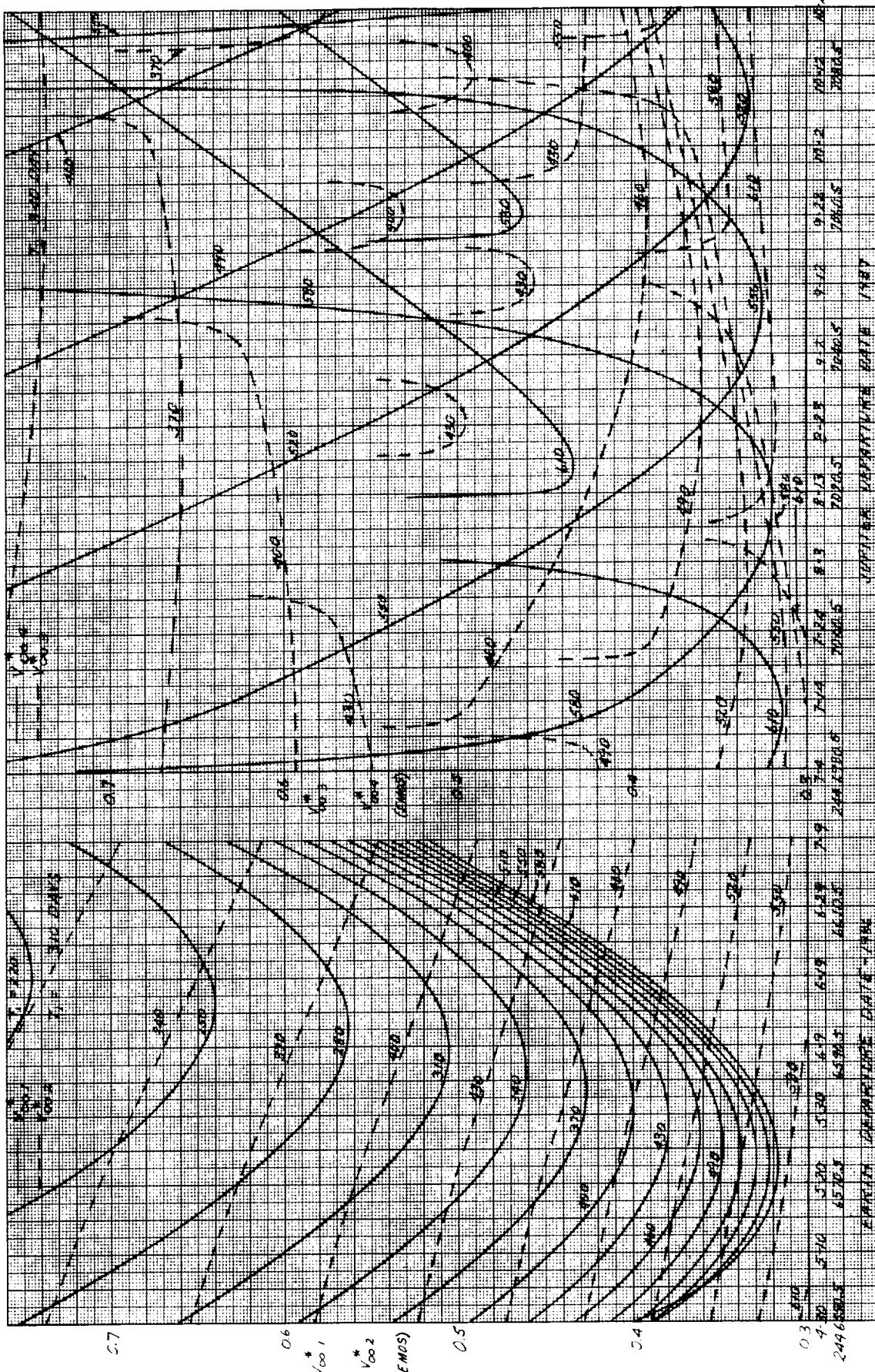


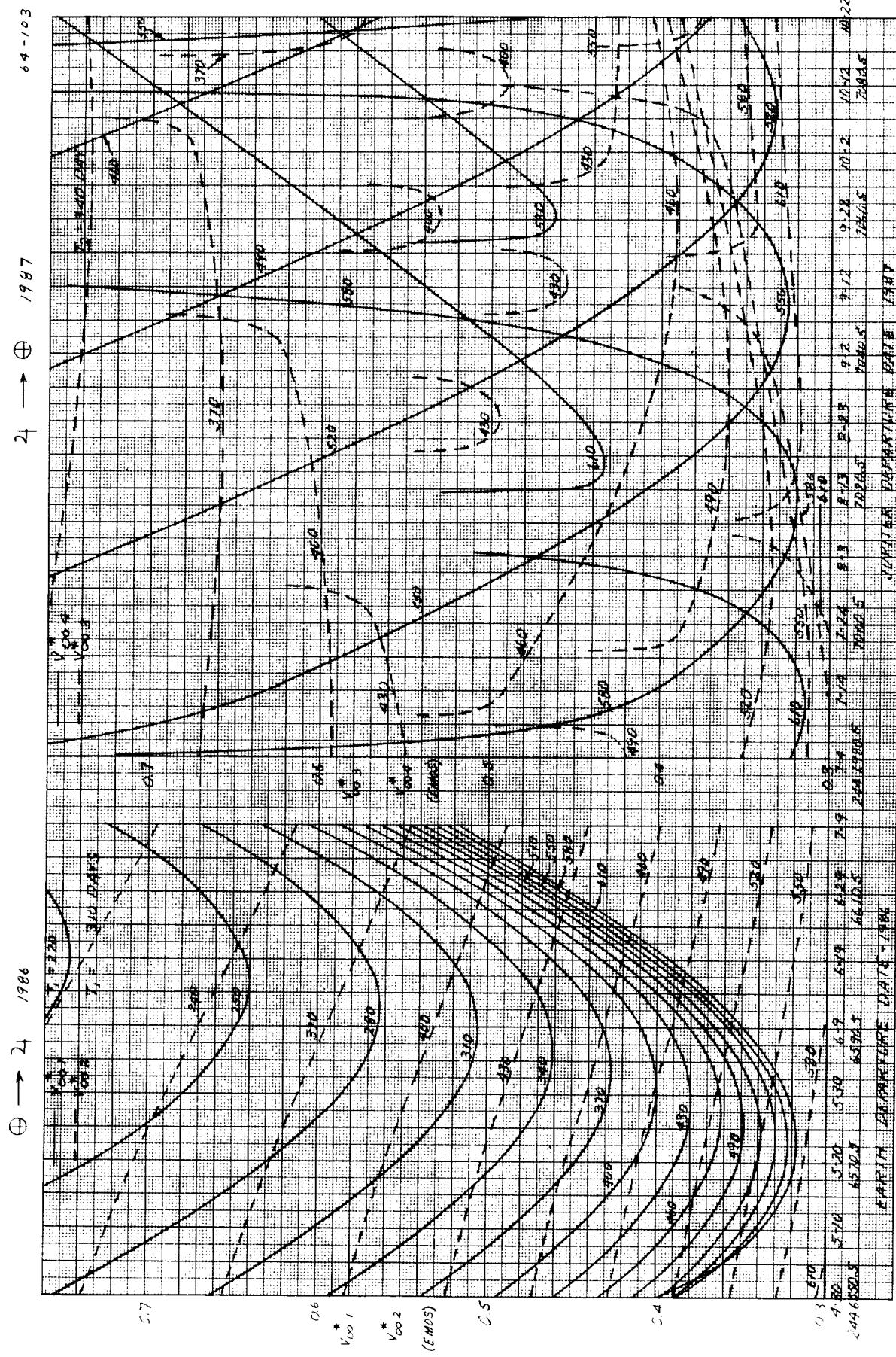


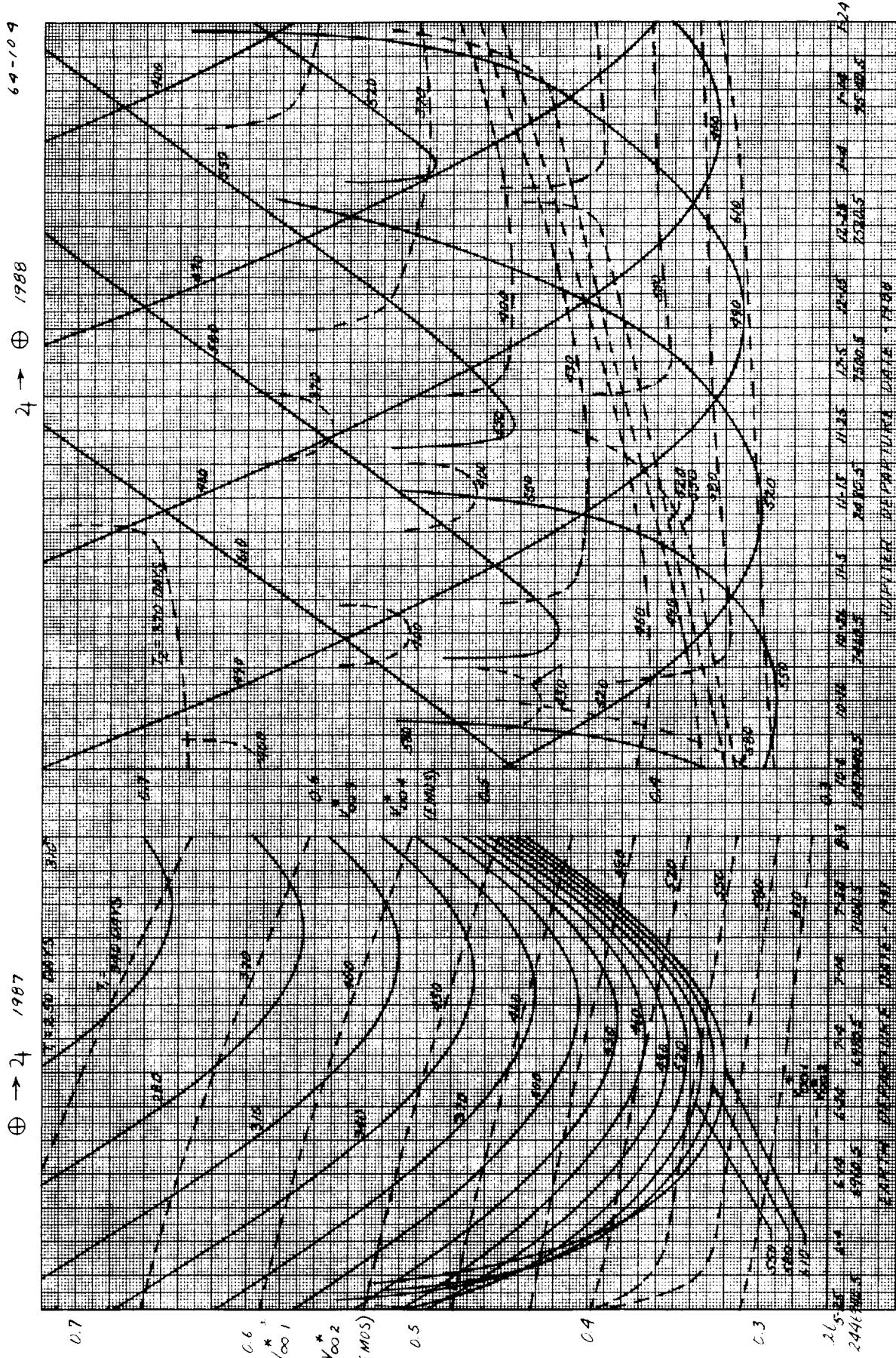
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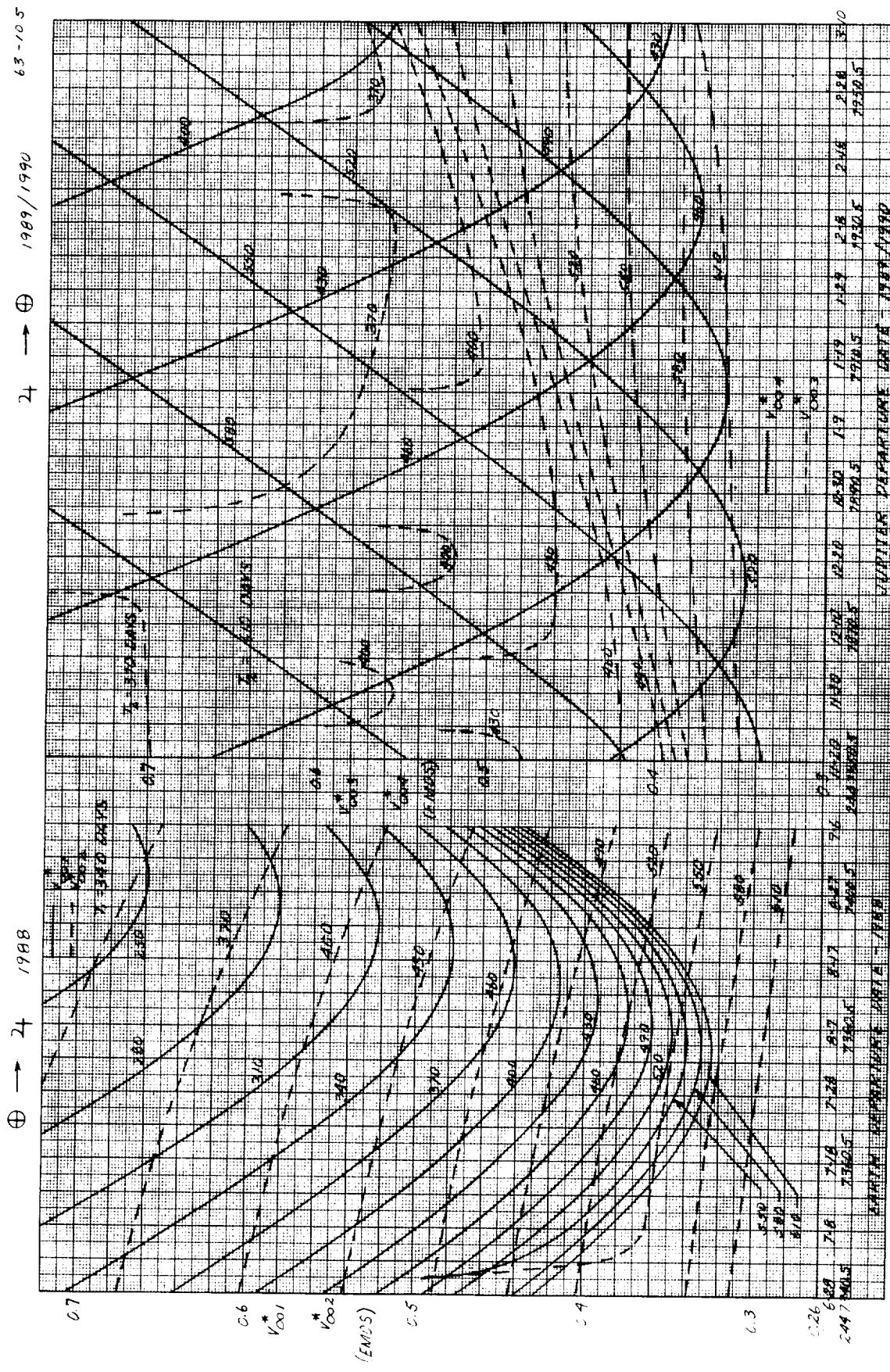
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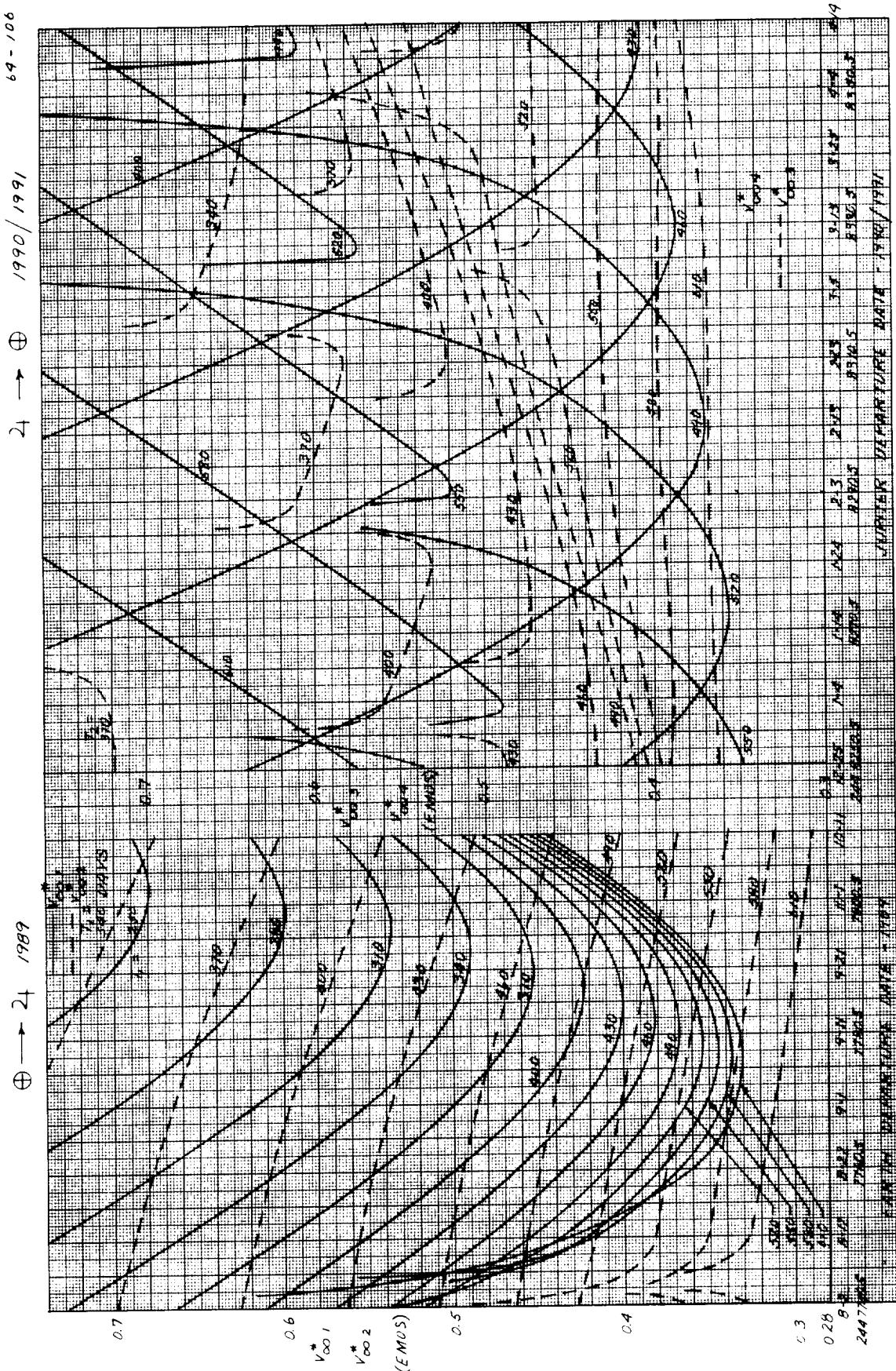
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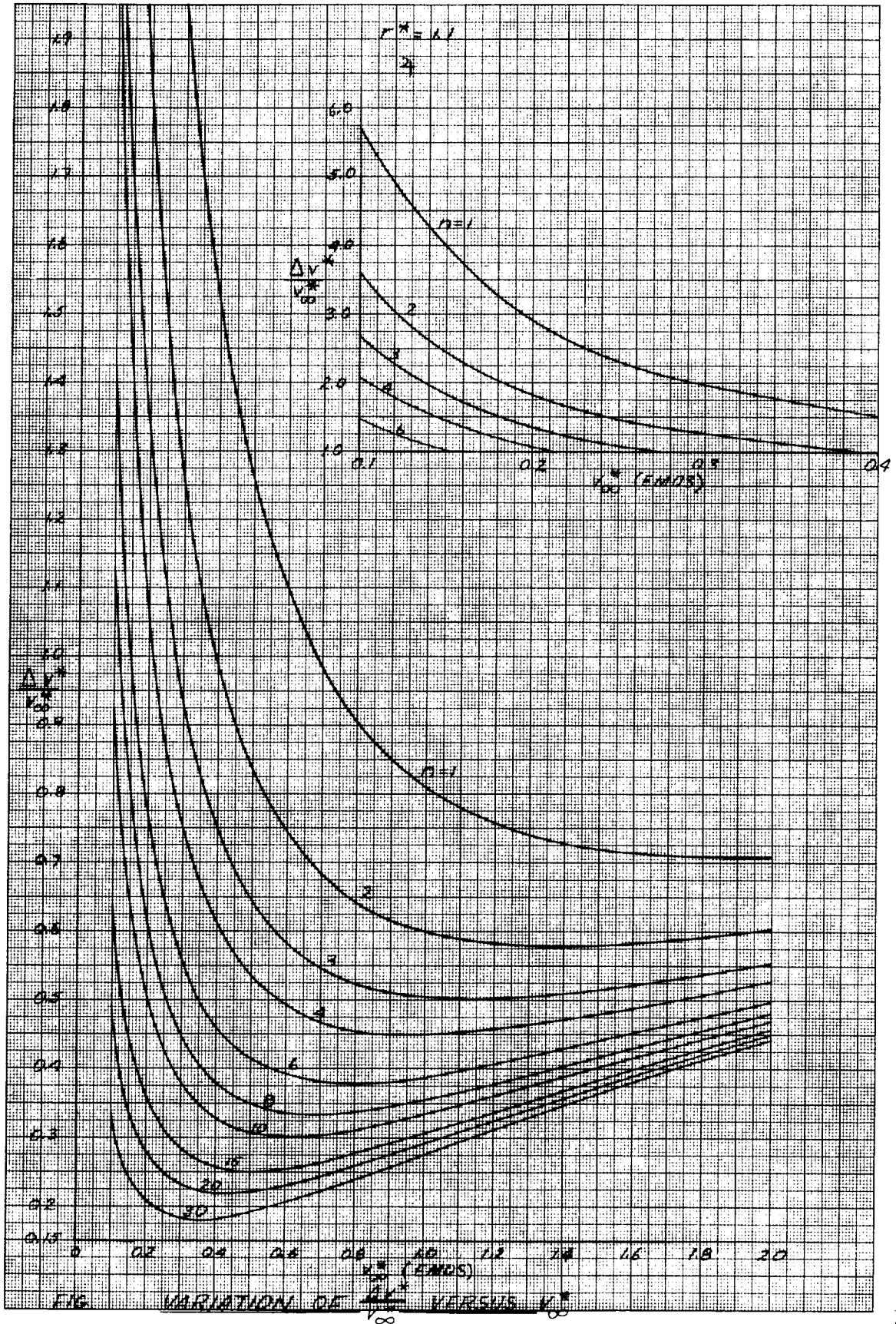


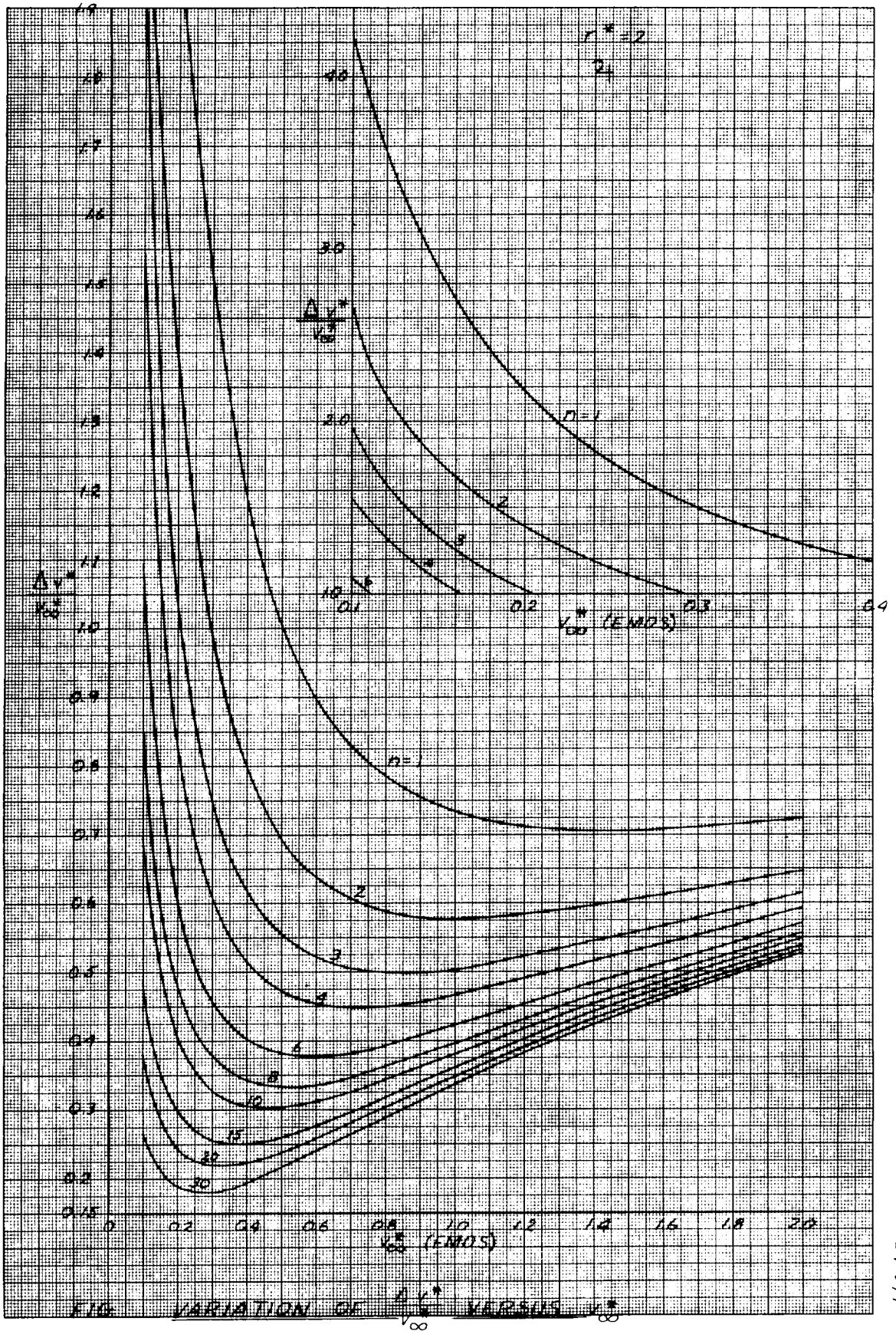


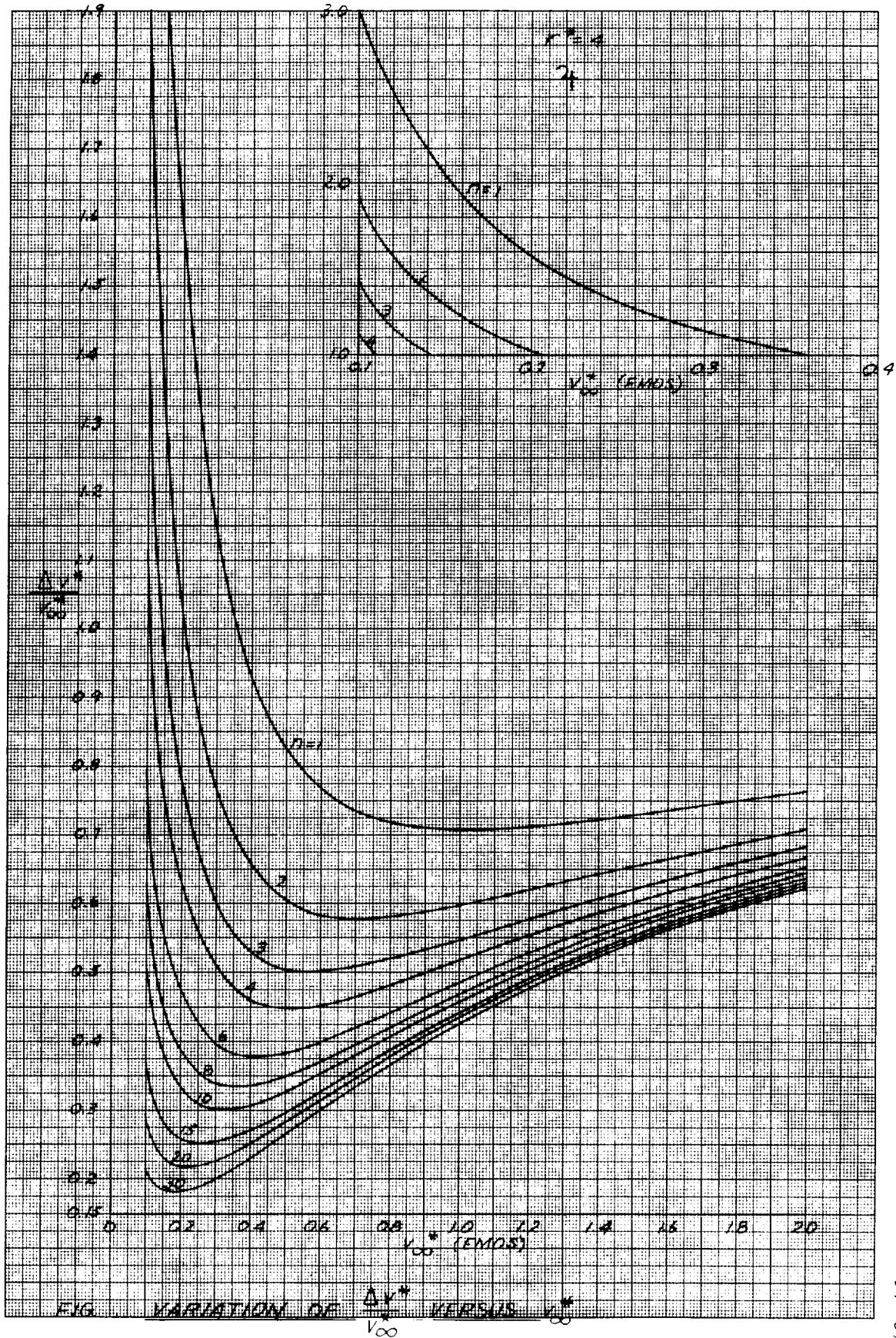




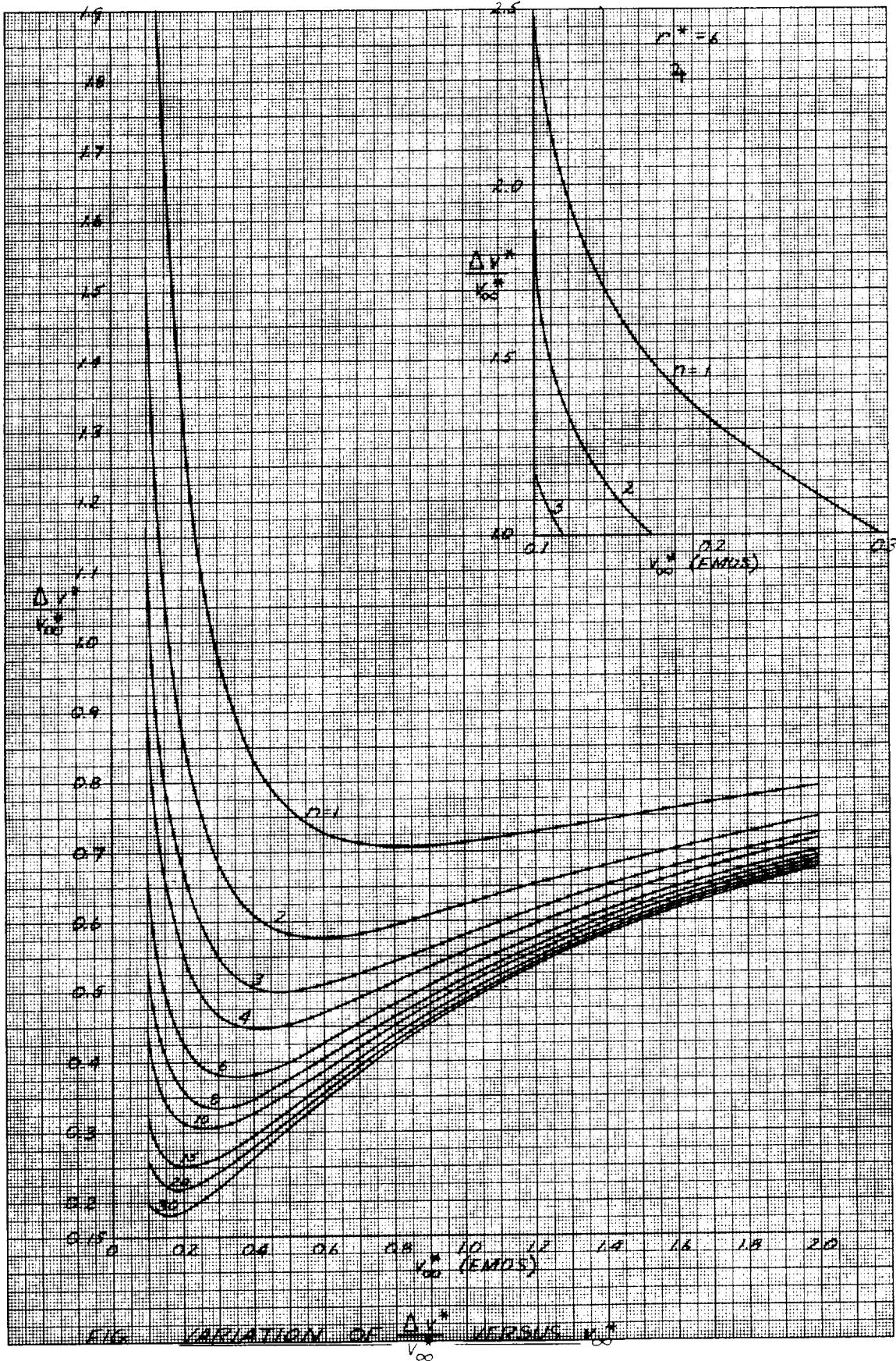




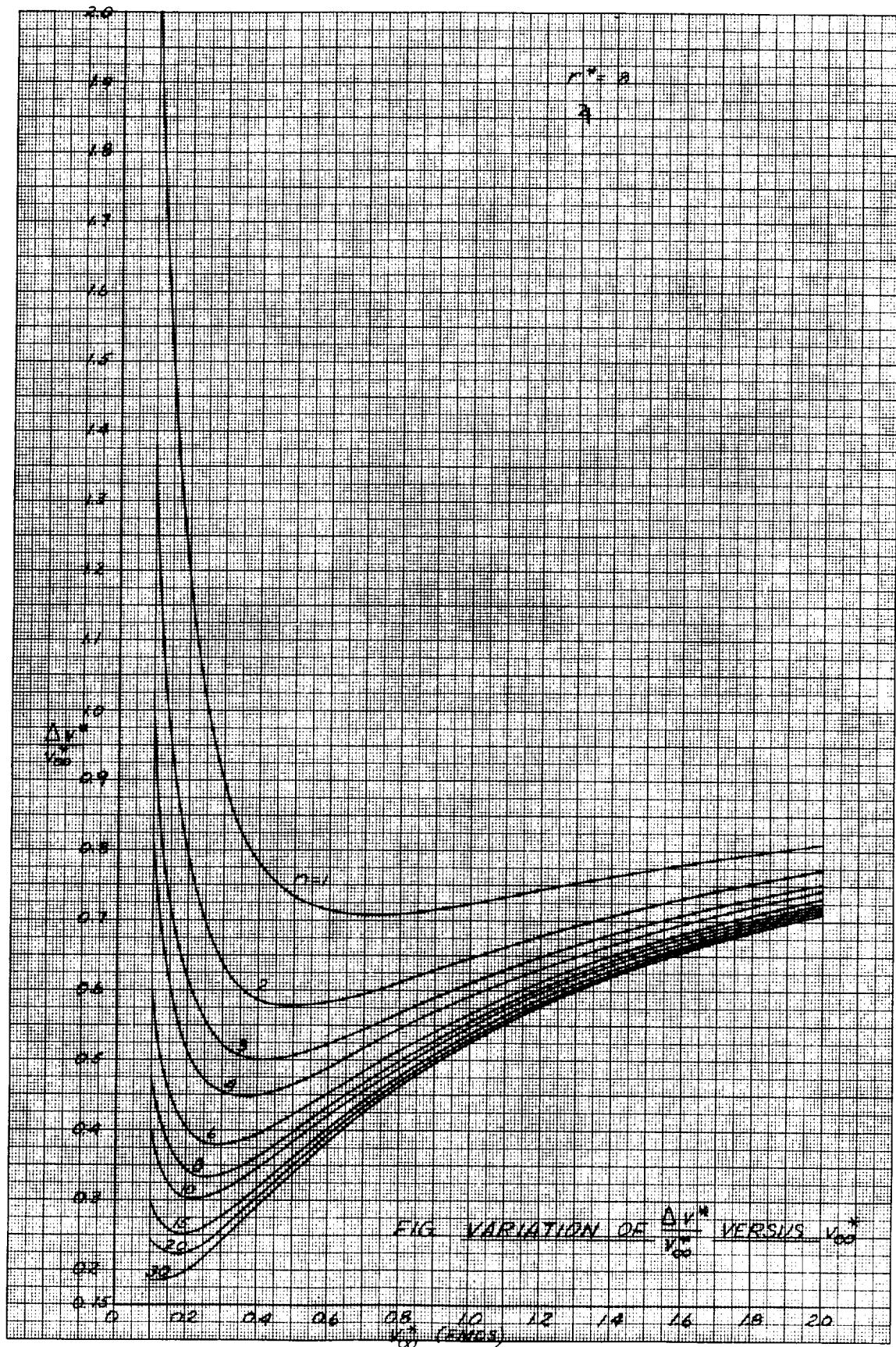




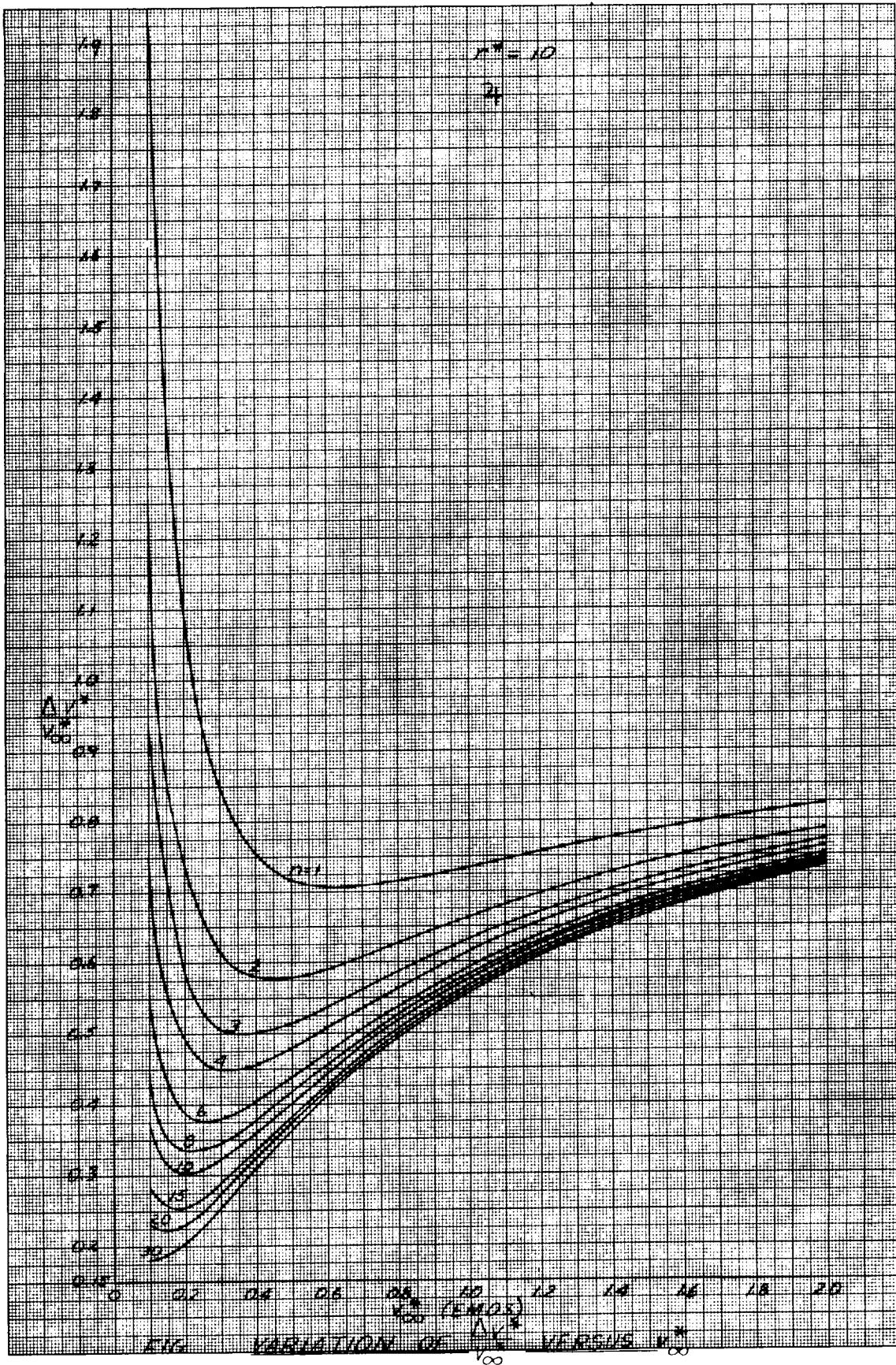
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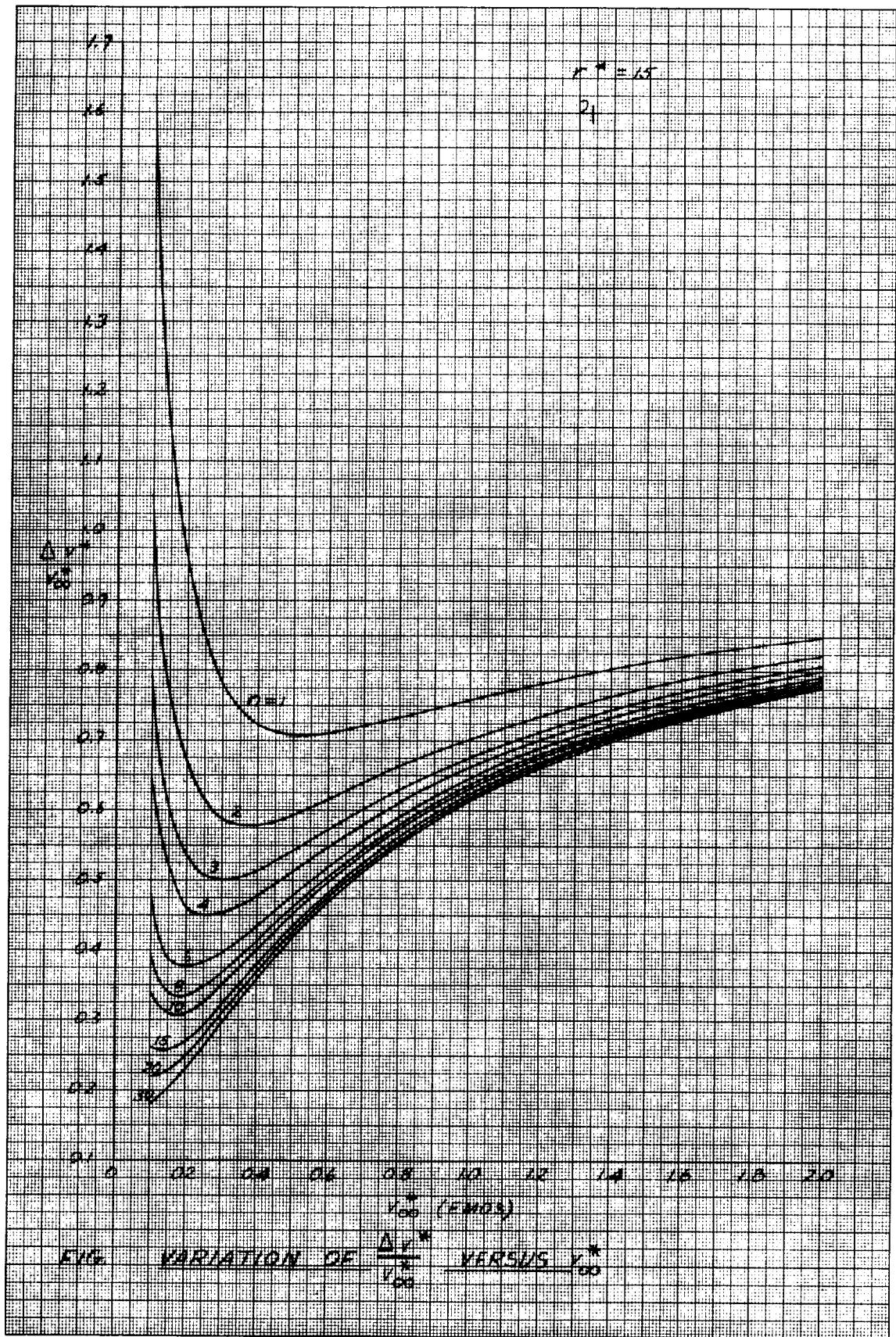


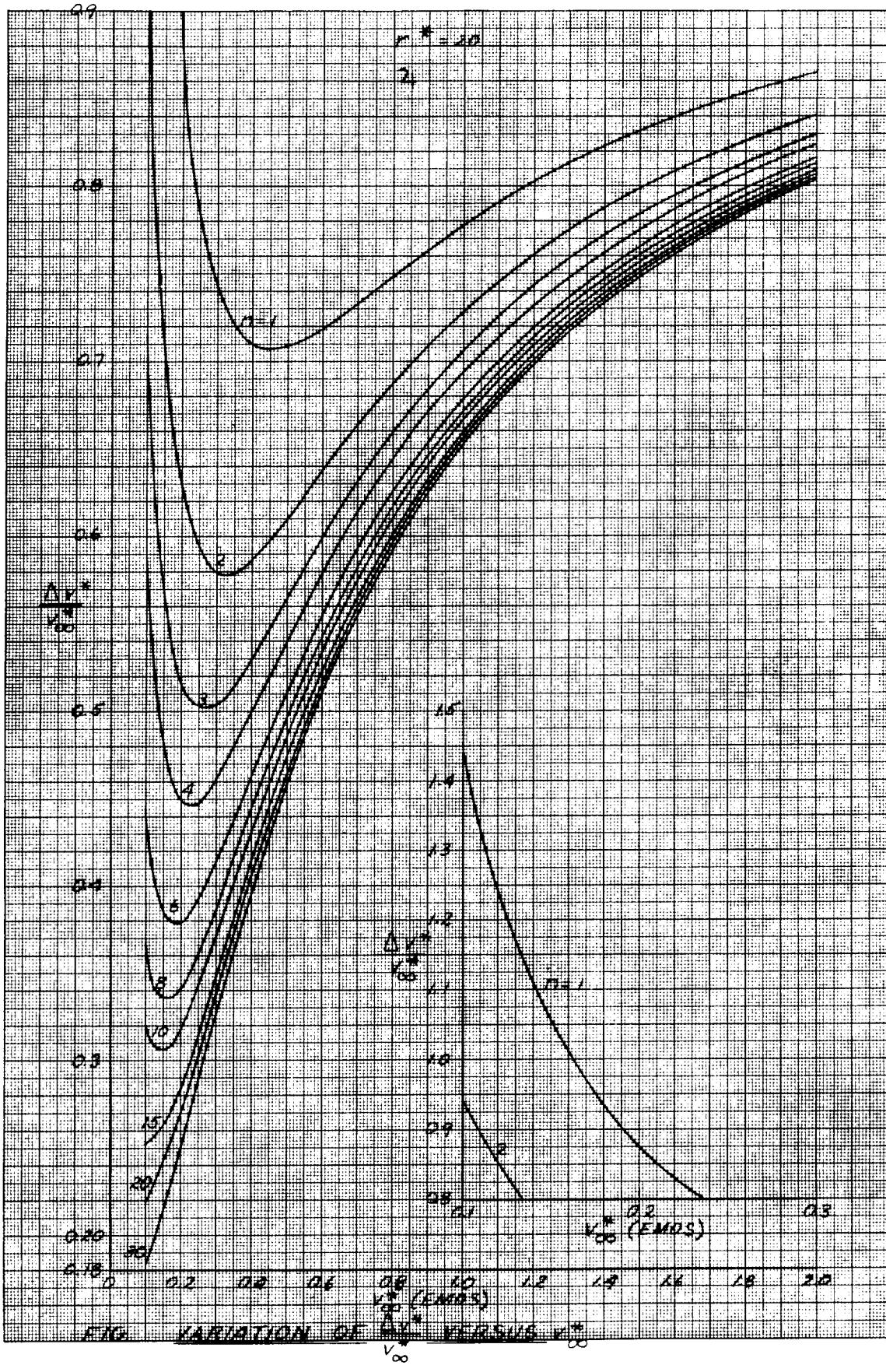
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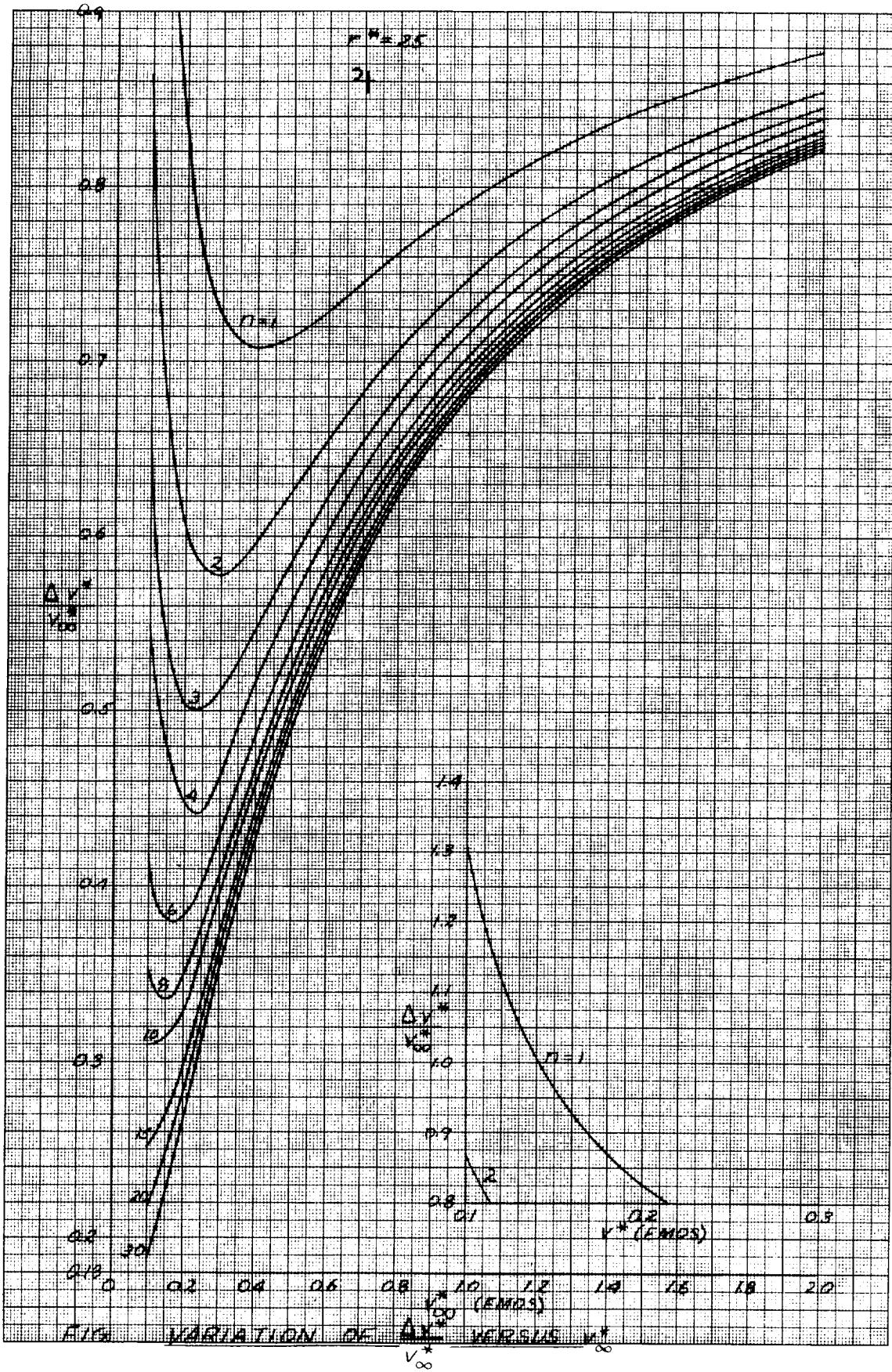


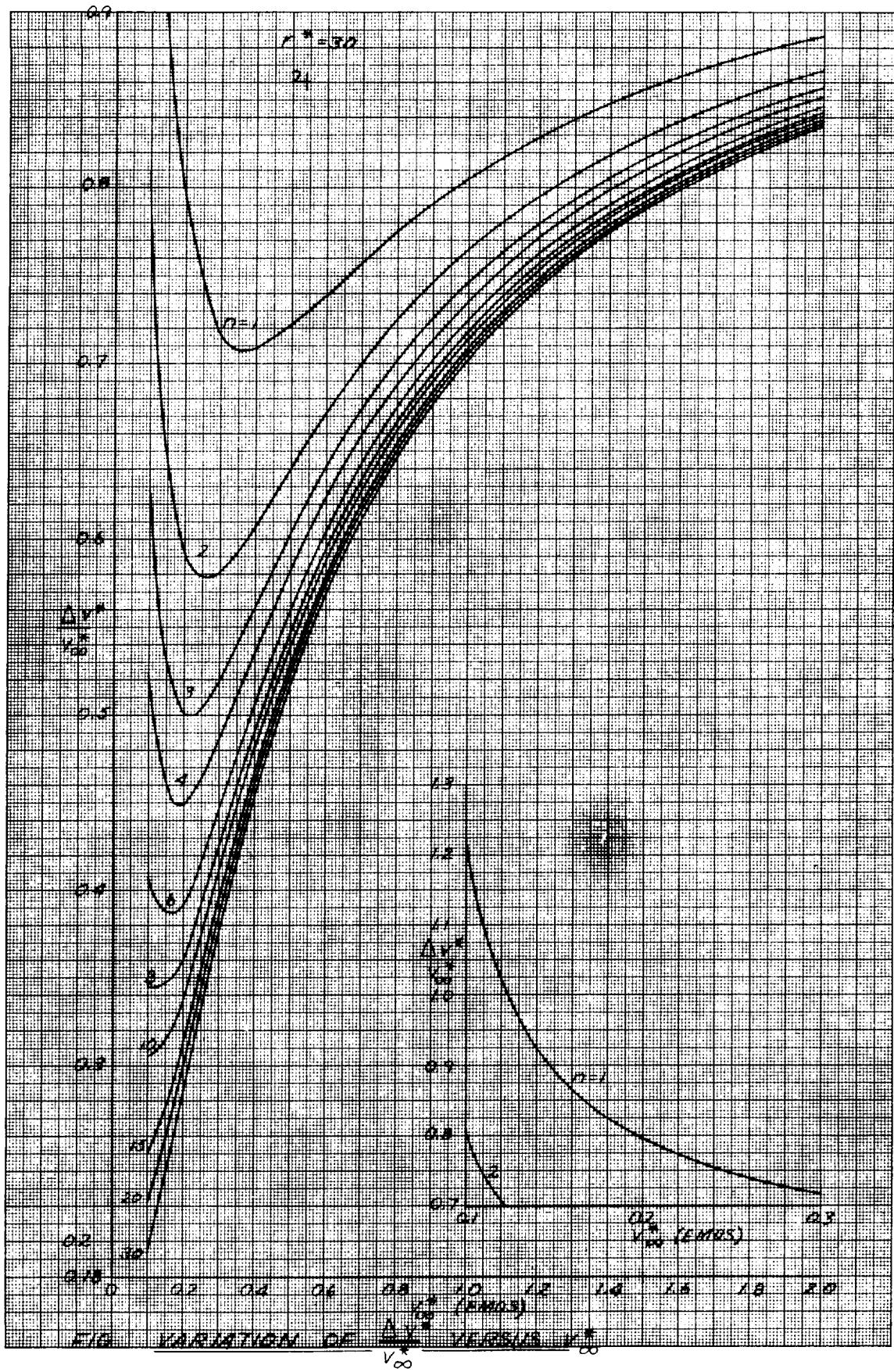
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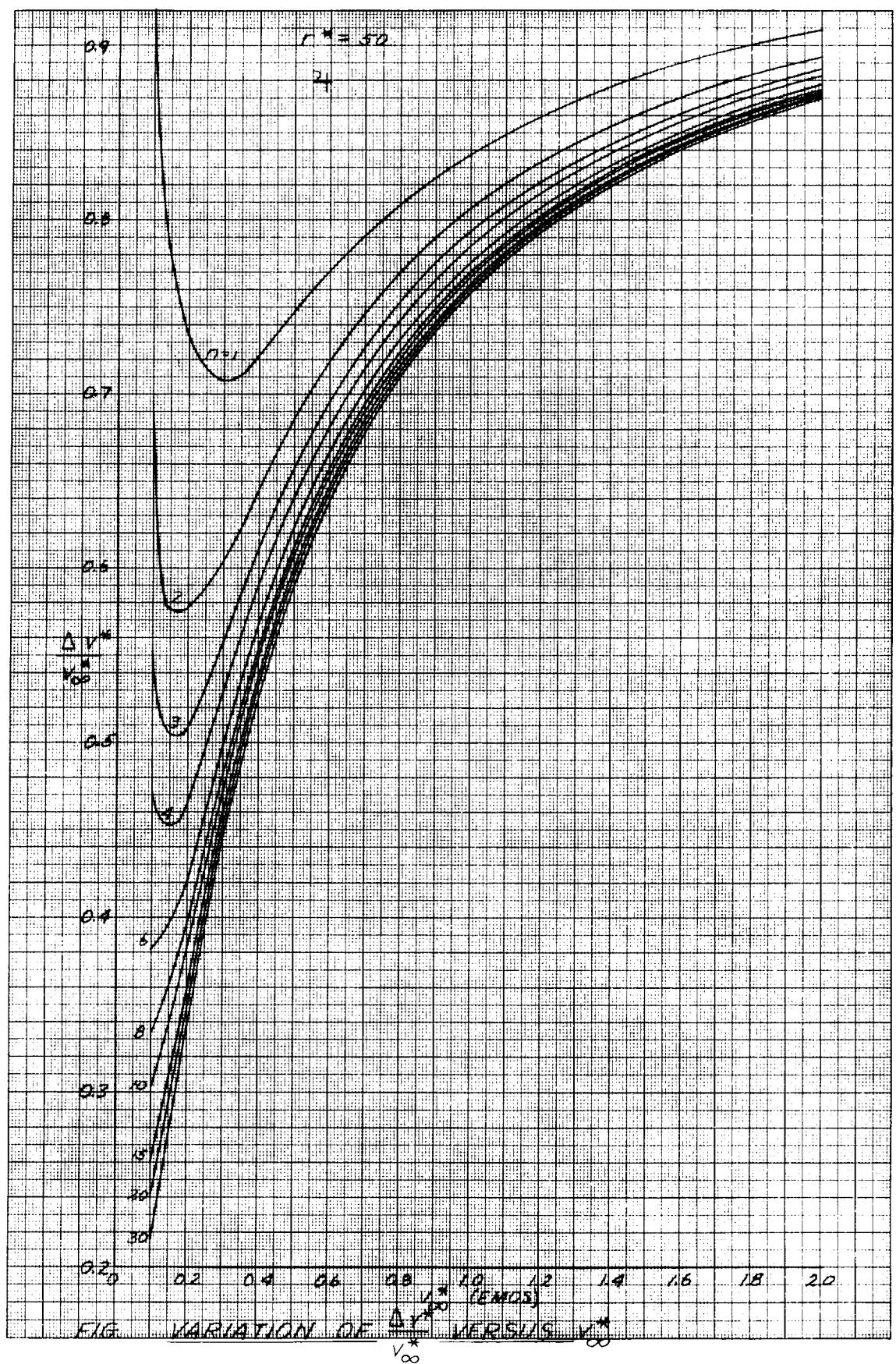


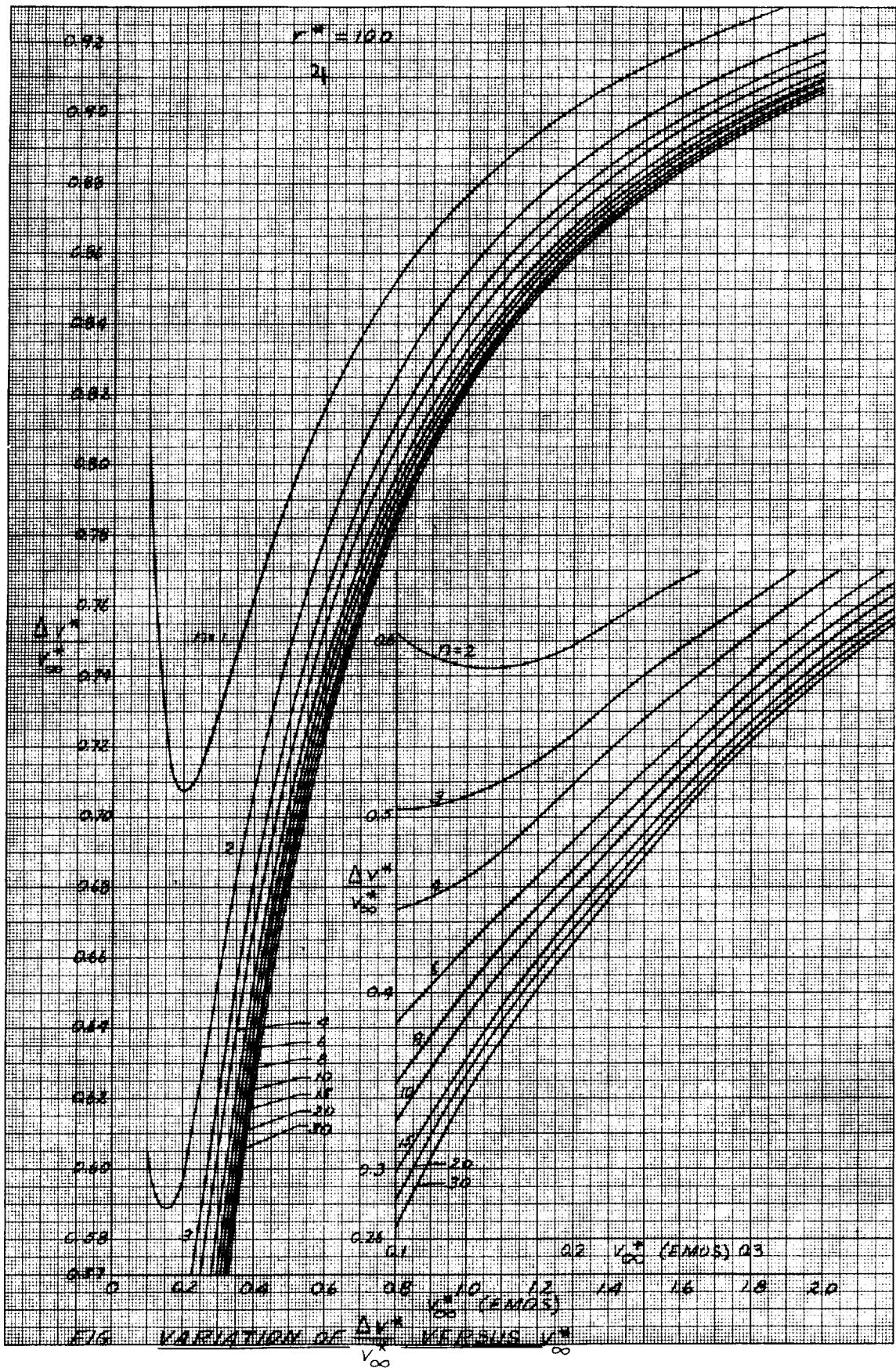




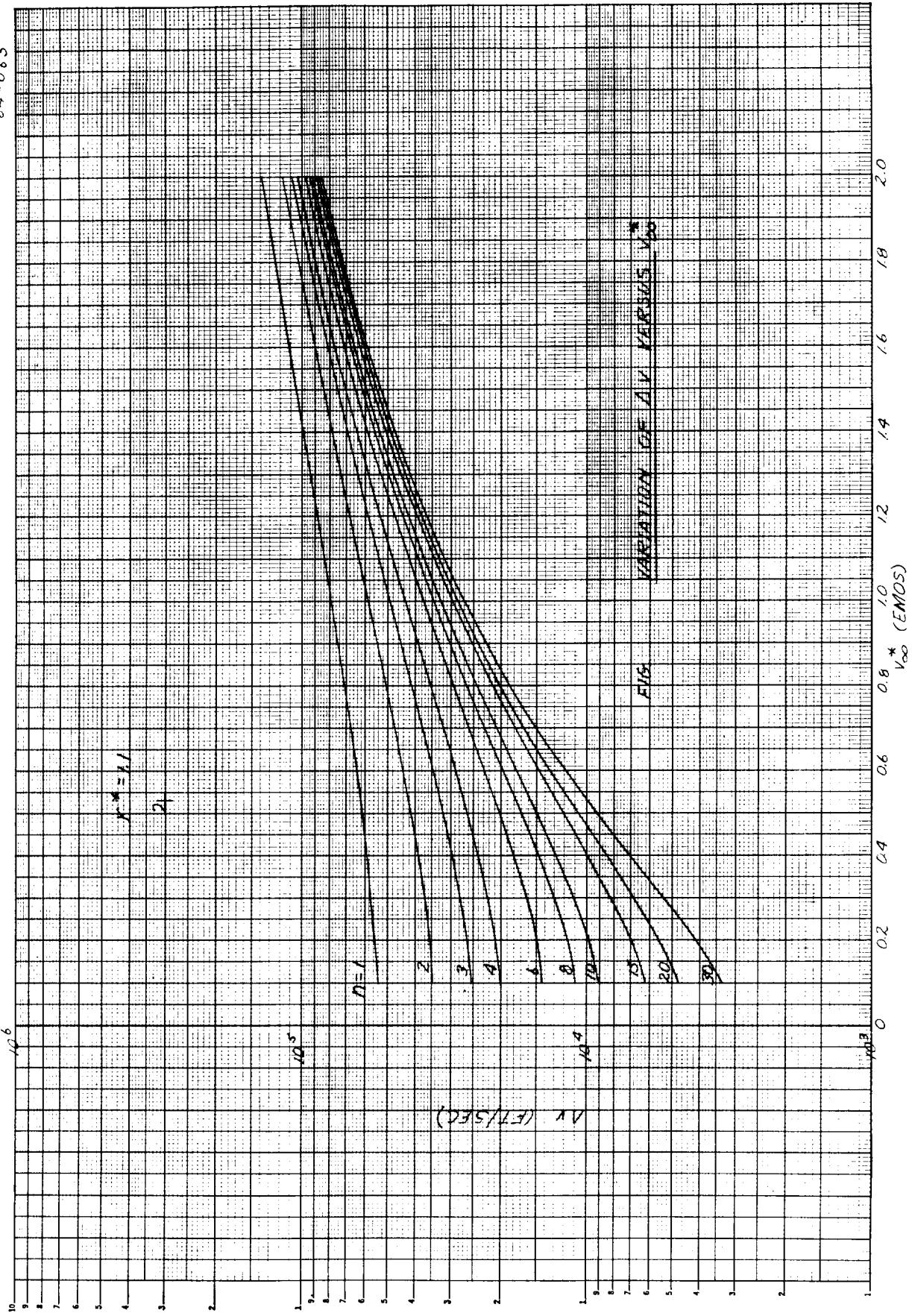




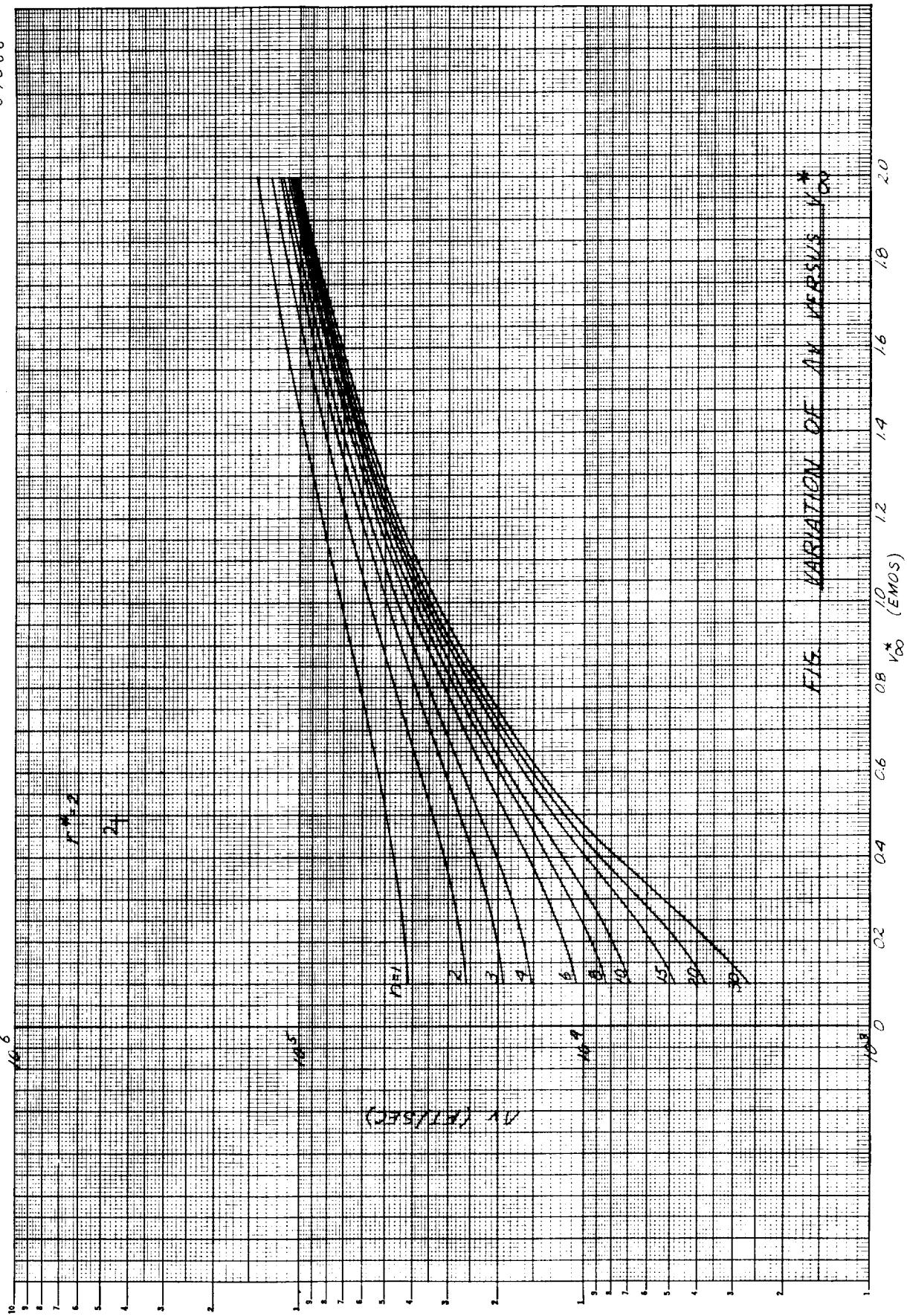


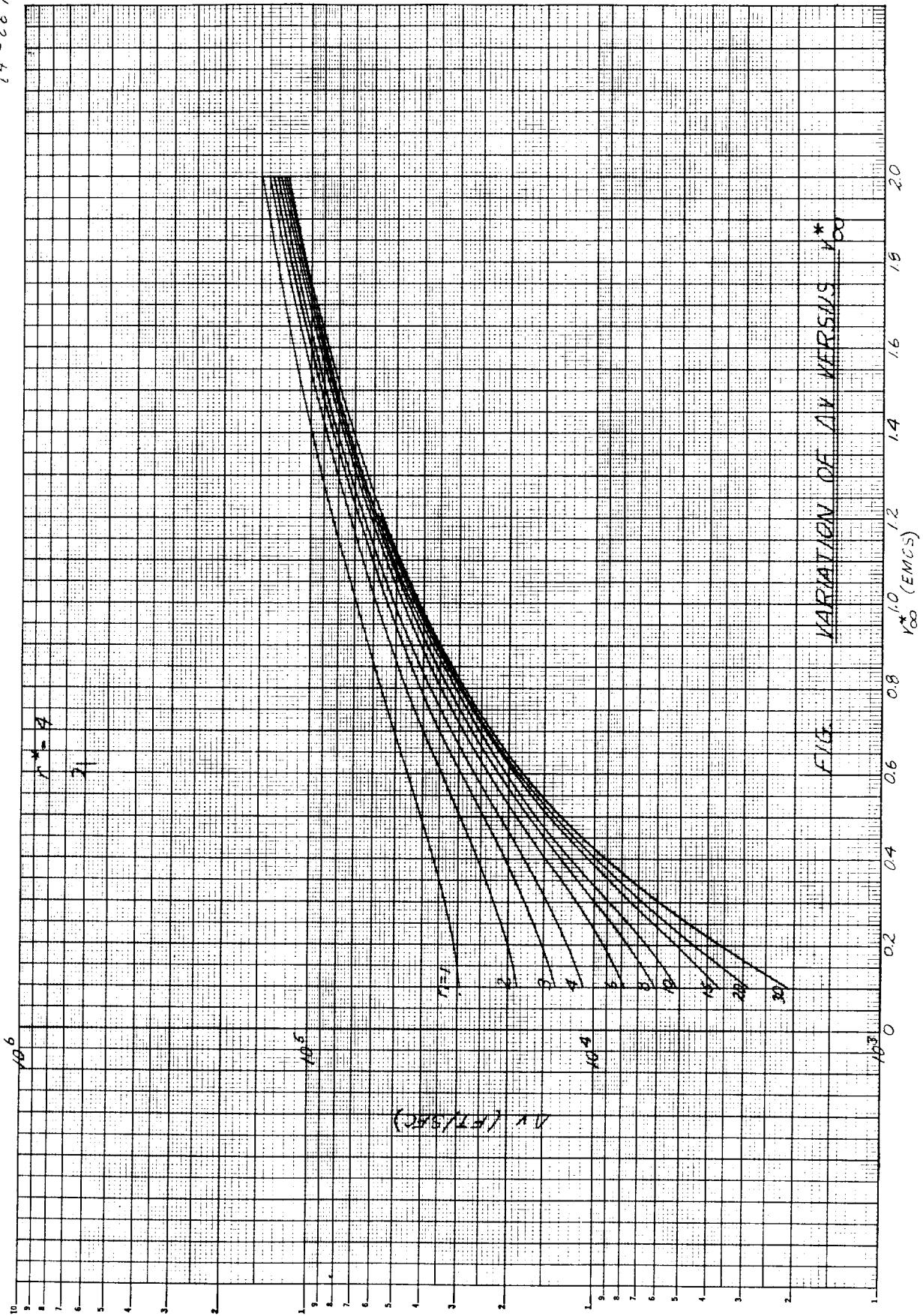


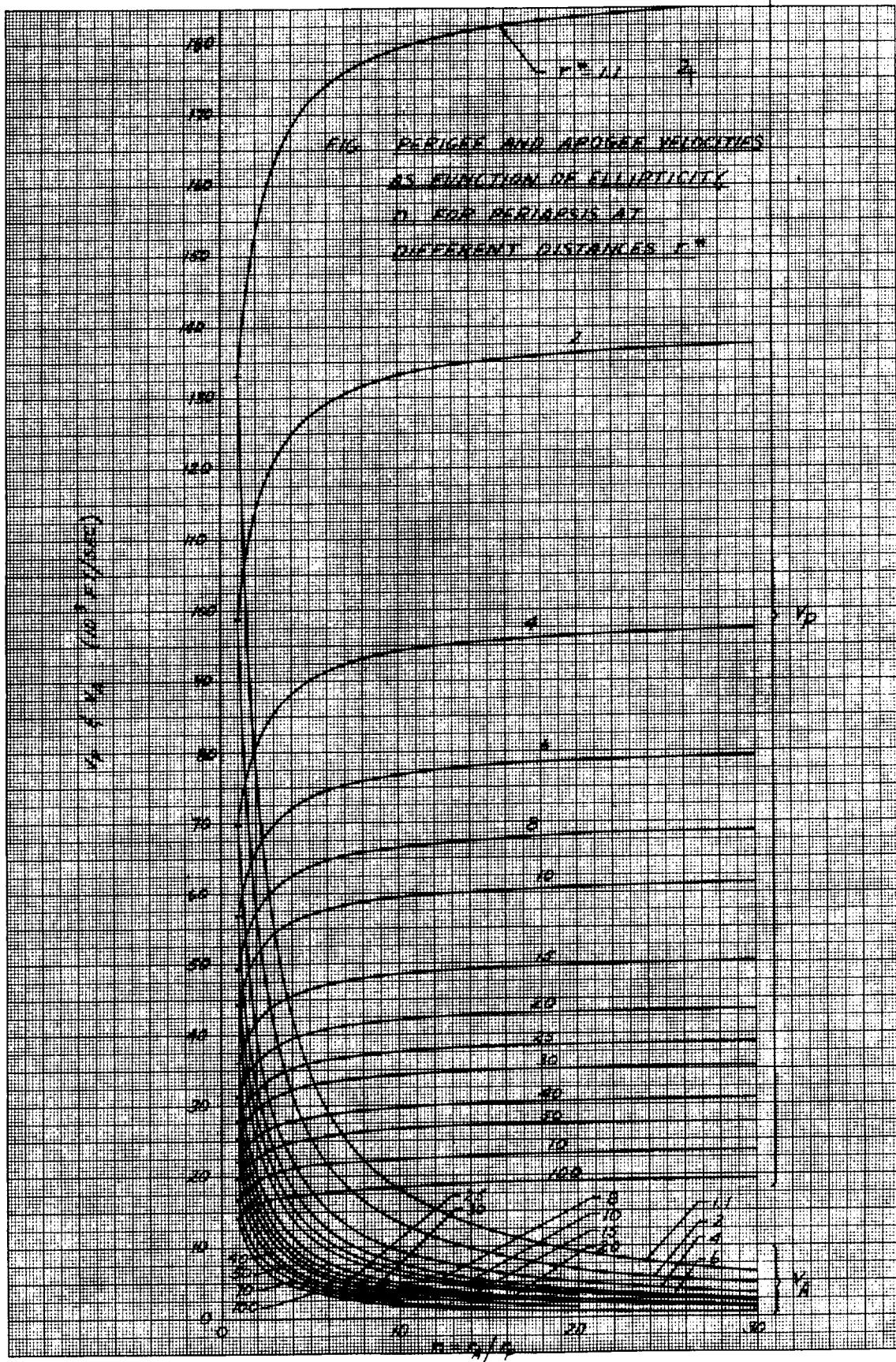
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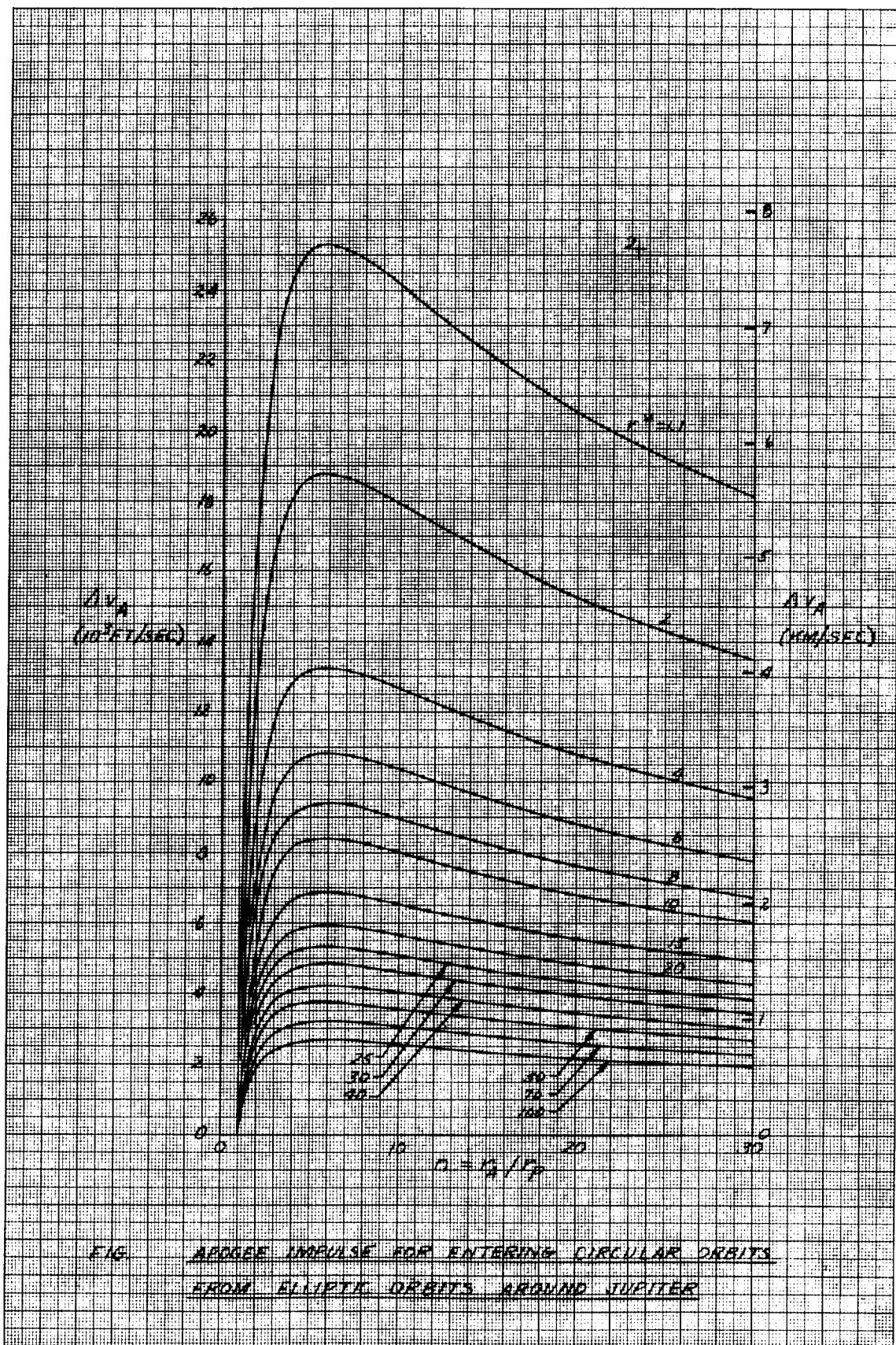


64066



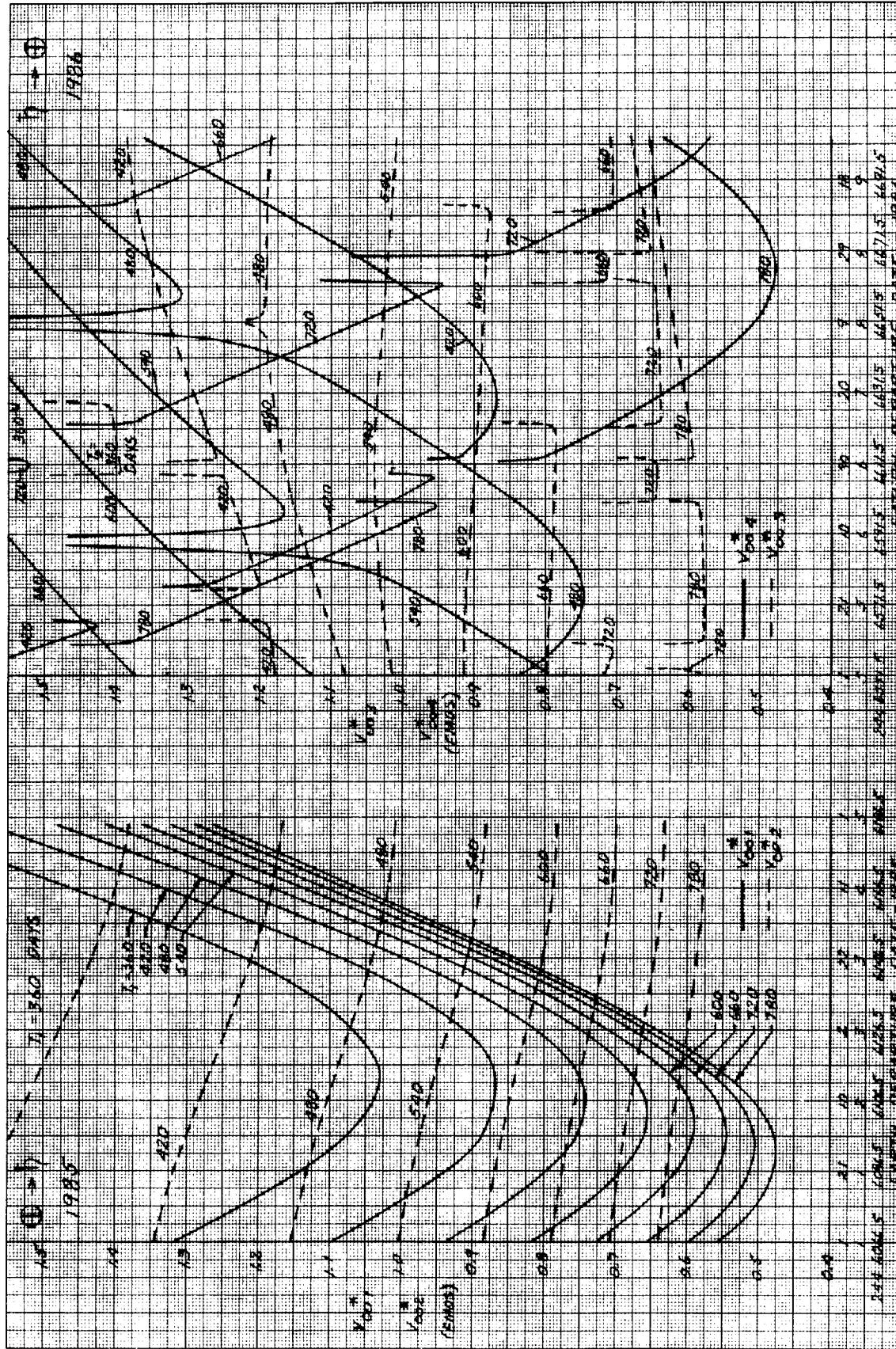


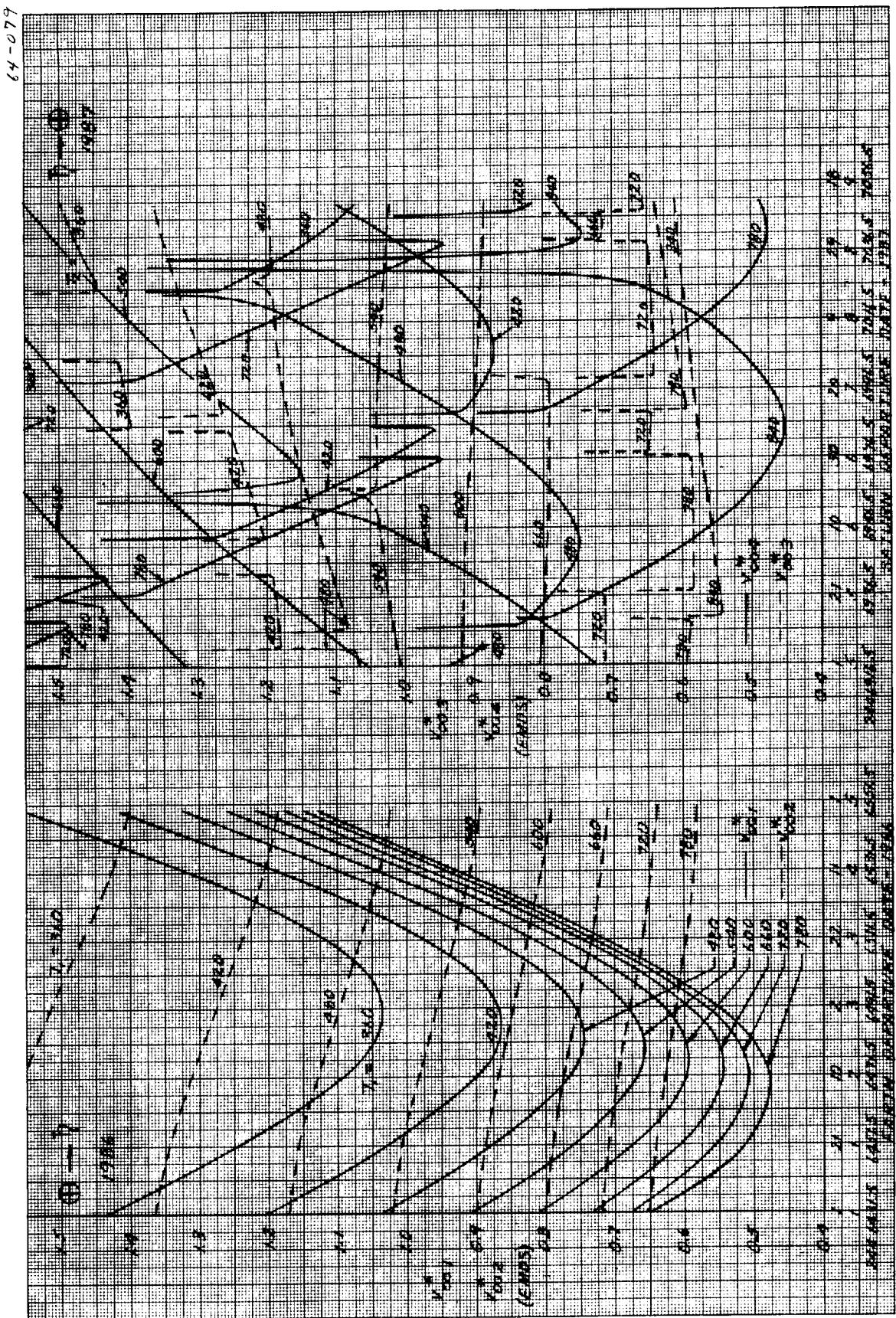


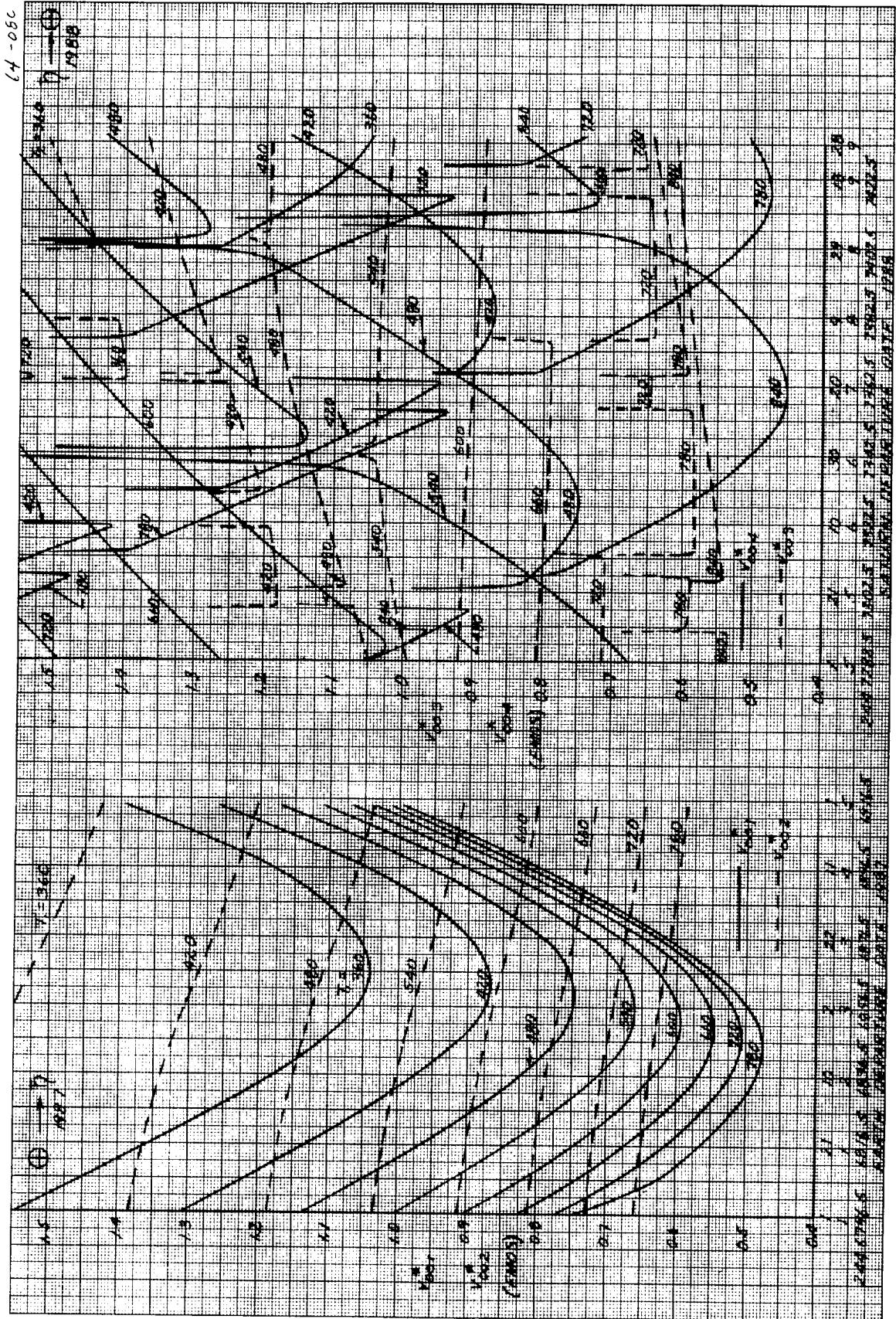


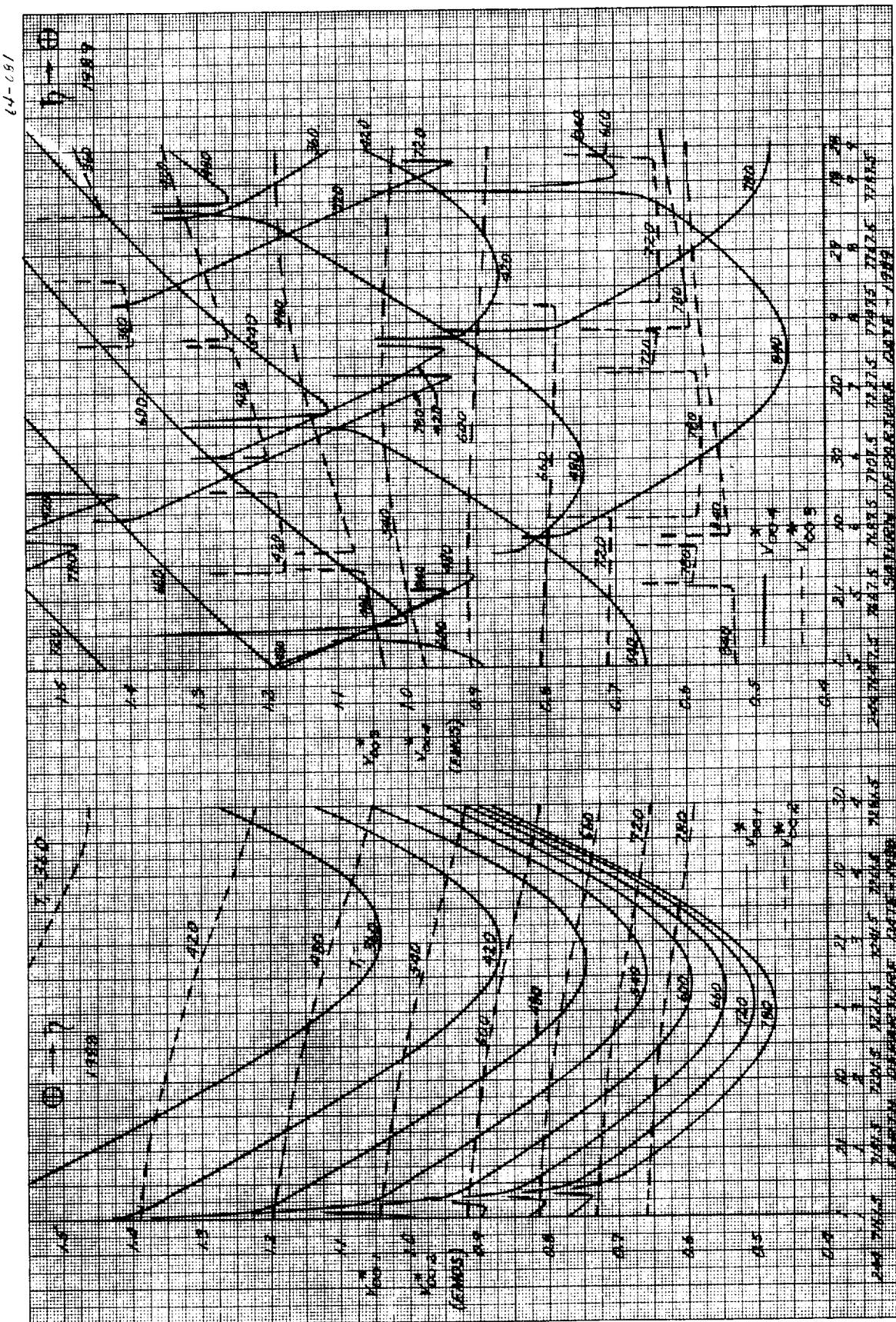
64-038

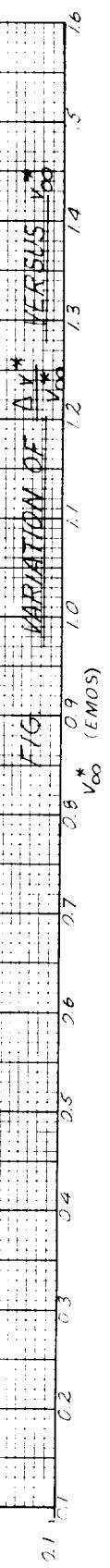
PART VI SATURN











$I_D = 0.83$

$I_D^* = I_D$

10^{-10}

10^{-9}

10^{-8}

10^{-7}

10^{-6}

$$\frac{\Delta V^*}{V_{CO}}$$

1

2

3

4

5

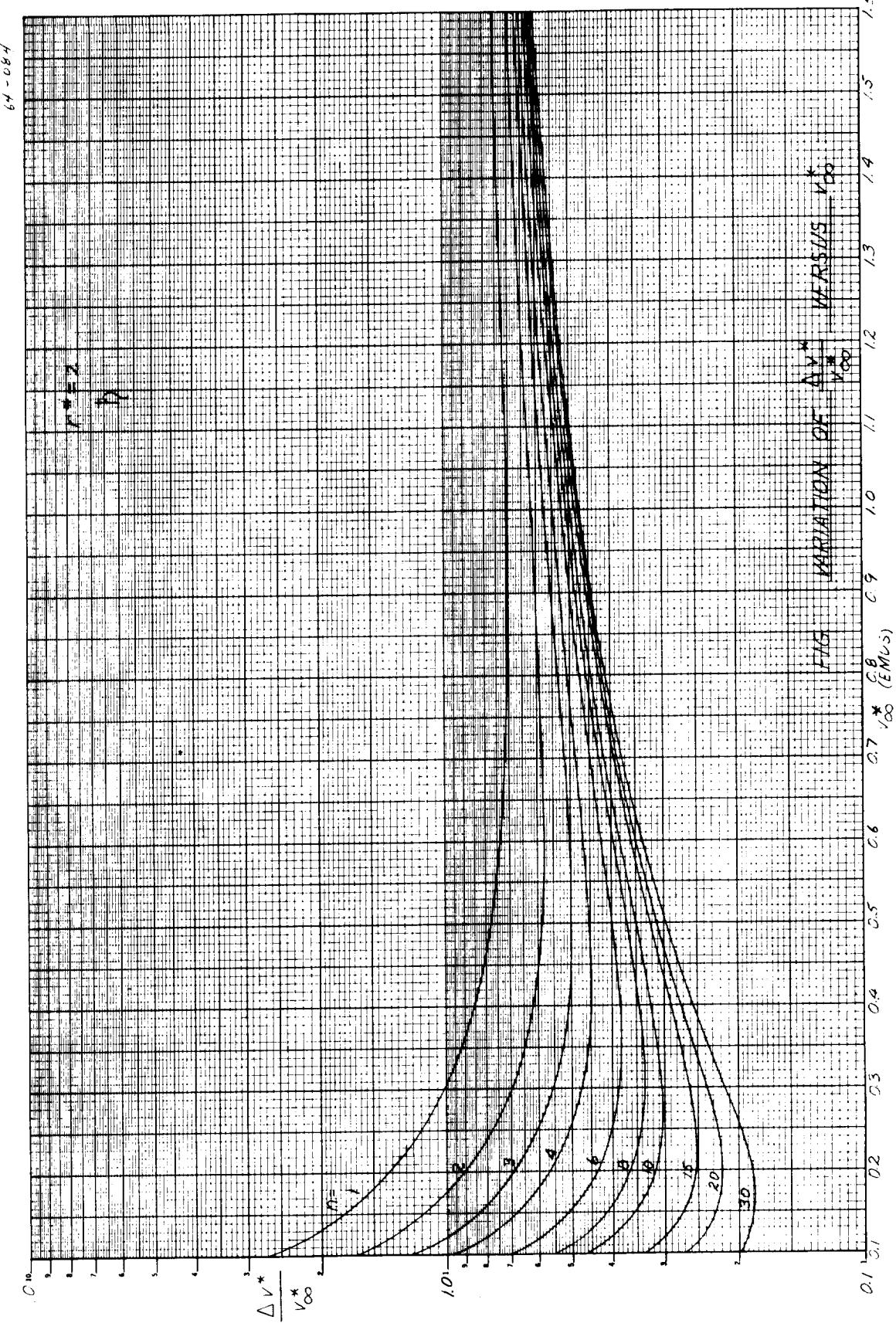
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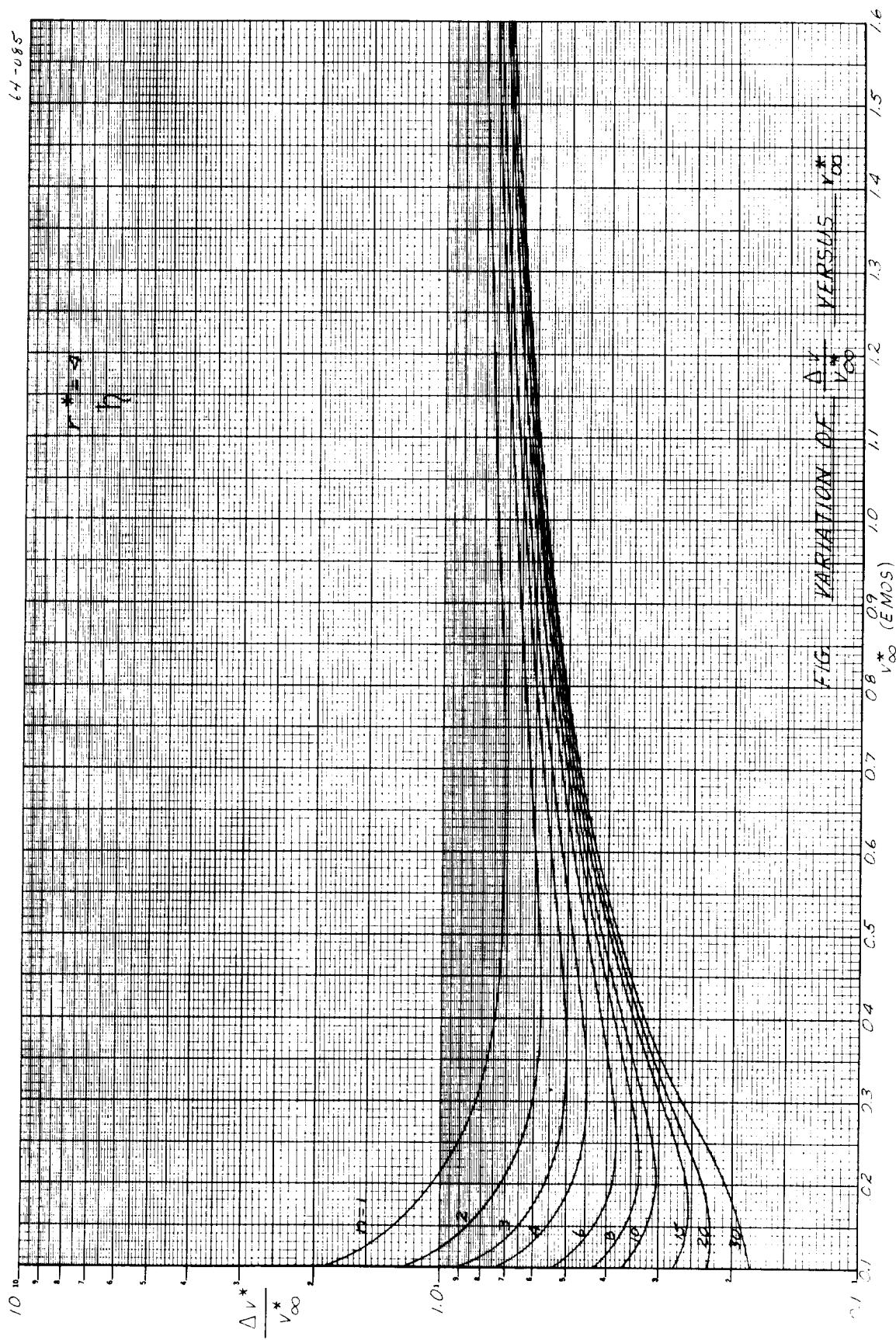
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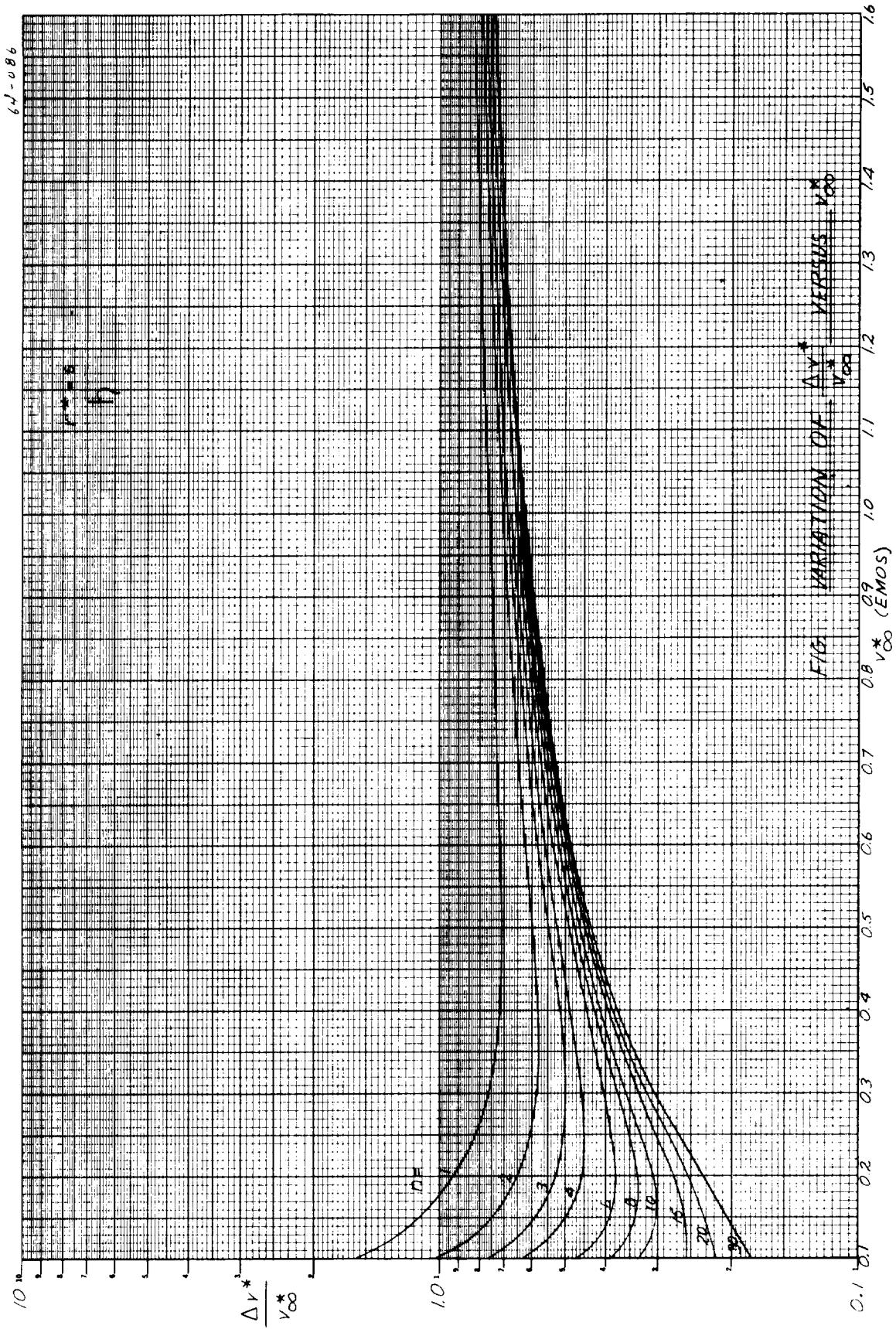
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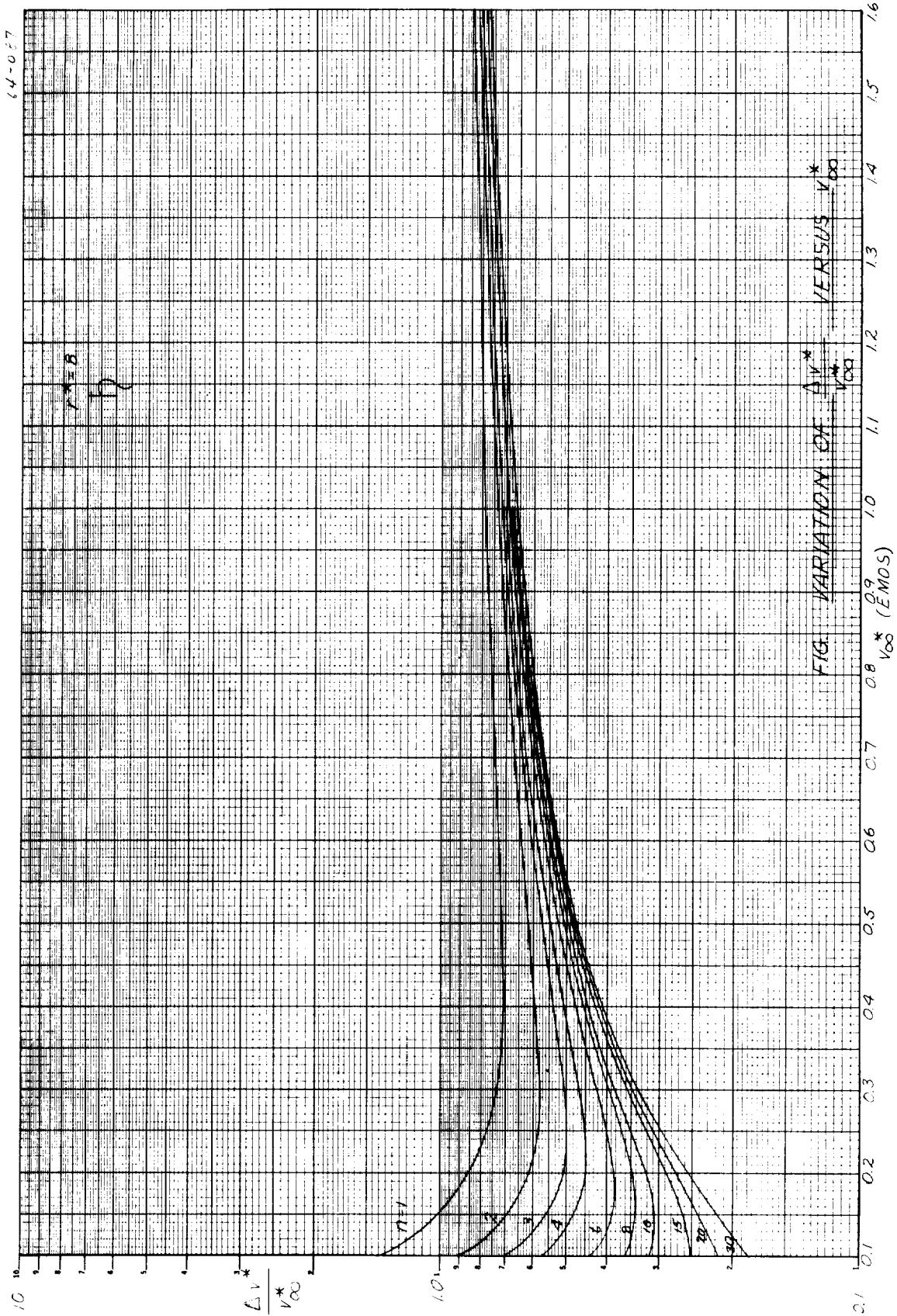
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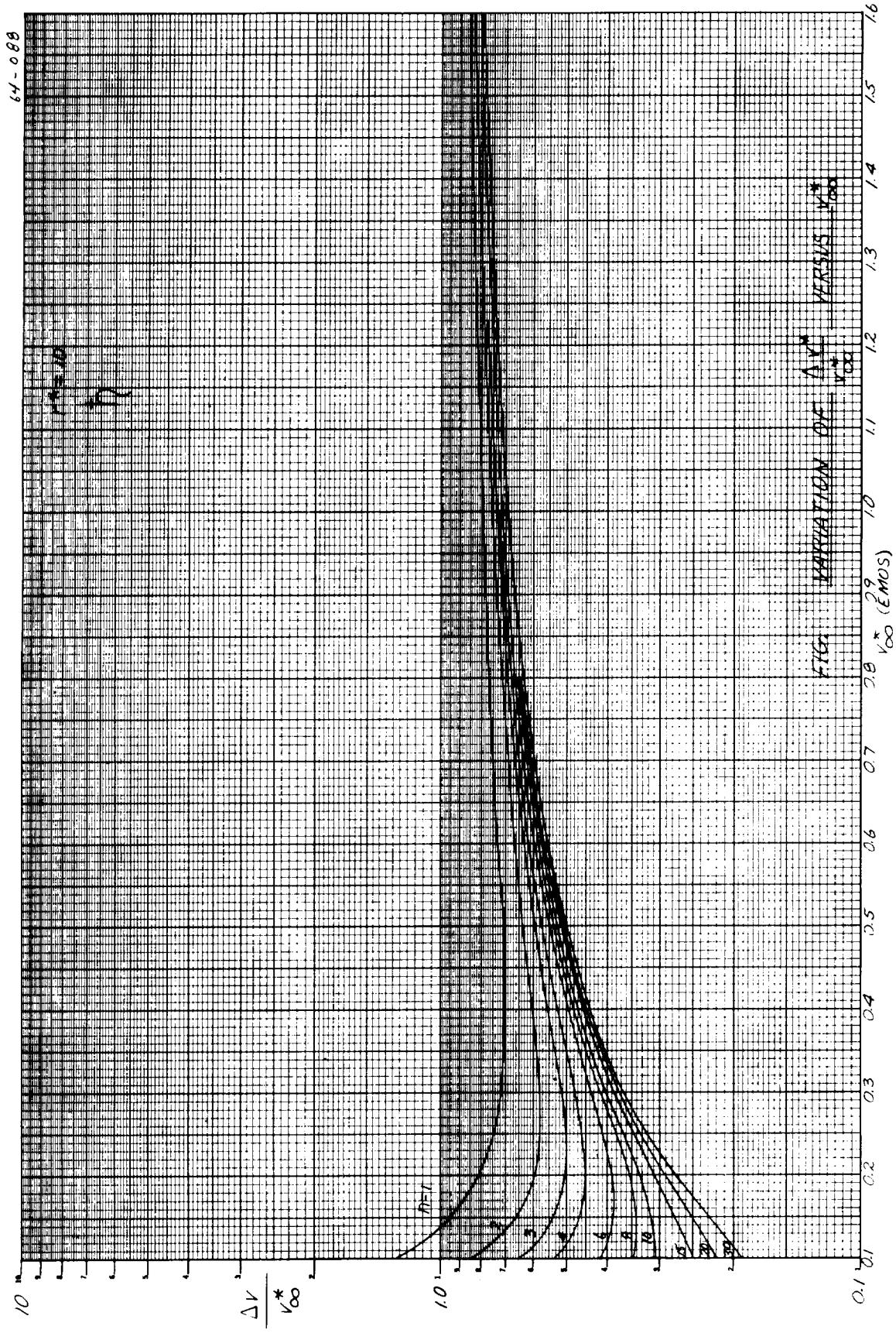
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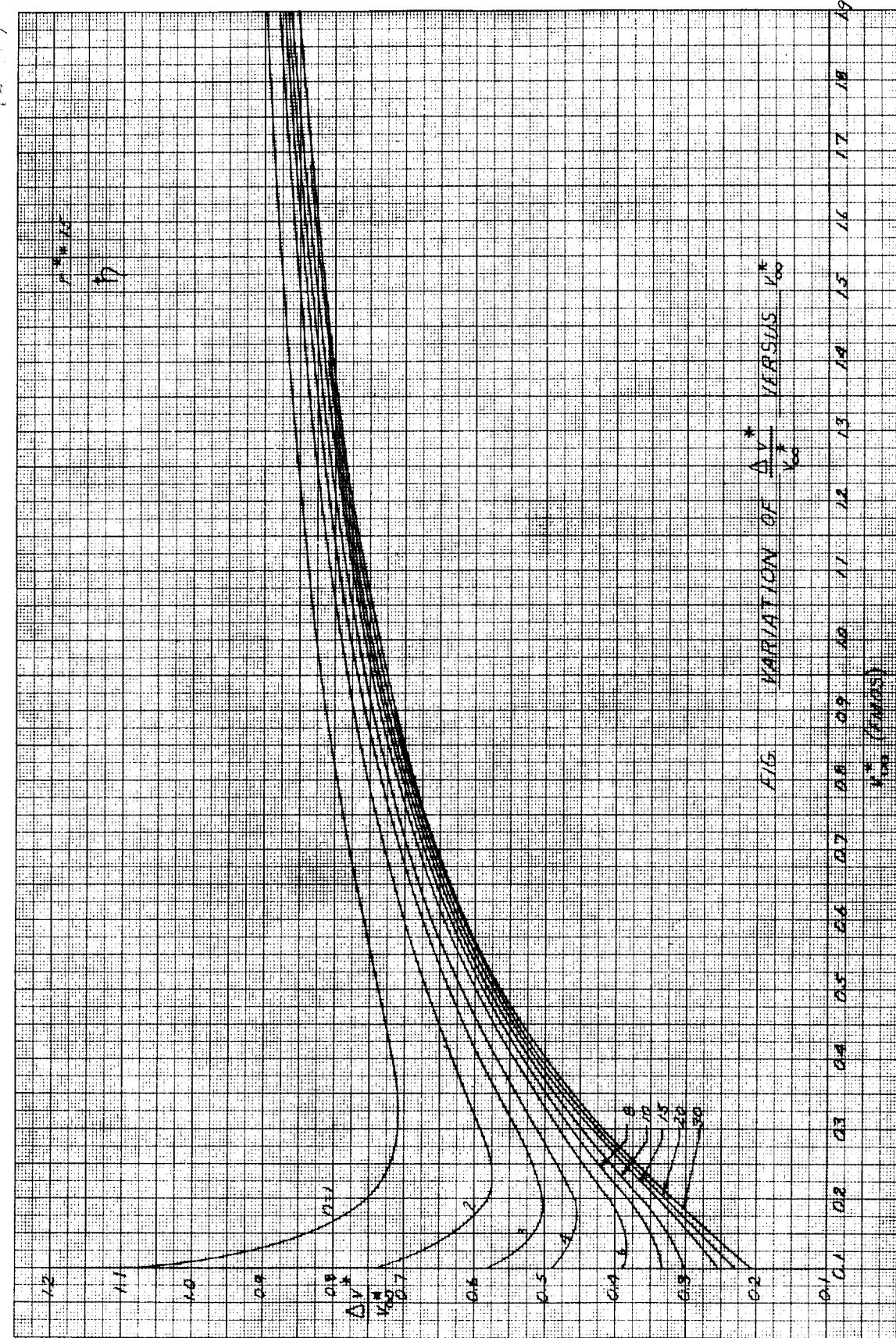


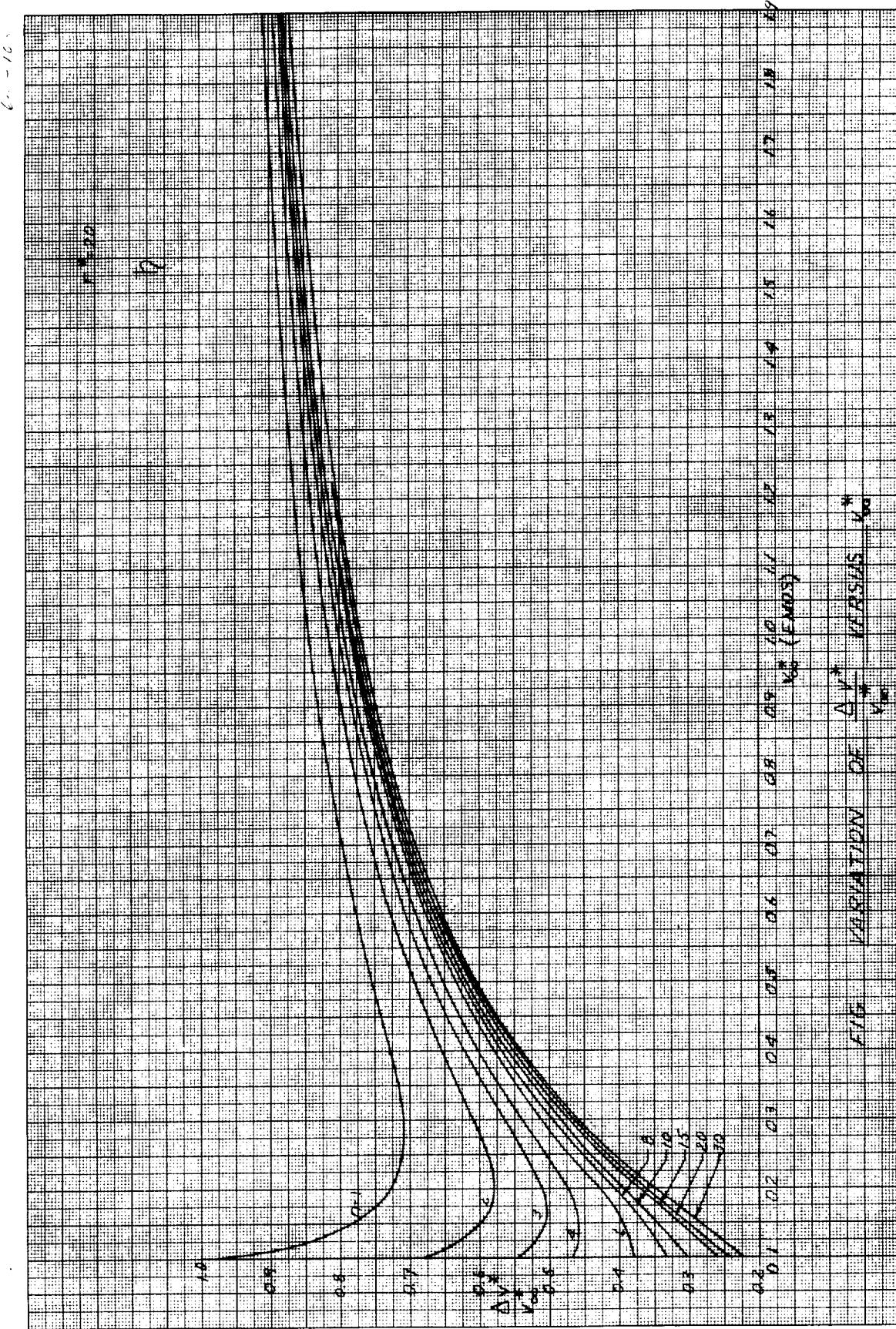


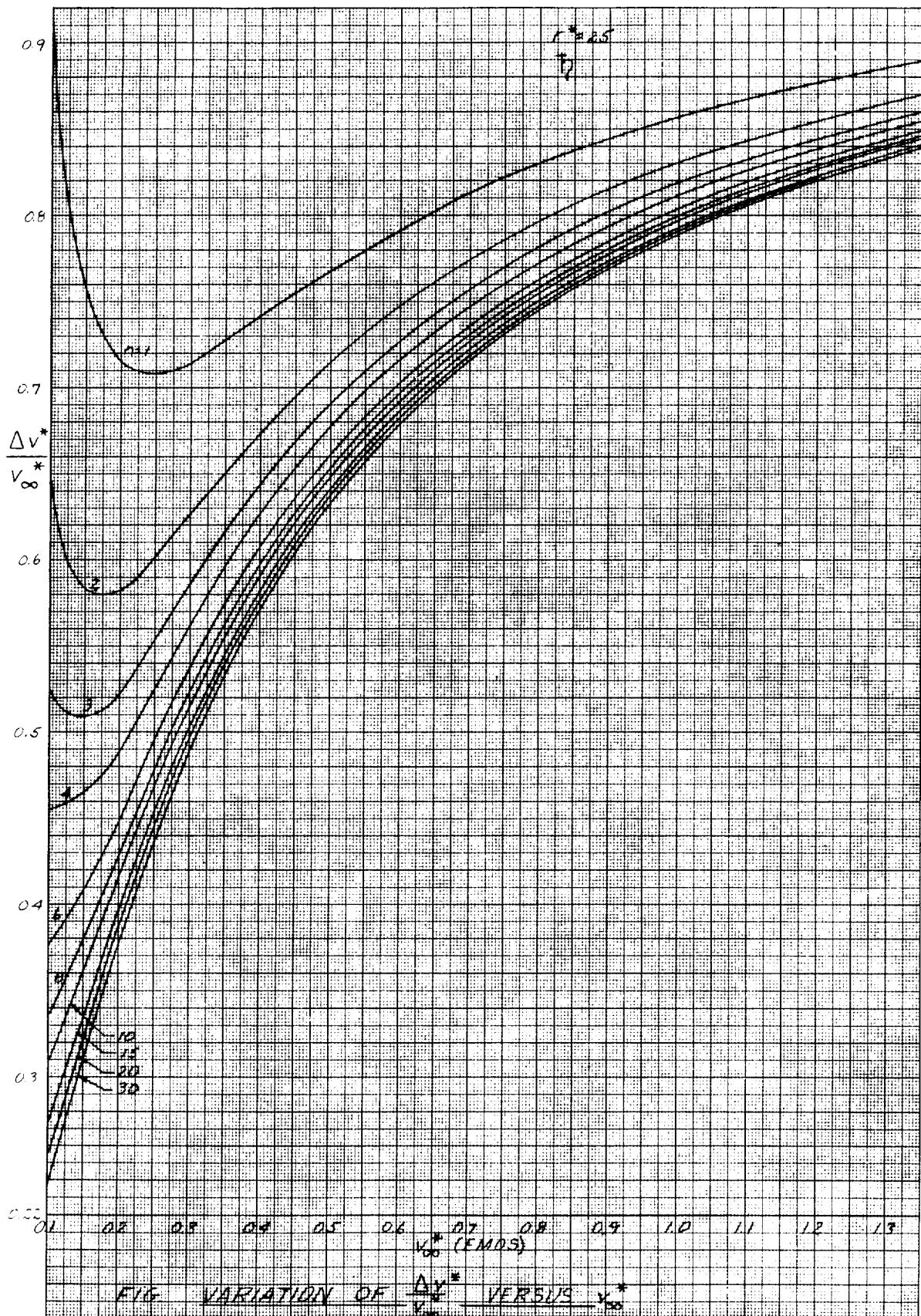


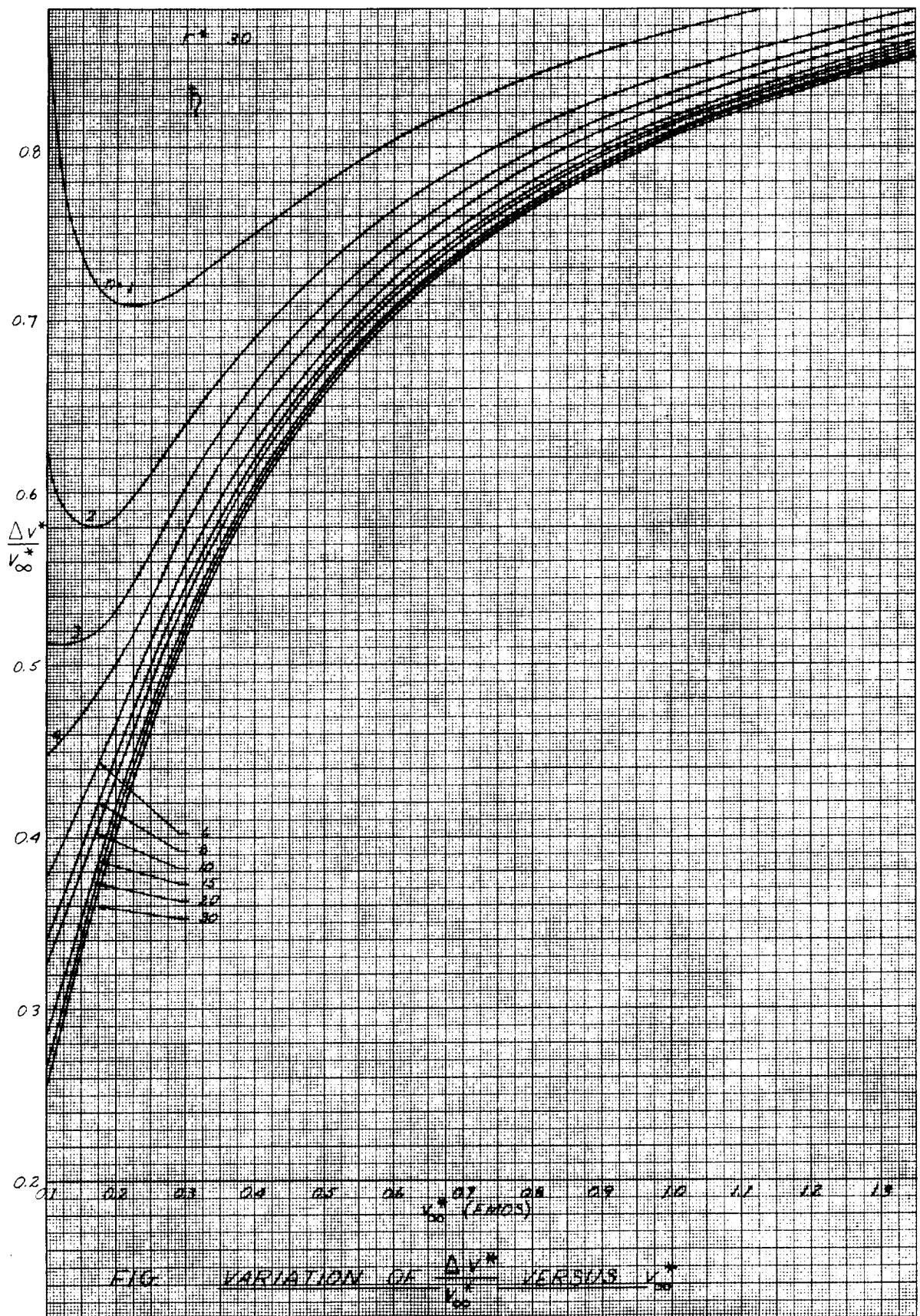


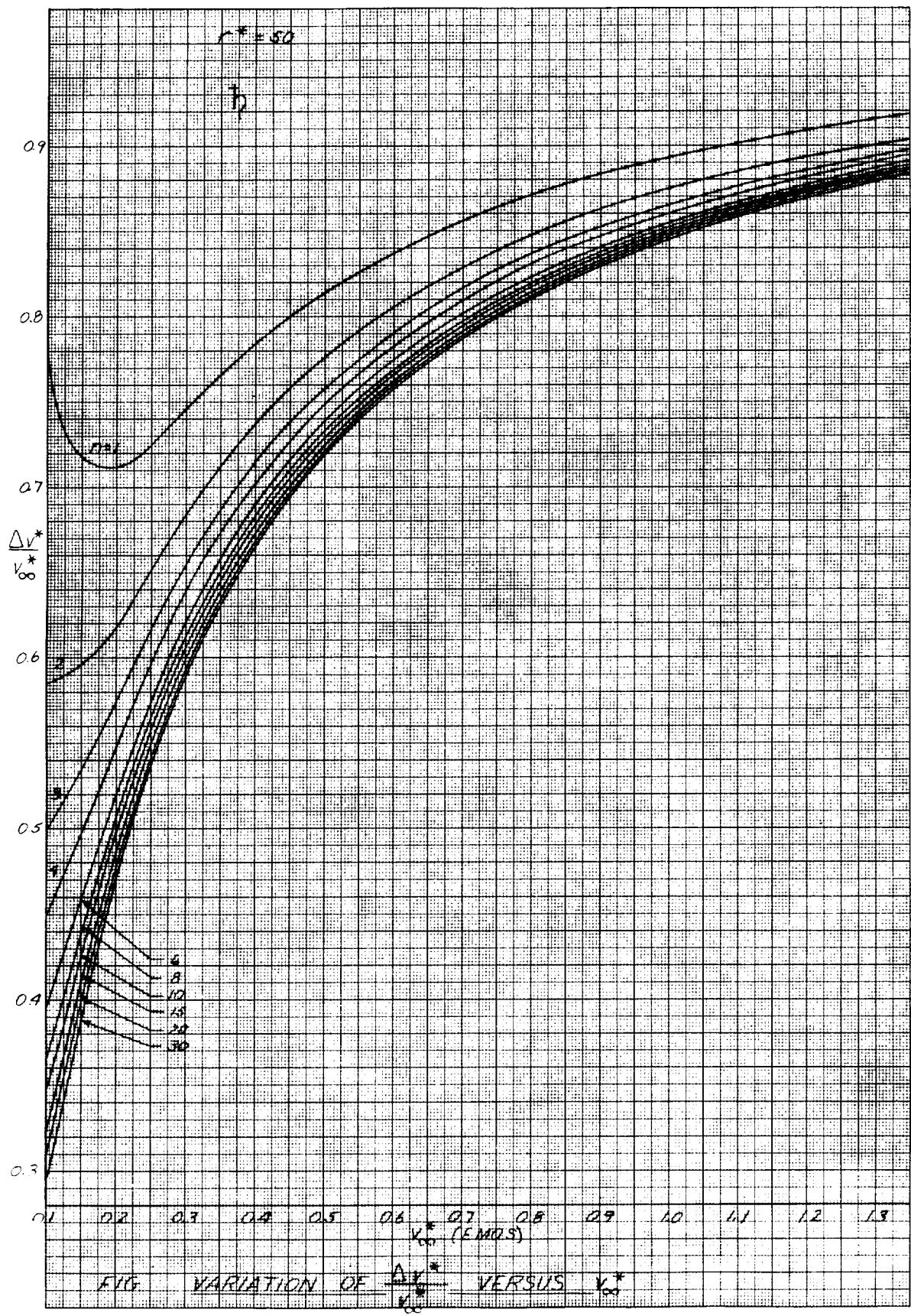


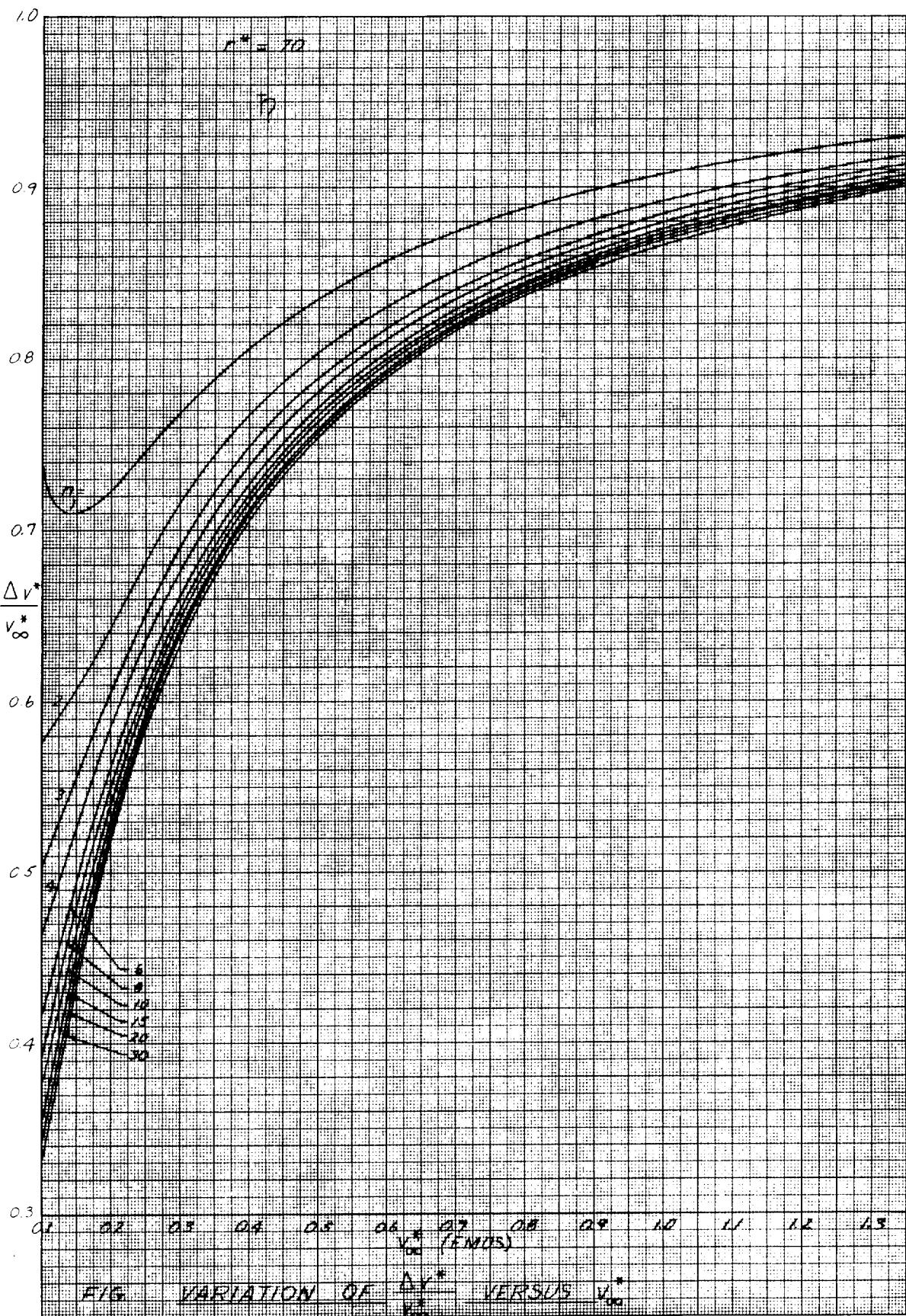


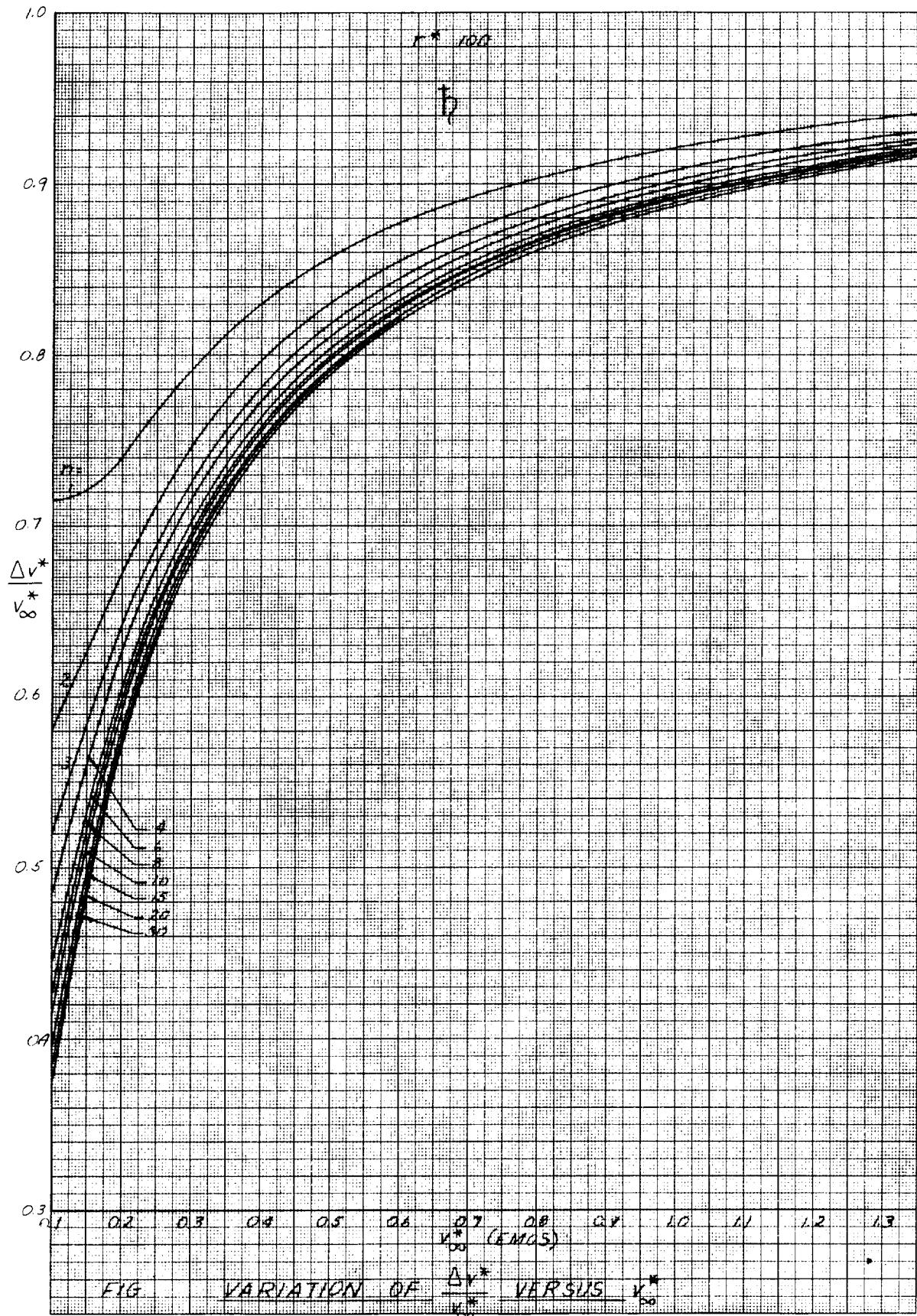


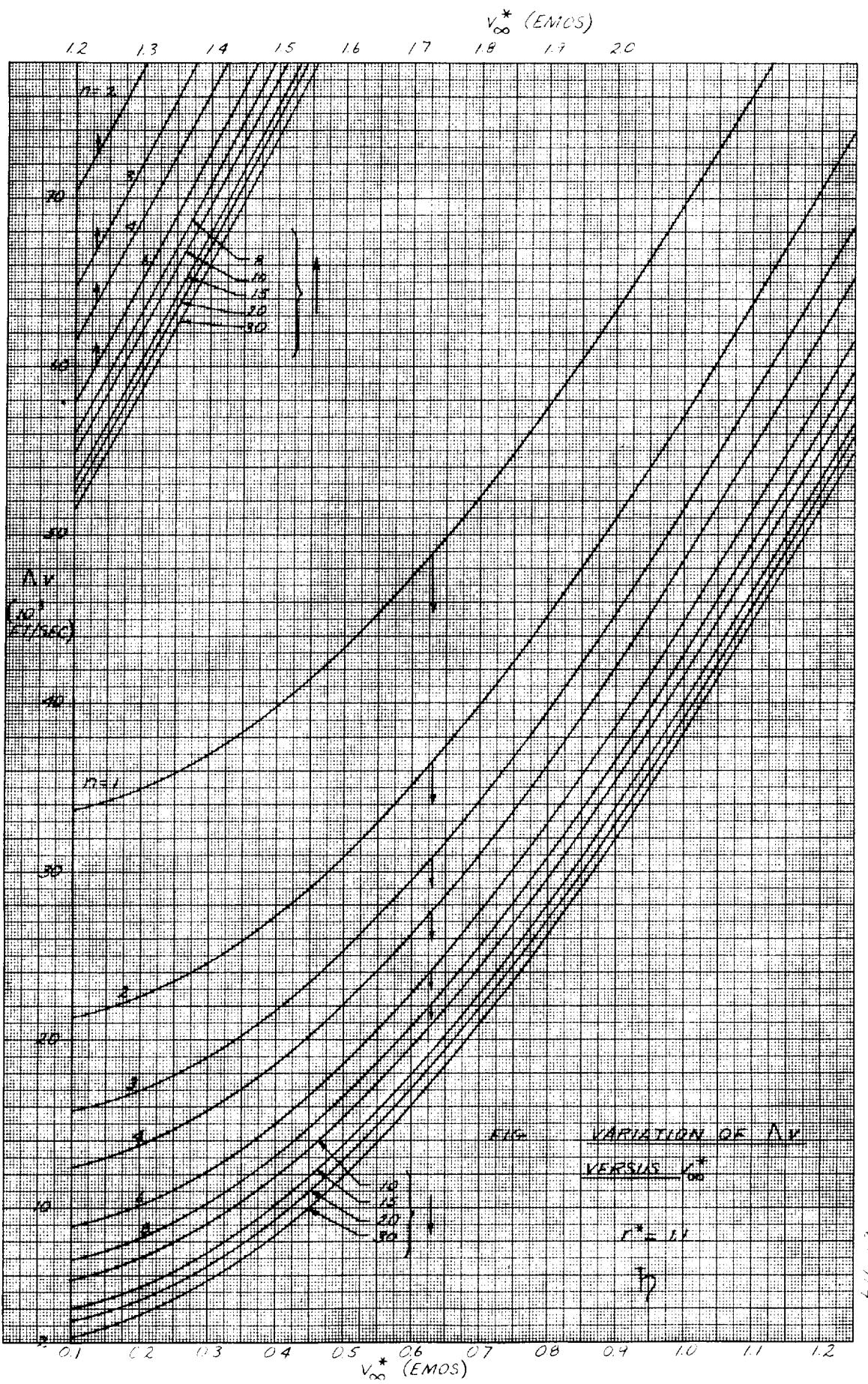


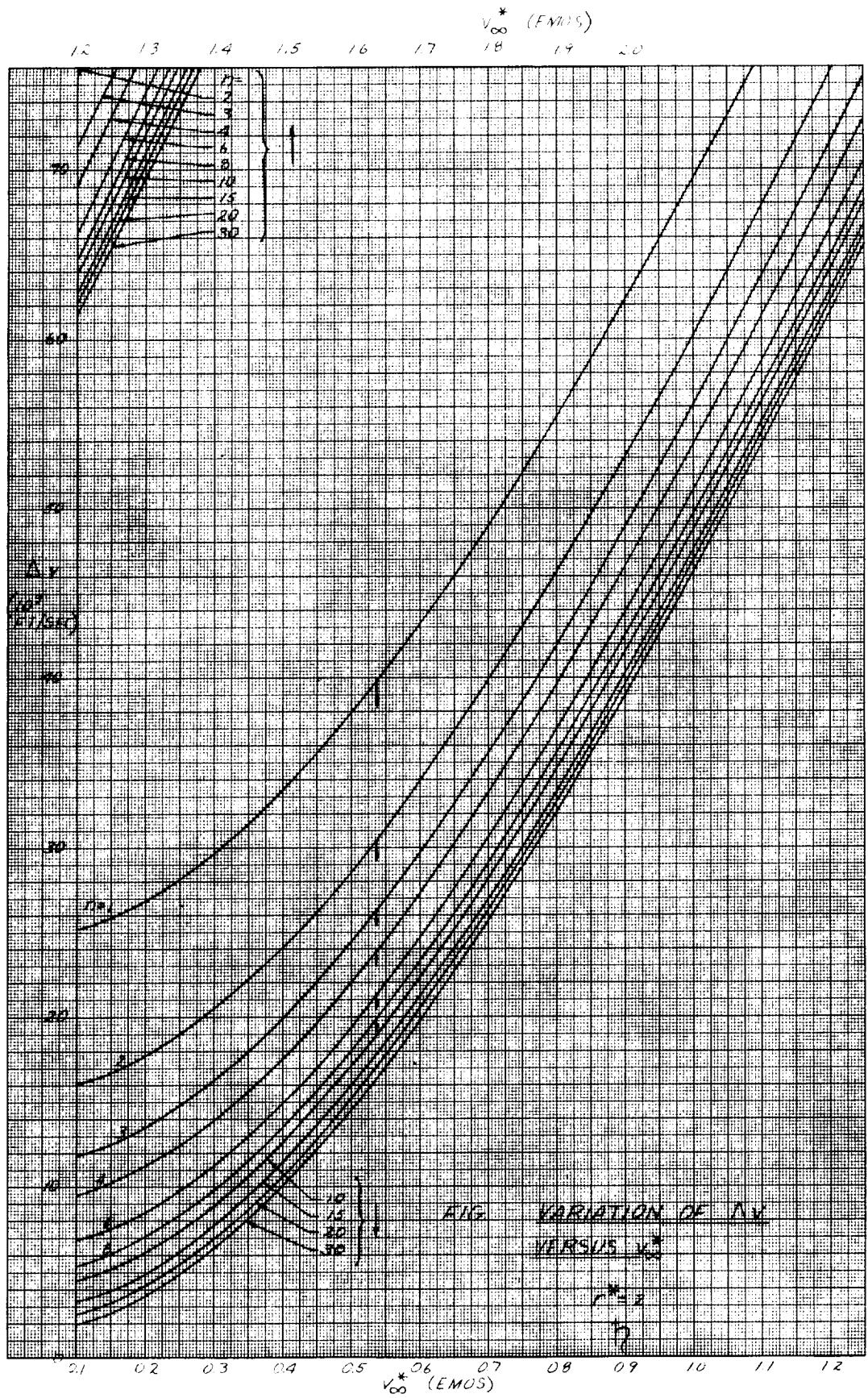


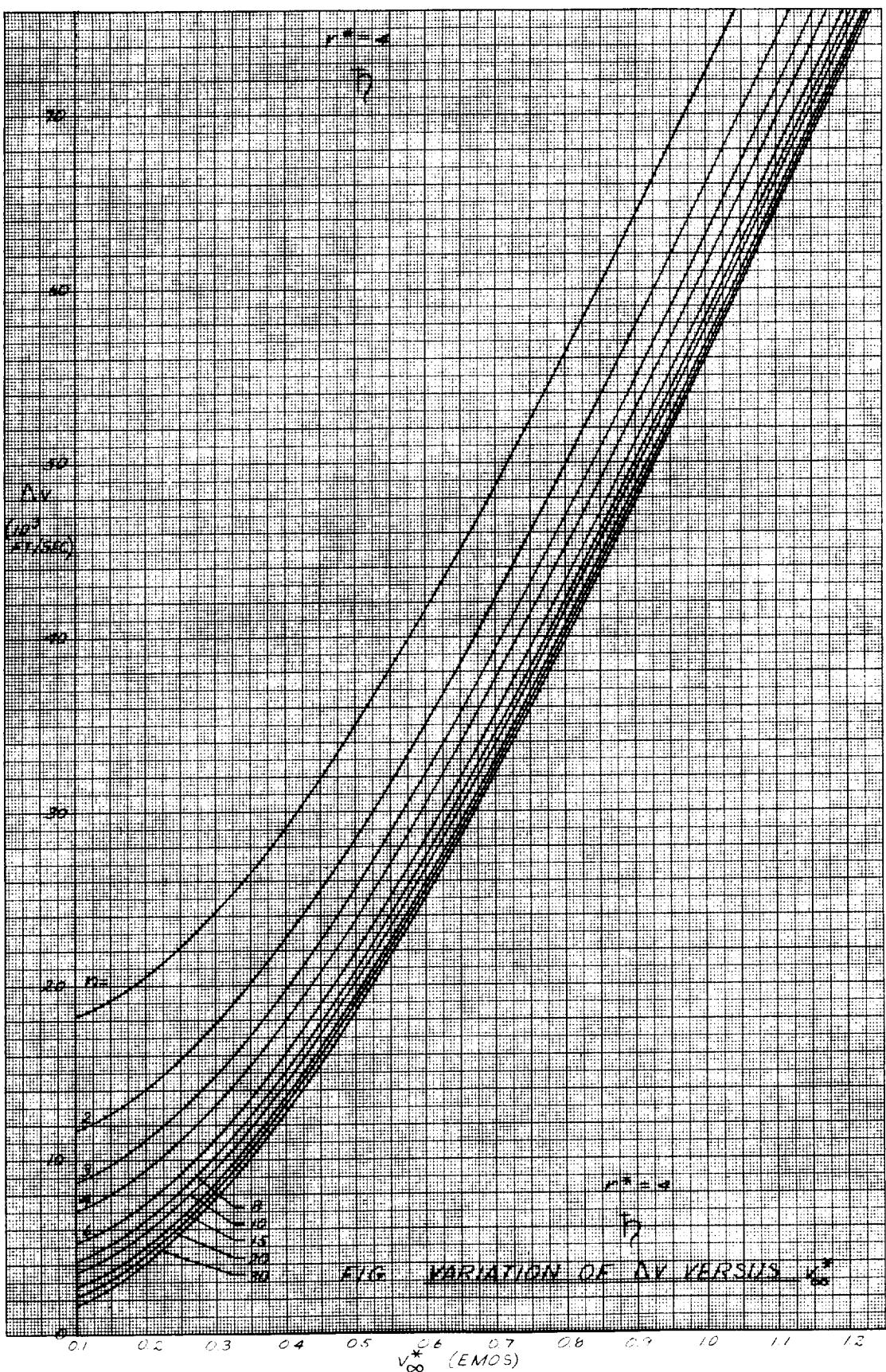


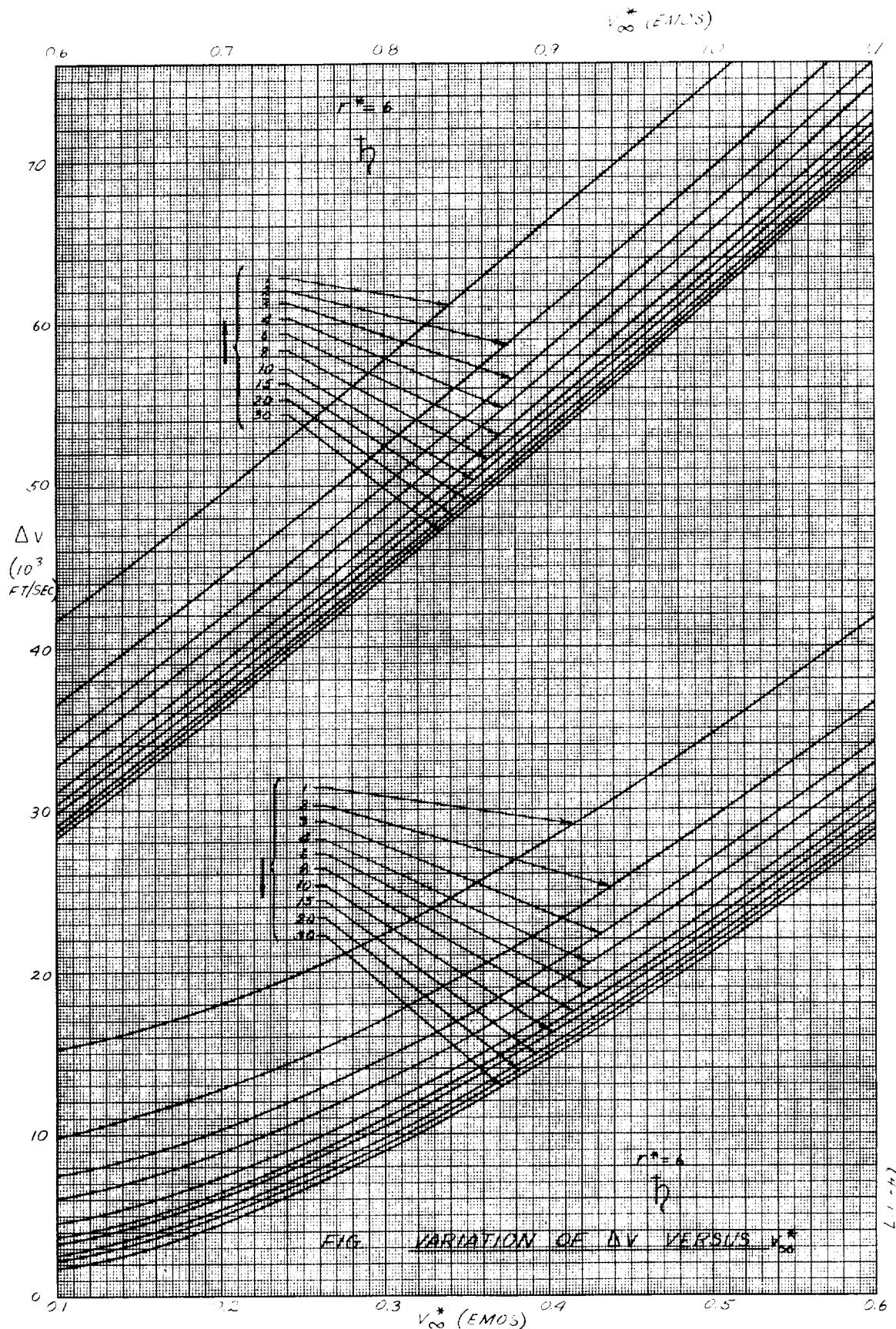


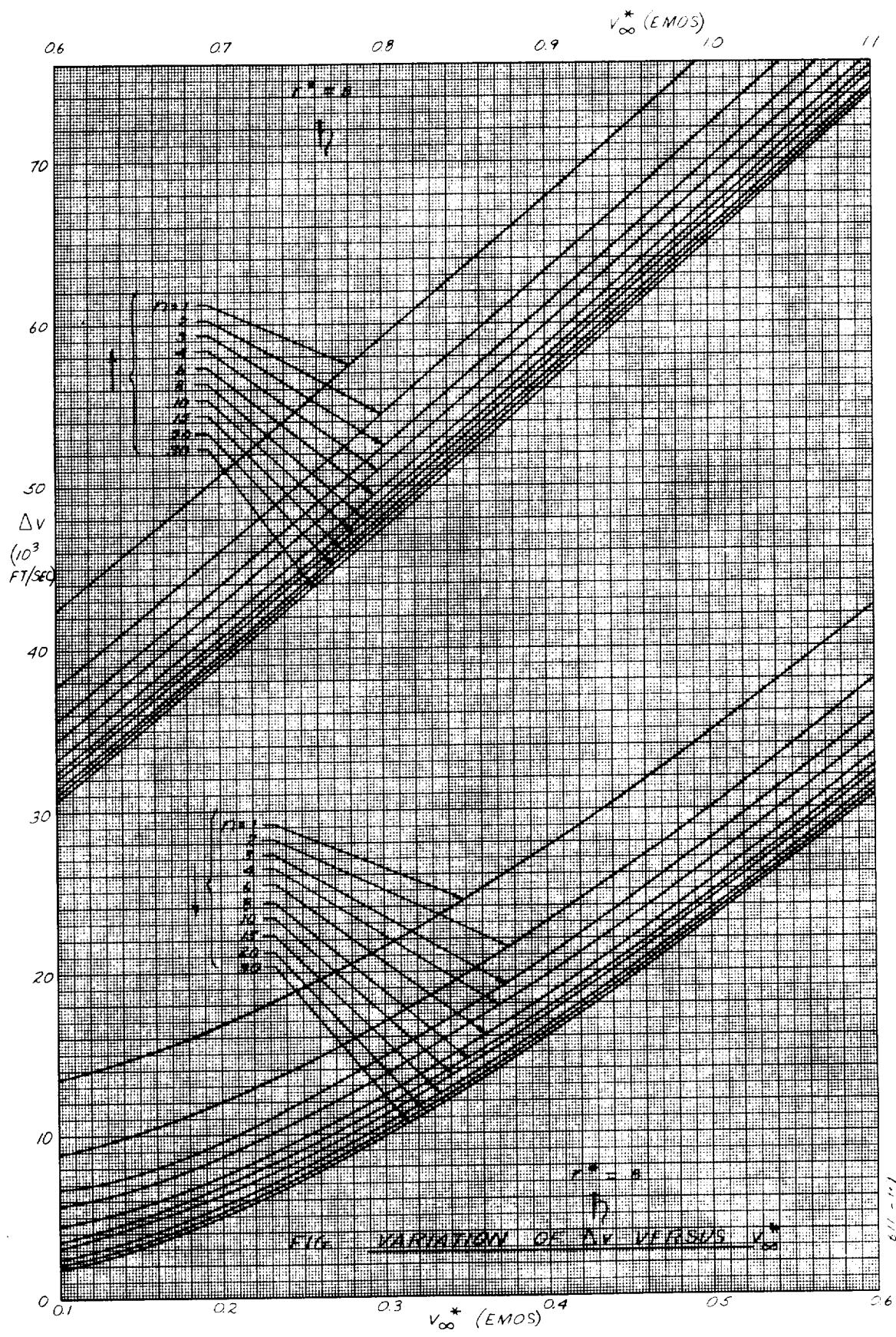












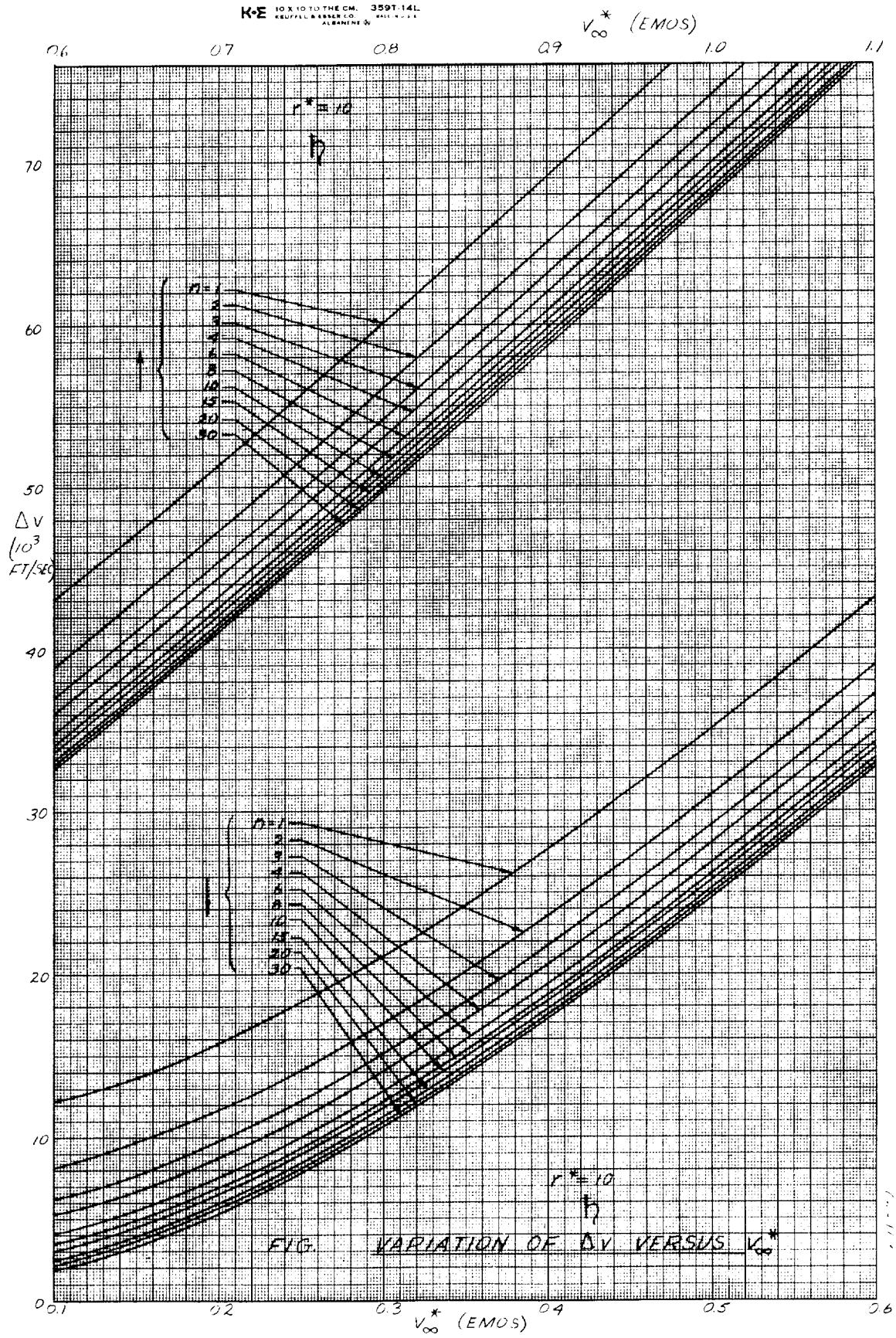


FIG. VARIATION OF ΔV VERSUS V_{∞}^*

